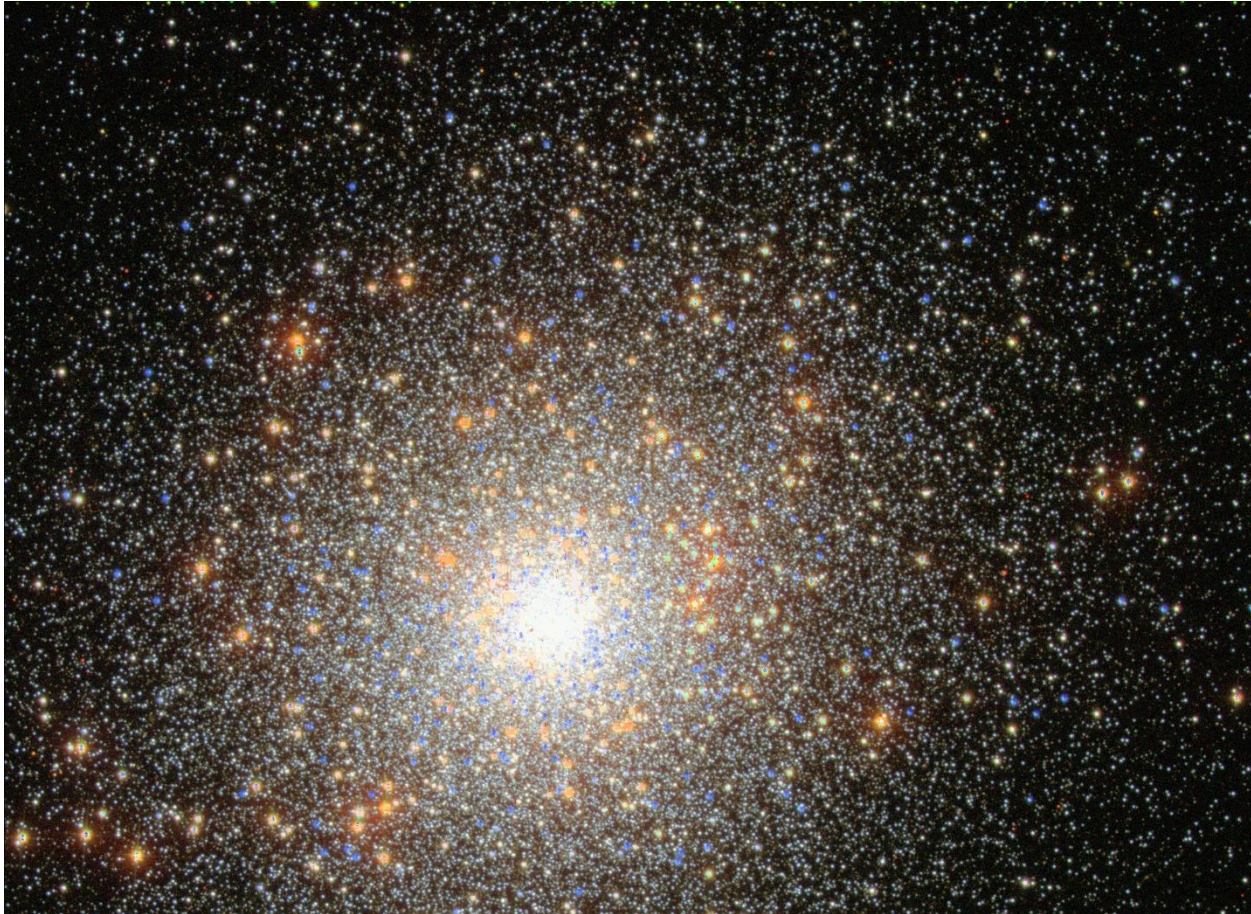


Probabilistic Cataloguing

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Harvard Astrostatistics Group



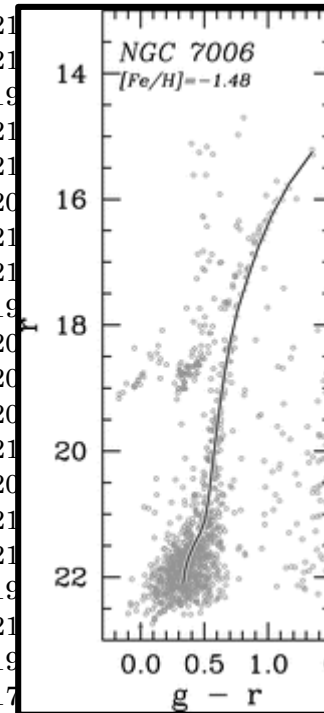
Telescopes don't make catalogues!



