

Quantification of Discovery in Astrophysics

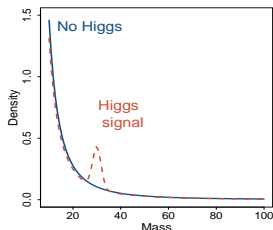
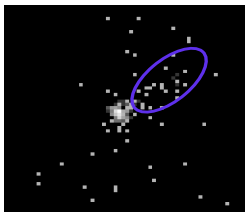
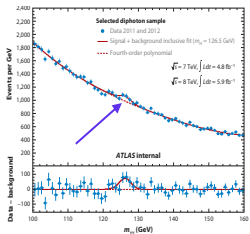
Frequentist and Bayesian Perspectives

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HEAD Meetings 2017, Sun Valley, Idaho

Searching for Structure



- **Bump Hunting:** Is there a bump? *E.g., spectral line or Higgs Boson.*
- Are circled photons due to background or a quasar jet?

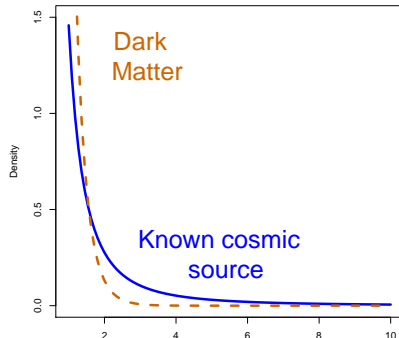
Scientific and Statistical Issues

- High-stakes science: discovery vs. estimation.
- Model selection is much harder than estimation.
- Frequentist and Bayesian methods: different conclusions.
- Is a non-partisan approach possible?

Comparing Models

Compare two different models

- Is a spectral continuum
 - a Bremsstrahlung *or*
 - a Power Law?
- Is γ -ray emission due to
 - known cosmic sources *or*
 - dark matter?



Neutrino Mass Hierarchy

- normal ($\Delta m_{32}^2 > 0$) vs inverted hierarchy ($\Delta m_{32}^2 < 0$)
- $|\Delta m_{32}^2|$ well constrained, degeneracy of sign with other parameters.

Outline

- 1 Using P-values For Discovery
- 2 Bayesian Discovery
- 3 Examples: Mass Hierarchy and Bump Hunting
- 4 Advice and Resources

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Statistical Framework for Discovery

Model / Hypothesis Testing

H_0 : The null hypothesis (e.g., no jet; known cosmic sources)

H_1 : The alternative hypothesis (e.g., jet; dark matter)

- Without further evidence, H_0 is presumed true.
- “Deciding” on H_1 means scientific discovery: new physics.
- **Model Selection**: No presumed model. (normal/inverted hierarchy)

Appropriate Statistical Approach Depends on

- Is H_0 the *presumed* model? or more than 2 possible models?
- Is H_0 a special case of H_1 , “nested models”
- Parameters: (i) Unknown values under H_0 ?
(ii) No “true value” under H_0 ?, (iii) Boundary concerns.
- Bayesian vs. Frequentist methods

Statistical Criterion for Discovery

The most common criterion is the p-value,

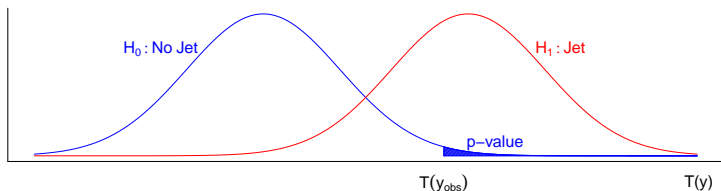
$$\text{p-value} = \Pr \left(T(y) \geq T(y_{\text{obs}}) \mid H_0 \right)$$

- $T(\cdot)$ is a *Test Statistic*, e.g., $\Delta\chi^2$ or likelihood ratio statistic

$$\text{Likelihood Ratio Test} = -2 \log \frac{\max_{\theta} p_0(y \mid \theta)}{\max_{\theta} p_1(y \mid \theta)}$$

Likelihood under H_0

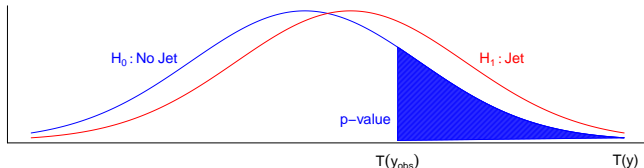
Likelihood under H_1



Computing p-values

The most common criterion is the p-value,

$$\text{p-value} = \Pr \left(T(y) \geq T(y_{\text{obs}}) \mid H_0 \right)$$



Requires distribution of $T(y)$ under H_0

- Distributions depend on unknown parameters
(e.g., continuum / background parameters)
- Standard Theory: models nested, all parameters have values under H_0 , “large” data set. *... often violated in astro/physics*
- Monte Carlo / Bootstrap infeasible with 5σ criterion.

