

# DEM for the statistically challenged $\chi^2$ vs. L1 norm minimization?



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Chloé Guennou

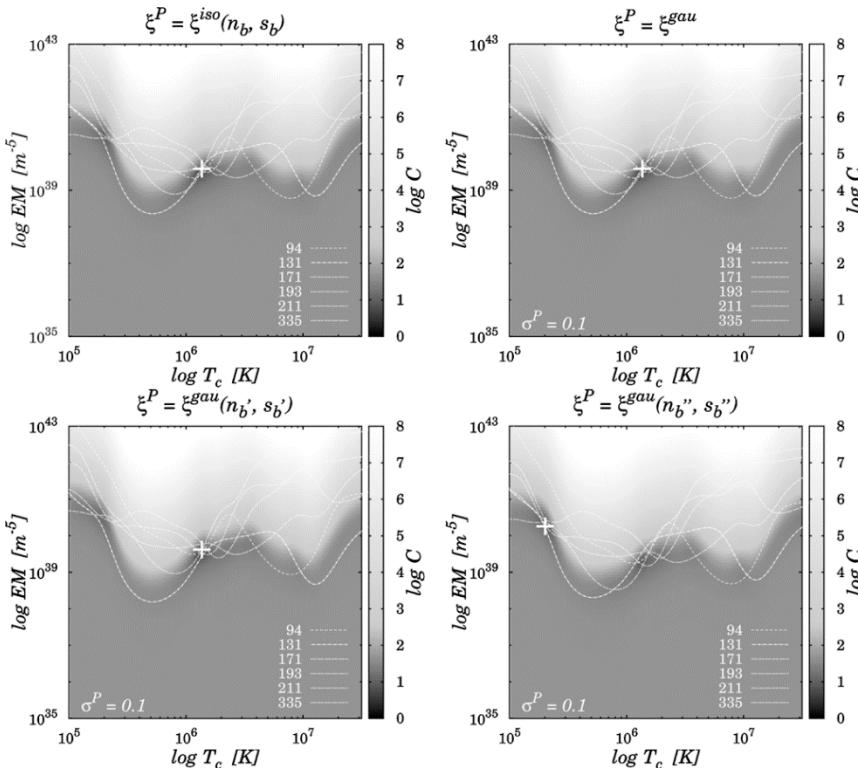
Instituto de Astrofísica de Canarias, Spain

# Merit function (a.k.a. objective function, criterion, etc.)

- Many DEM inversion algorithms based on  $\chi^2$  minimization

$$\chi^2 = \min \left[ \sum_{b=1}^{N_b} \left( \frac{I_b^{obs} - I_b^{th}(\xi)}{\sigma_{unc}} \right)^2 \right]$$

- Example of  $\chi^2$  merit function for isothermal inversion ( DEM  $\xi = EM \delta(T_c)$  )



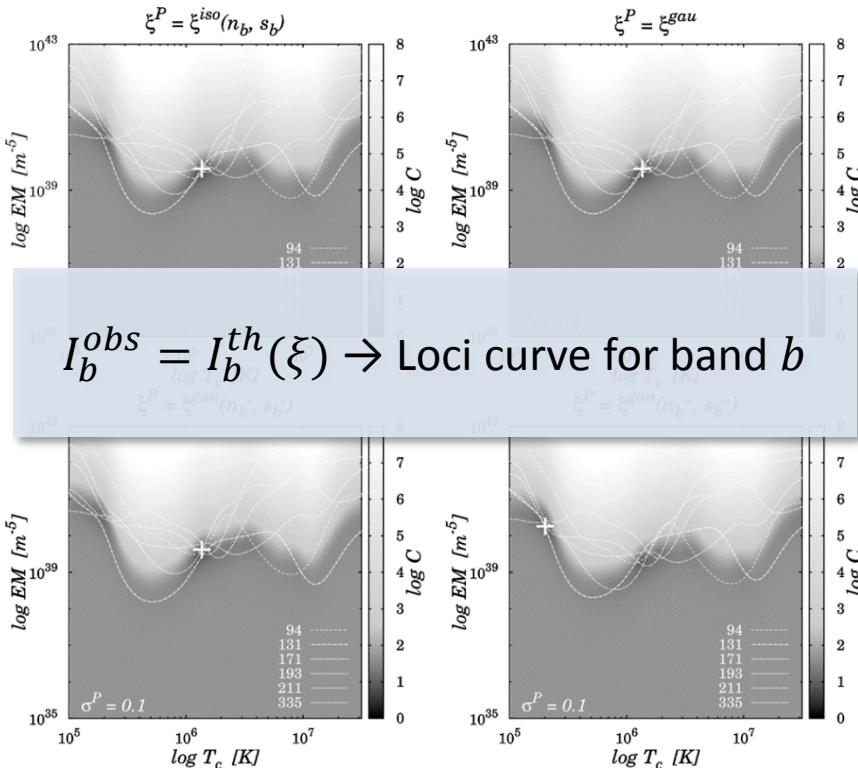
- Goal **is not** to derive a minimization algorithm but to understand the **properties of the merit function**
- Results apply to **all  $\chi^2$ -based inversion schemes**
- Fundamental equivalence** between noise and multithermality