People make catalogues



RA (J2000)	DEC (J2000)	g	r
229.4351	2.010923	19.344	19.114
229.4353	1.990166	23.07	21.877
229.4358	2.033374	21.809	21.877
229.4361	2.070269	20.107	19.114
229.4362	1.997957	22.894	21.877
229.4364	2.048578	22.386	21.877
229.4366	2.053515	20.853	20.114
229.4369	2.103516	21.827	21.877
229.4369	2.043476	23.067	21.877
229.437	2.051732	19.96	19.114
229.4371	2.102266	20.813	20.114
229.4373	2.052342	20.785	20.114
229.4374	1.996688	21.161	20.114
229.4376	2.13321	22.476	21.877
229.4378	2.039289	20.883	20.114
229.438	2.077996	22.682	21.877
229.438	2.043483	22.884	21.877
229.4381	2.045585	20.111	19.114
229.4382	2.011463	22.069	21.877
229.4382	2.029807	19.625	19.114
229.4382	2.030182	17.835	17.877
229.4385	2.157053	22.193	21.877
229.4385	2.147021	22.492	21.546
...

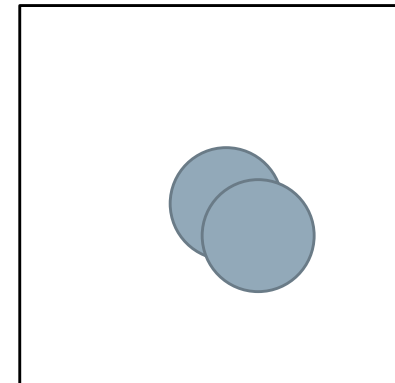
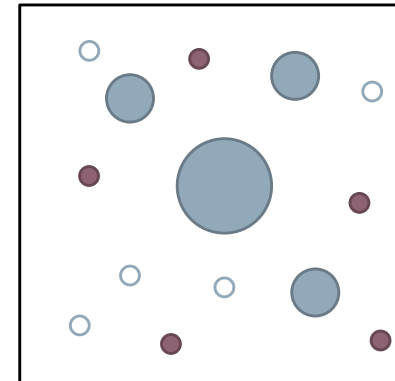


(Deterministic) Catalogues

- A (deterministic) catalogue is a list of point source candidates above some inclusion threshold TS_{incl}

$$Data, TS_{incl} \rightarrow \{\ell_i \pm \sigma_{\ell_i}, b_i \pm \sigma_{b_i}, F_i \pm \sigma_{F_i}\}_{i=1}^N$$

- **Inclusion threshold = detection threshold:**
Almost all catalogue sources are true sources
But faint true sources are not in the catalogue
- **Inclusion threshold < detection threshold:**
More faint true sources are included in the catalogue
But many catalogue sources are not true sources
The data is overfitted
- Overlapping point sources may not be deblended