Misuse of P-values

The most common criterion is the p-value,

$$\text{p-value} = \Pr\left(T(y) \geq T(y_{\text{obs}}) \mid H_0\right) \text{ with } T = \text{test statistic}$$

But....

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Psychology journal bans *P* values

Test for reliability of results 'too easy to pass', say editors.

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26 February 2015 | Clarified: 09 March 2015

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Statisticians issue warning over misuse of P values

Policy statement aims to halt missteps in the quest for certainty.

Monya Baker

07 March 2016

(ASA Statement on Statistical Significance and P-values)

February 5, 2016

The Problem with P-values

The misuse of P-values:

- **Do not measure relative likelihood of hypotheses.**
- Large p-values do not validate H_0 .
- May depend on bits of H_0 that are of no interest.
- **Single filter** for publication / judging quality of research.
- **Should be viewed as a data summary, not the summary**

*Reviewers, Editors, and Readers want a simple
black-and-white rule: $p < 0.05$, or $> 5\sigma$.*

*But, statistics is about quantifying uncertainty,
not expressing certainty.*

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- 2 Bayesian Discovery**
- 3 Examples: Mass Hierarchy and Bump Hunting
- 4 Advice and Resources

A Bayesian Criterion for Discovery

When trying to detect a jet, suppose we find

$$\text{p-value} = \Pr\left(T(y) \geq T(y_{\text{obs}}) \mid \text{No Jet}\right) = 0.0001$$

Questions

- Can we conclude that there is probably a Jet?
- Does $\Pr(\text{Data} \mid \text{No Jet})$ small imply $\Pr(\text{No Jet} \mid \text{Data})$ is small?

Order of conditioning matters!

Consider $\Pr(A \mid B)$ and $\Pr(B \mid A)$ with

A:

B:

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Order of conditioning matters!

Consider $\Pr(A \mid B)$ and $\Pr(B \mid A)$ with

A: A person is a woman.

B: A person is pregnant.

Bayesian Methods

Bayes Theorem

$$\Pr(\text{Jet} \mid \text{Data}) = \frac{\Pr(\text{Data} \mid \text{Jet}) \Pr(\text{Jet})}{\Pr(\text{Data} \mid \text{Jet}) \Pr(\text{Jet}) + \Pr(\text{Data} \mid \text{No Jet}) \Pr(\text{No Jet})}$$

Bayesian methods

- have cleaner mathematical foundations
- more directly answer scientific questions

... *but they depend on **prior distributions***

- $\Pr(\text{Jet})$ = probability of a Jet before seeing data.

Prior distributions must also be specified for model parameters.

The Problem with Priors

Bayesian Criteria for Discovery:

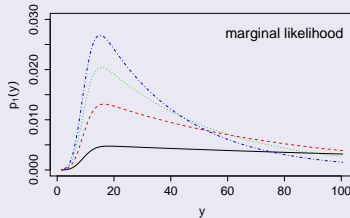
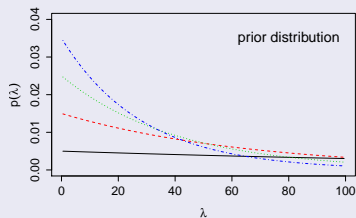
$$\text{Bayes Factor} = \frac{p_0(y)}{p_1(y)} \text{ with } p_i(y) = \int p_i(y|\theta)p_i(\theta)d\theta.$$

$$\Pr(H_0 | y) = \frac{p_0(y)\pi_0}{p_0(y)\pi_0 + p_1(y)\pi_1} = \frac{\pi_0}{\pi_0 + \pi_1(\text{Bayes Factor})^{-1}}$$

Example: (simplified) Higgs search

Likelihood: $y|\lambda \sim \text{Poisson}(10+\lambda)$

Test: $\lambda = 0$ vs $\lambda > 0$



Value of $p_1(y)$ depends on prior!