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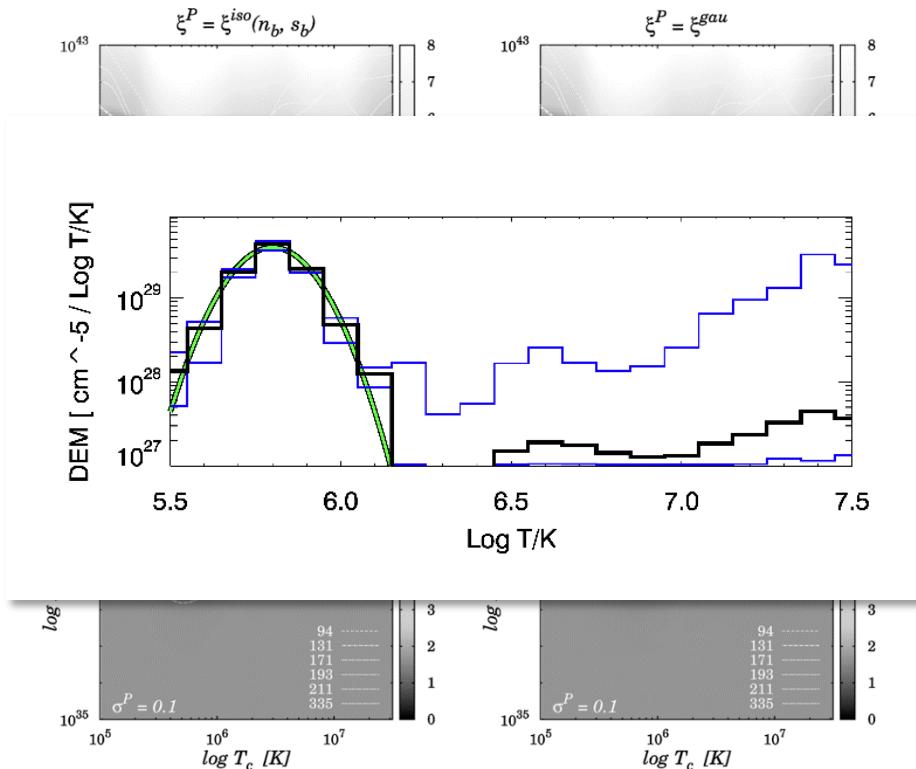
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# Chloé's approach

- 6 AIA bands → can't fit a very complex DEM
- Systematic search of all solutions for a simple test case
  - Gaussian (log-normal) DEM plasma input

$$\xi_{gau}^P = \frac{EM}{\sigma\sqrt{2\pi}} \exp\left(-\frac{[\log T_e - \log T_c]^2}{2\sigma^2}\right)$$