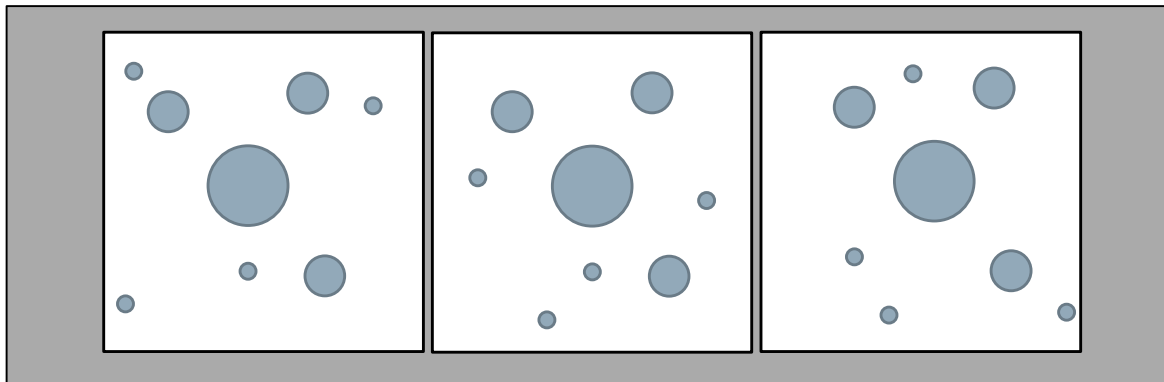


Probabilistic Catalogues

- A probabilistic catalogue is a posterior probability distribution over the space of lists of point source candidates

$$\begin{aligned} &P(\{\ell_i, \mathcal{L}_i, F_i\}_{i=1}^N | Data) \\ &= \pi(\{\ell_i, \mathcal{L}_i, F_i\}_{i=1}^N) \mathcal{L}(Data | \{\ell_i, \mathcal{L}_i, F_i\}_{i=1}^N) \end{aligned}$$

- Sampling the probabilistic catalogue provides an *ensemble of catalogues* inferred from the data



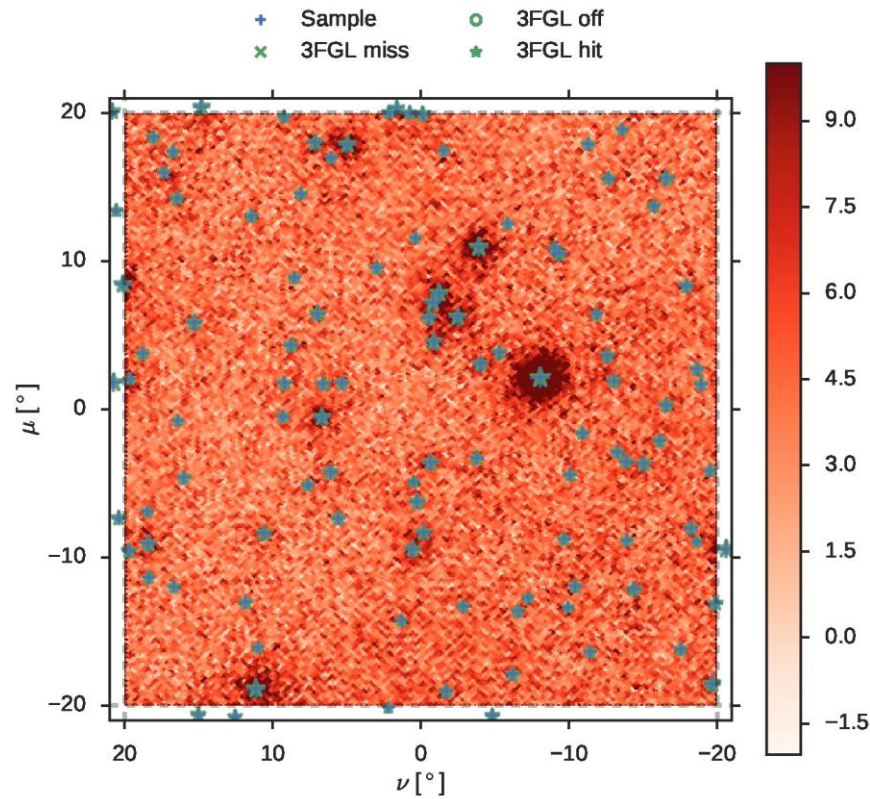
Why Probabilistic Catalogues?