Choice of Prior Matters!

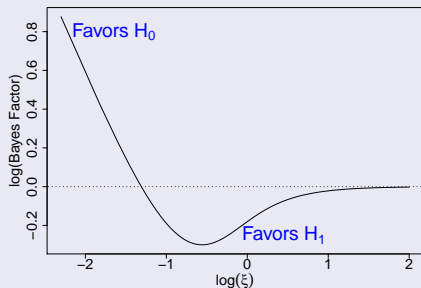
Bayes Factor

$$H_0 : y \sim \text{Poisson}(10).$$

$$H_1 : y \sim \text{Poisson}(10 + \lambda).$$

with $\lambda \sim \exp(\xi)$

- Observe $y = 15$
- $\log(\text{Bayes Factor})$



Must think hard about choice of prior and report!

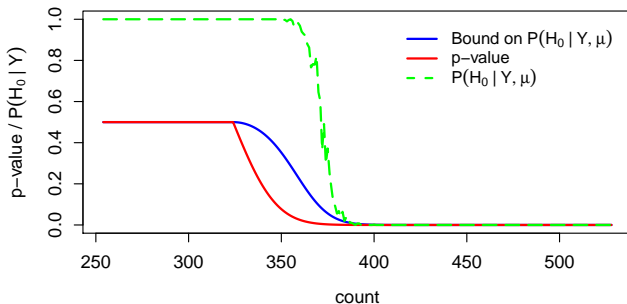
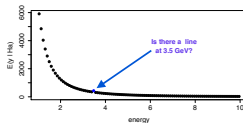
Frequentist vs Bayesian: Does it Matter?

Model Testing and Model Selection

- Frequency and Bayesian methods **may not agree**.
 - Bayes automatically penalizes larger models (*Occam's Razor*)
 - and adjusts for trial factors / look elsewhere effect.
- Choice of prior distribution **is often critical**.
- **Difficult cases:** Dimension of model parameters differ.
 - Higgs search: location and intensity of bump above bkgd.
 - Added structure in image.
- Anti-conservative: $p\text{-value} \ll \Pr(H_0 | y)$.
- *Remember:*
 $p\text{-value}$ and $\Pr(H_0 | y)$ quantify different things!

Interpreting $p\text{-value}$ as $\Pr(H_0 | y)$ may significantly overstate evidence for discovery.

Example: Searching for a bump above background.



.... but researchers interpret p -value as $\Pr(H_0 | y)$.

Solution: Report both.

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Normal Hierarchy versus Inverted Hierarchy

Non-nested parameterized models

H_0 : normal hierarchy i.e., $\Delta m_{32}^2 \leq 0$

H_1 : inverted hierarchy i.e., $\Delta m_{32}^2 > 0$

... recall $|\Delta m_{32}^2|$ is well constrained.

Computing a p-value using LRT

- Non-nested models: If no unknown parameters in either model.
 - LRT follows a Gaussian distribution under H_0 or H_1 .
- With unknown parameters *(e.g., Bremsstrahlung vs. Power Law)*
 - Std theory (Wilks, Chernoff) does not apply: dist'n of LRT unknown.
 - Problem-specific theory, requires strong assumptions.
 - What about uncertainty in $|\Delta m_{32}^2|$?
 - PPP-values / parametric bootstrap, *(e.g., Protassov et al., ApJ, 2002)*.

Back to Monte Carlo / Bootstrap? at 5σ ??

Is There an Easier Solution?

Two paradigms for statistical inference:

Likelihood: inference based on $p(y | \theta)$ and *LRT, p-value, etc.*

Bayesian: inference based on $p(\theta | y) \propto p(y | \theta)p(\theta)$.

Model Fitting

- Specify one model, fit parameters, estimate uncertainty.
- Frequency and Bayesian methods tend to agree.
- Choice of prior distribution is often not critical.

Some “model selection” can be accomplished via model fitting, e.g., confidence intervals.

Normal versus Inverted Hierarchy: Easier Way?