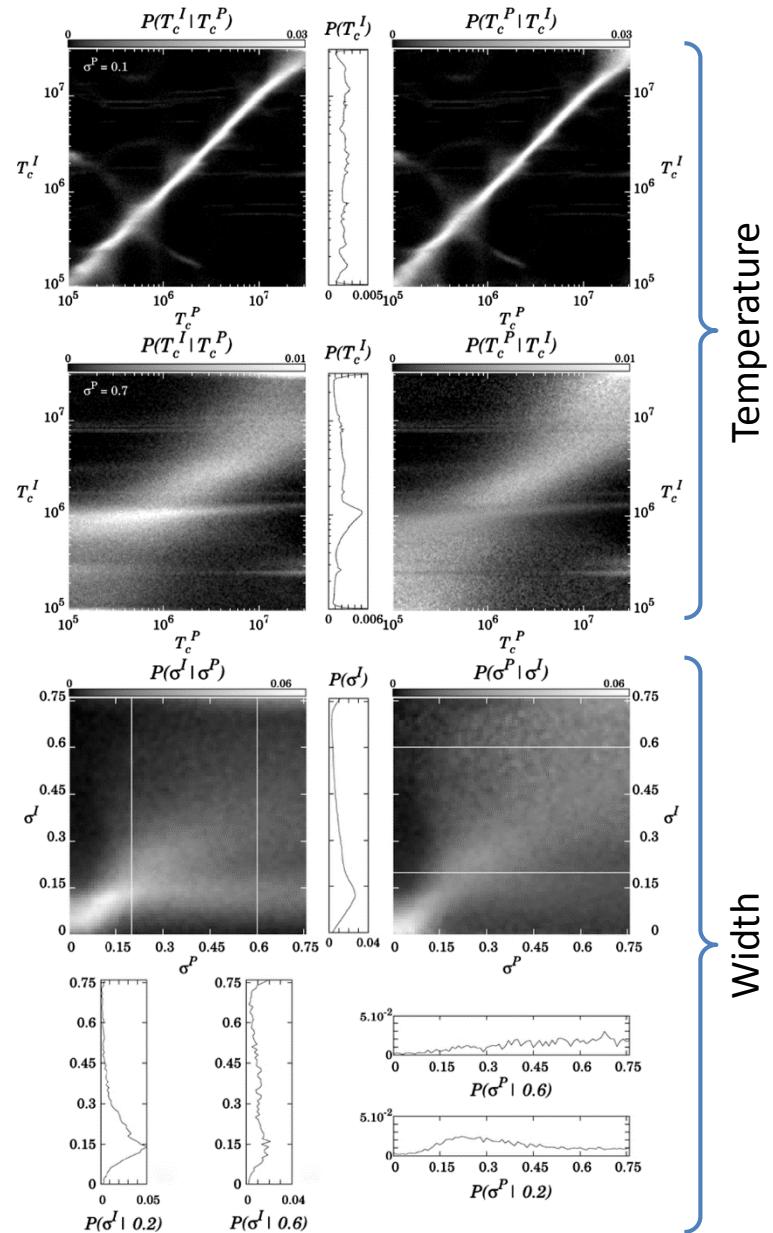
$$= EM \times \mathcal{N}(\log T_e - \log T_c)$$

- Search for Gaussian solutions

$$\chi^2 = \min \left[ \sum_{b=1}^{N_b} \left( \frac{I_b^{obs} - I_b^{th}(\xi)}{\sigma_{unc}} \right)^2 \right]$$

- Detailed uncertainties
  - Photon noise
  - 25% calibration & atomic physics

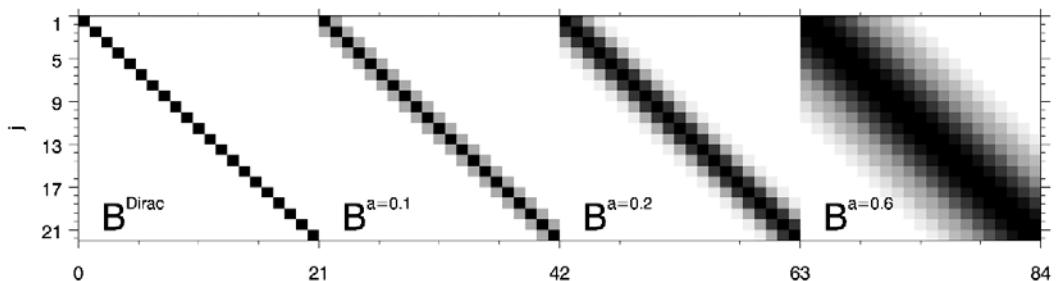
- (some of the) results
  - Broad DEMs poorly constrained
  - Possible bias of the solutions towards
    - $T_c = 1 \text{ MK}$  &  $\sigma = 0.1 \log T$
    - Similar to Weber et al. 2005, ApJ, 635, L101  
(Guennou, C. et al. 2012a, ApJS, 203, 25,  
Guennou, C. et al. 2012b, ApJS, 203, 26)