- The reality of a single faint point source candidate will be very uncertain, but the properties of a faint population are constrained
- The uncertainty in deblending sources with overlapping PSFs can be captured
- Provides a framework to marginalize over uncertainties (modelling, instrumental, calibration, etc.)
- Probabilistic cataloguing more fully captures the information contained in the data and the *inherent degeneracies* of point source identification

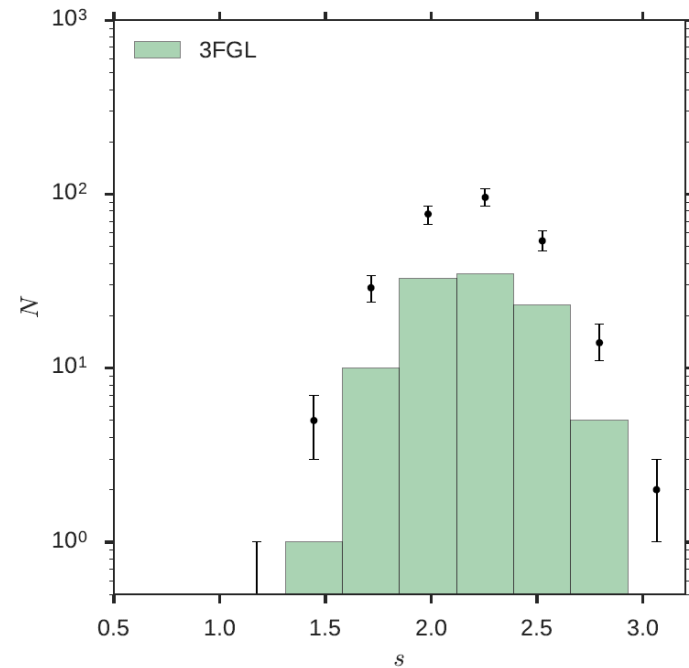
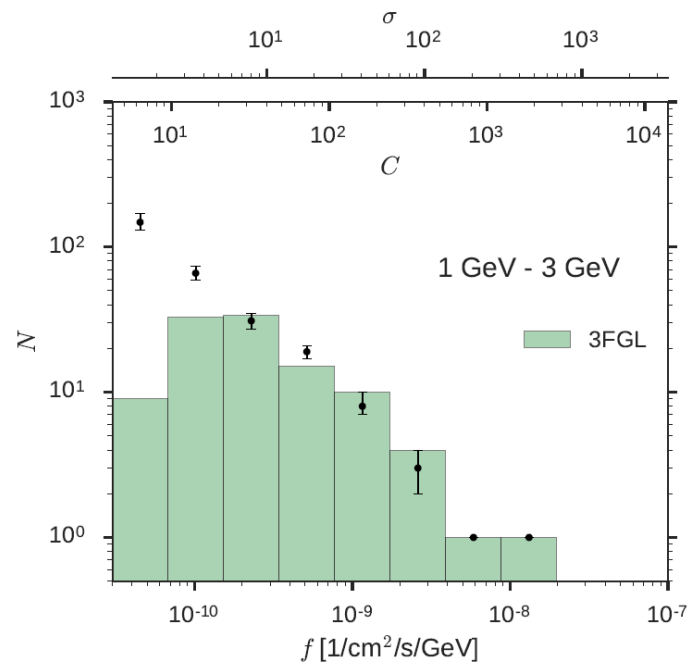
Application I: *Fermi* High Latitude

- North Galactic Pole $20^\circ \times 20^\circ$ ($N_{pix} = 29\,880$)
- 3 energy bins: 0.3-1 GeV, 1-3 GeV, 3-10 GeV
- Region includes 108 3FGL sources
- Run with ~ 250 CPU-hours
- Diffuse sources:
 - Galactic diffuse emission
 - Isotropic emission
- Point source population:
 - Mostly distant active galaxies
 - Assumed to be isotropically distributed
 - Unknown flux distribution parameterized as power law