Non-nested parameterized models

H_0 : normal hierarchy i.e., $\Delta m_{32}^2 \leq 0$

H_1 : inverted hierarchy i.e., $\Delta m_{32}^2 > 0$

Is there an easier solution??

Why not just compute $\Pr(H_0 | y) = \Pr(\Delta m_{32}^2 \leq 0 | y)$?

In this case Bayes Criterion is particularly easy:

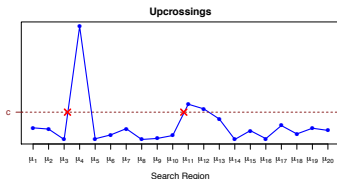
$$\text{Posterior Odds} = \frac{\Pr(\Delta m_{32}^2 \leq 0 | y)}{\Pr(\Delta m_{32}^2 > 0 | y)}$$

...model fitting with Δm_{32}^2 a free parameter.

*One model and one prior, easy to compute,
not sensitive to prior... what's not to like?*

Bayesian solution is easier in this case.

Bump Hunting: Frequency vs Bayes



Frequency Methods:

- Fixed bump location: *standard methods apply*
- Multiple testing problem.

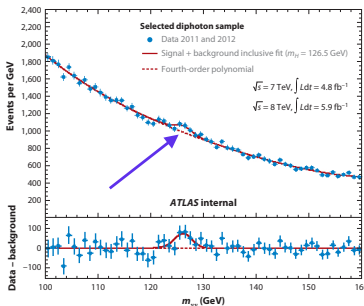
(Algeri, van Dyk et al., 2016)

Bayesian Methods:

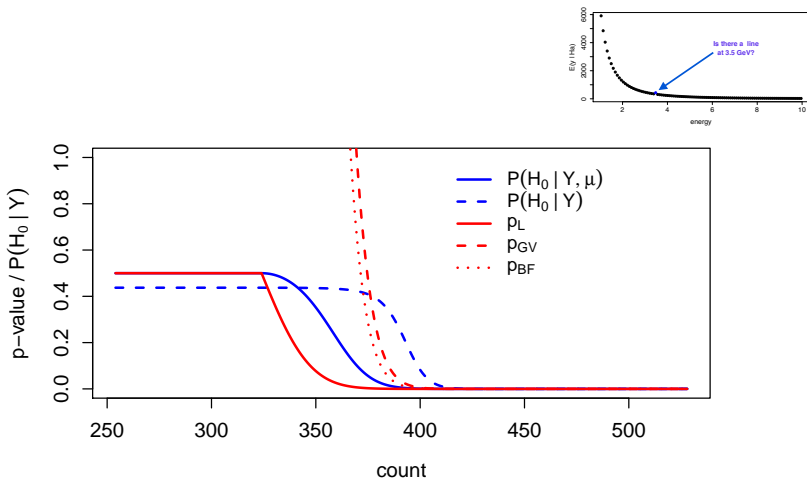
– *Prior specification is key.*

- Intensity parameter
P-values favor H_1
Use prior most favorable for H_1 .
Bound $\Pr(H_0 \mid \text{Data})$.
- Location:
Prior automatically corrects
for multiple testing.

(van Dyk and Jones, 2017+)



Bump Hunting: Frequency vs Bayes



Prior on location naturally and simply corrects for multiple testing.

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Frequentist or Bayesian?

Do you have to choose??

- Bayes prescribes methodology.
- Frequentists evaluate methods.
- Frequency evaluation of Bayesian methods.
- Model fitting: often little difference in fits and errors.
- Why not control rate of false detection
and assess probability of new physics?
- Why throw away half of your tool box?

Be open to both Bayesian and Frequency based methods.

- Now lots of Bayesian and Frequentist methods in HEA.
- My experience with cosmologists and particle physicists.

Strategies

What is a astrophysicist to do?






- Controlling false discovery is critical in physical sciences.
- Comparing p-values with a predetermined significant level can control false discovery.... *if used with care, e.g., no cherry picking!*
- When confronted with small p-values researchers *...even statisticians!!...* may believe H_0 is unlikely.
- Bayesian solutions can better quantify likelihood of H_0 / H_1 .
- **Solution:** Compute both *global* p-value *and* Bayes Factor.

But be Careful...

- 1 *quantification of p-values in non-standard problems*
- 2 *choice and validation of prior distributions*

remain challenging!

References

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