# Alternative to $\chi^2$

- Cheung, M., Boerner, P., Schrijver, C. et al. 2015, "Thermal Diagnostics with the Atmospheric Imaging Assembly onboard the Solar Dynamics Observatory: A Validated Method for Differential Emission Measure Inversions", ApJ, in press

- Dictionary-based inversion

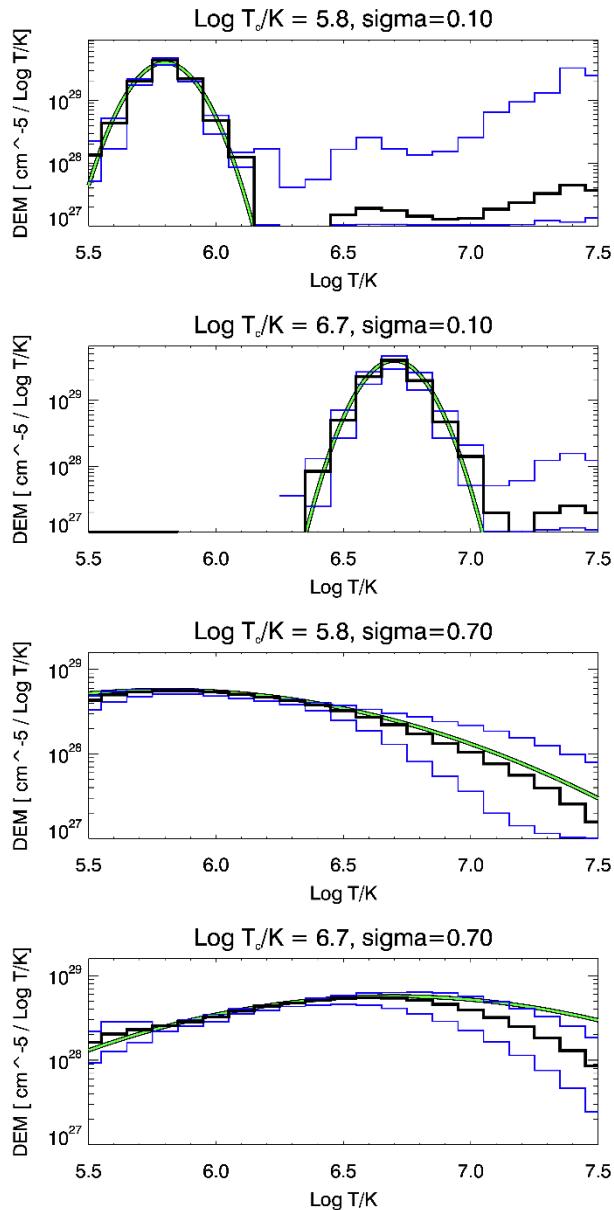


- Not  $\chi^2$ -based
- Minimizes the **L1 norm** of the coefficients  $x_j$ , i.e.

$$\text{LP1 : minimize } \sum_{j=1}^n x_j \text{ subject to}$$

$$\begin{aligned} \mathbf{D}\vec{x} &\leq \vec{y} + \vec{\eta}, \\ \mathbf{D}\vec{x} &\geq \max(\vec{y} - \vec{\eta}, 0), \\ \vec{x} &\geq 0. \end{aligned}$$