Catalogue Samples



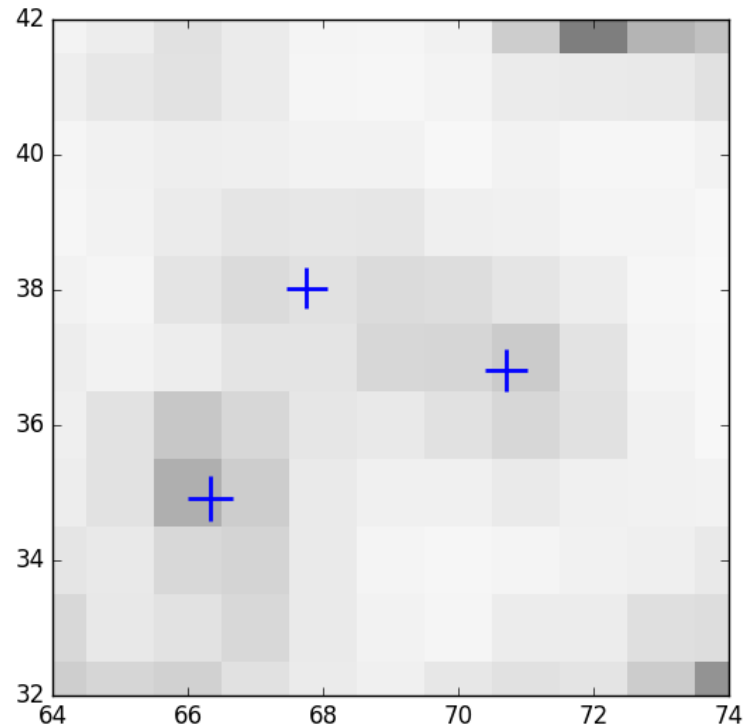
Flux and Colour Distributions



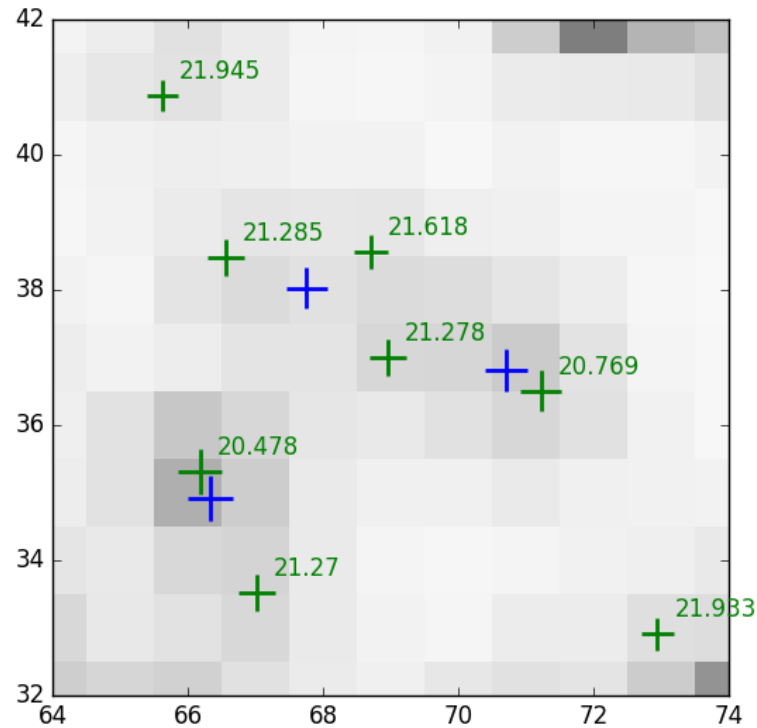
Application II: SDSS Globular Cluster

- Messier 2 $40'' \times 40''$ ($N_{pix} = 10\,000$)
- Region includes 337 DAOPhot sources
- Run with ~ 250 CPU-hours
- Region has also been observed with HST, which has better angular resolution, identifying 1 000 sources

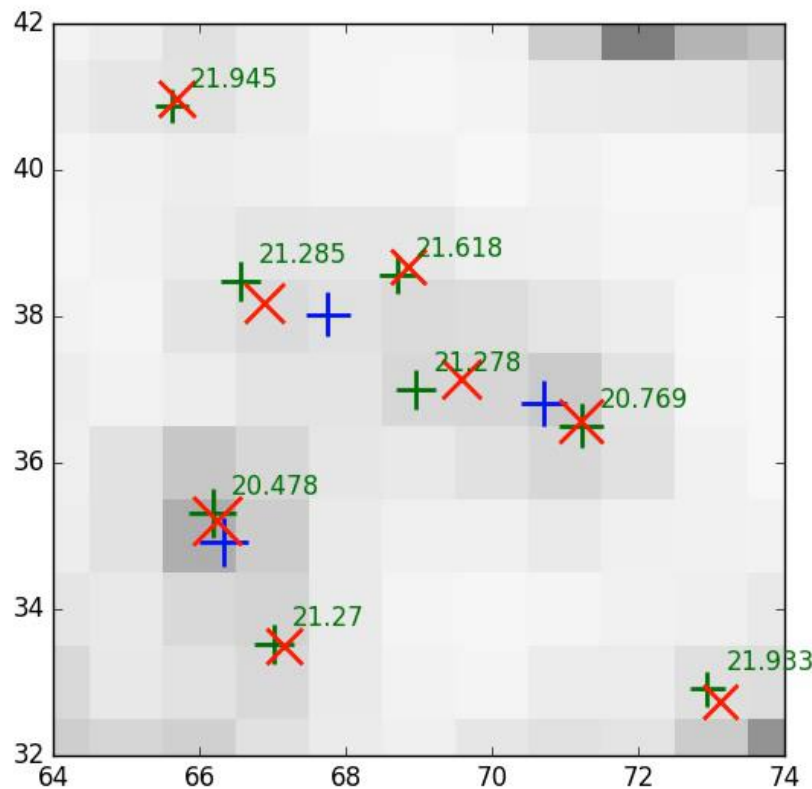
Deterministic Catalogue of SDSS Data



Deterministic Catalogue of HST Data

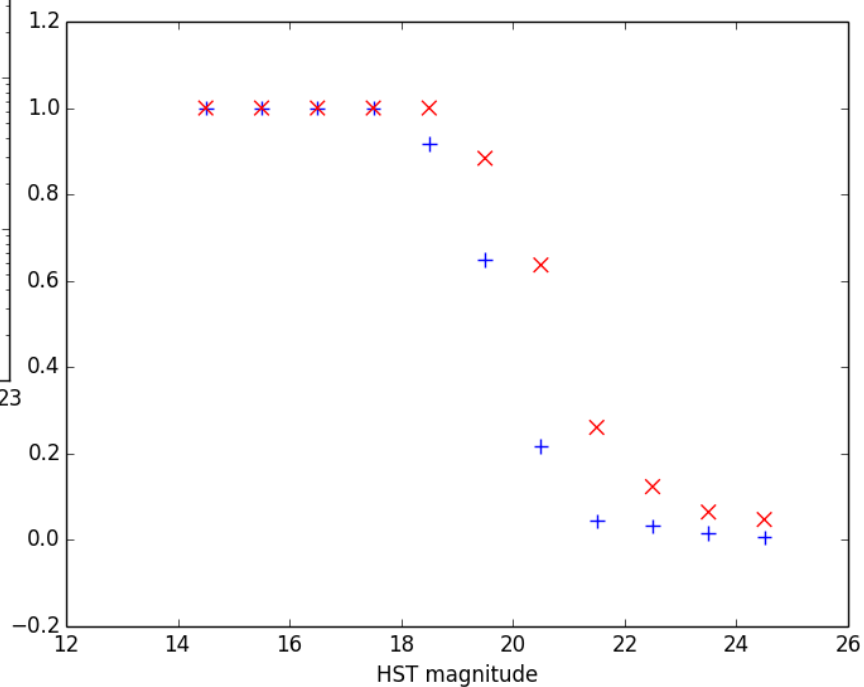
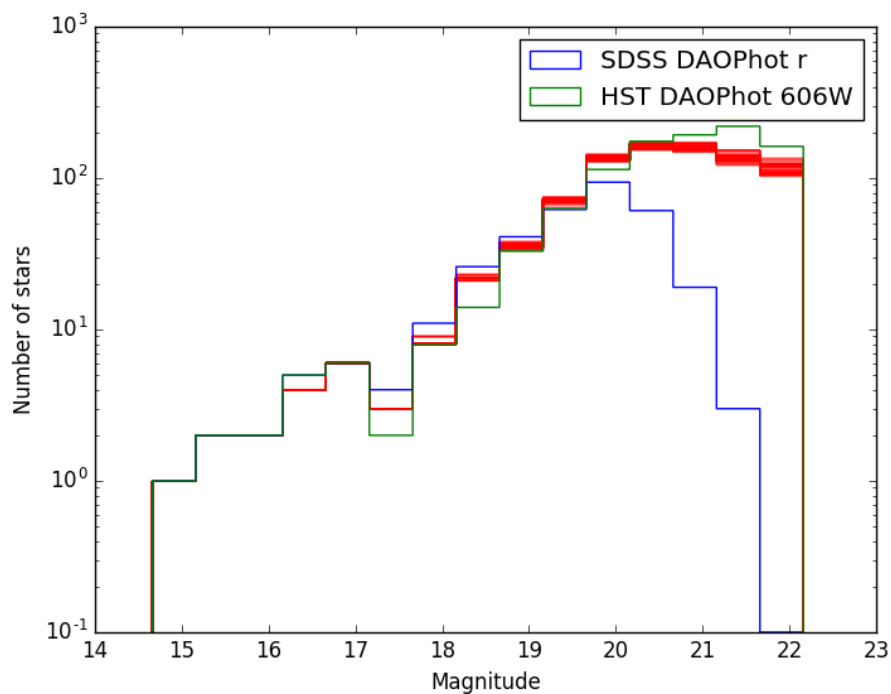


Probabilistic Catalogue of SDSS Data



SDSS DAOPhot
HST DAOPhot
SDSS PCat-Dnest

Completeness



Reversible Jump MCMC

- Allows proposals to change dimensionality of model
 - Move m takes x and generates auxiliary u to propose x'
 - Move m' takes x' and generates auxiliary u' to propose x
 - $\dim x + \dim u = \dim x' + \dim u'$ and $(x, u) \leftrightarrow (x', u')$ one-to-one

$$\alpha(x \rightarrow x') = \min \left(1, \frac{\pi(x') \mathcal{L}(x'|D) j_{m'}(x') g(u')}{\pi(x) \mathcal{L}(x|D) j_m(x) g(u)} \left| \frac{\partial(x', u')}{\partial(x, u)} \right| \right)$$

- For example, birth/death between $x = \{x_1, \dots, x_N\}$ and $x' = \{x_1, \dots, x_{N+1}\}$ has $u = x_{N+1}$ and $u' = \emptyset$
 - If birth and death equally likely, sources independent in prior and new source x_2 generated from prior

$$\alpha(x \rightarrow x') = \min \left(1, \frac{\pi(N+1) \mathcal{L}(x'|D)}{\pi(N) \mathcal{L}(x|D)} \right)$$

Catalogue Priors

- Prior that sources are independent and described by population parameters β :

$$\pi(\{\ell_i, \mathcal{L}_i, F_i\}_{i=1}^N, \beta) = \pi(\beta)\pi(N|\beta) \prod_{i=1}^N \pi(\ell_i, \mathcal{L}_i, F_i|\beta)$$

- β can describe both spatial and flux distributions
- What should the prior on the number of sources look like? What do we mean by “the number of sources”?

How many sources are there with a flux above F_{min} ?

- Prior on N through putting a log uniform prior on expected number of sources $\langle N \rangle$?

$$\log \frac{\pi(N+1)}{\pi(N)} = \log N - \log(N+1) \approx -\frac{1}{N}$$

Source Number Prior

- But is this prior enough to counteract the fact that models with more sources will fit better?
- What about a prior that penalizes the $(N + 1)^{\text{th}}$ source based on the expected improvement in χ^2 under the null hypothesis that there are N sources?

$$\log \frac{\pi(N + 1)}{\pi(N)} = -\frac{3}{2}$$

How many sources meaningfully affect the current data?



What is the most compact representation of the data?

Conclusion

- Probabilistic catalogue samples are an ensemble of catalogues inferred from the data
- A point source population can be distinguished from a diffuse source, even if the individual sources are below the detection threshold
- Overlapping point sources can be better deblended
- This ensemble of catalogues captures the inherent degeneracies of point source identification