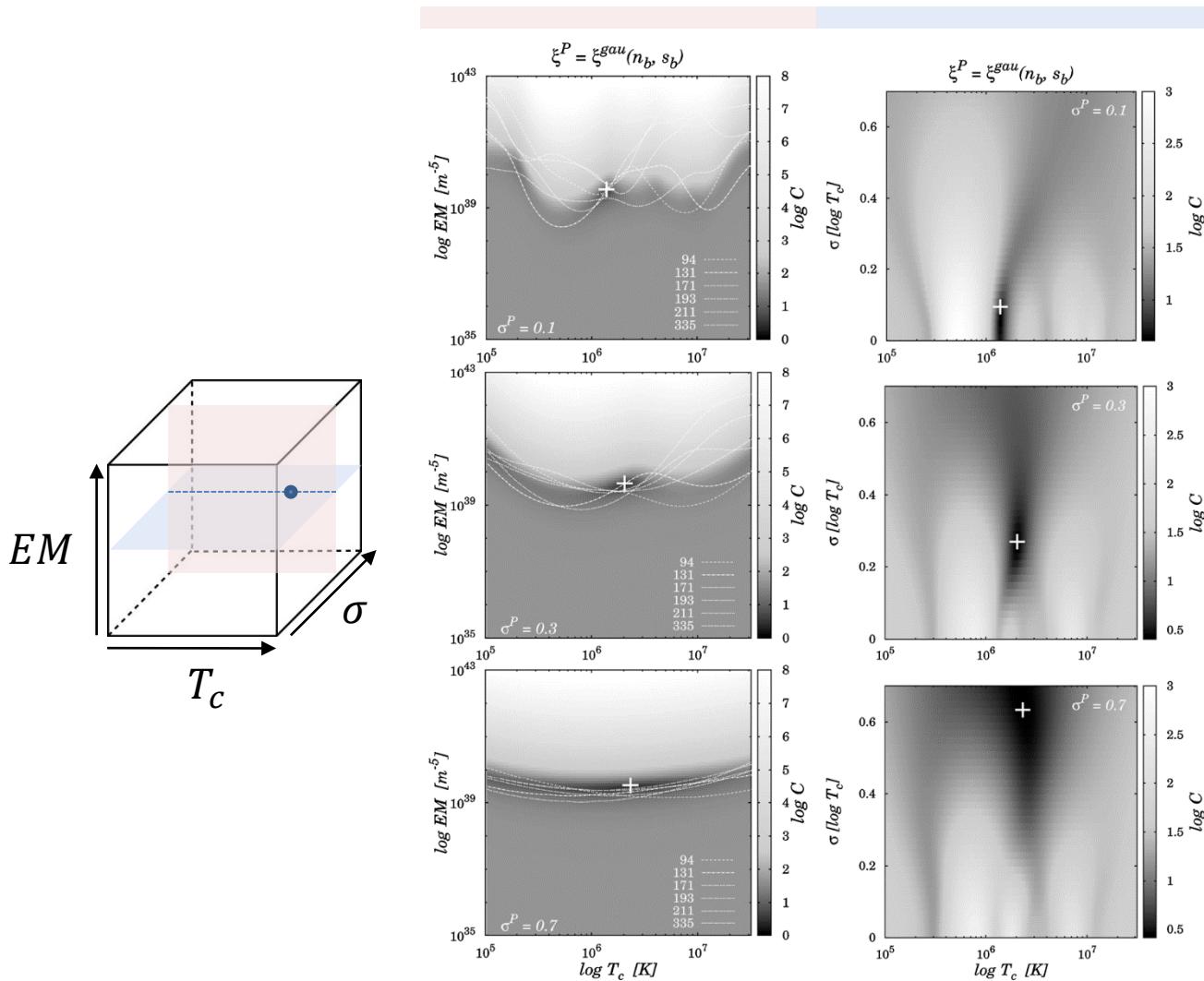
- If  $B = B^{\text{dirac}}$  only,  $\sum_{j=1}^n x_j = EM$
- More robust than  $\chi^2$  for wide DEMs ?



# Back to $\chi^2$ : why are broad DEMs poorly constrained?

$$\chi^2 = \min \left[ \sum_{b=1}^{N_b} \left( \frac{I_b^{obs} - I_b^{th}(\xi)}{\sigma_{unc}} \right)^2 \right]$$

Merit function (a.k.a. objective function, criterion, etc.)

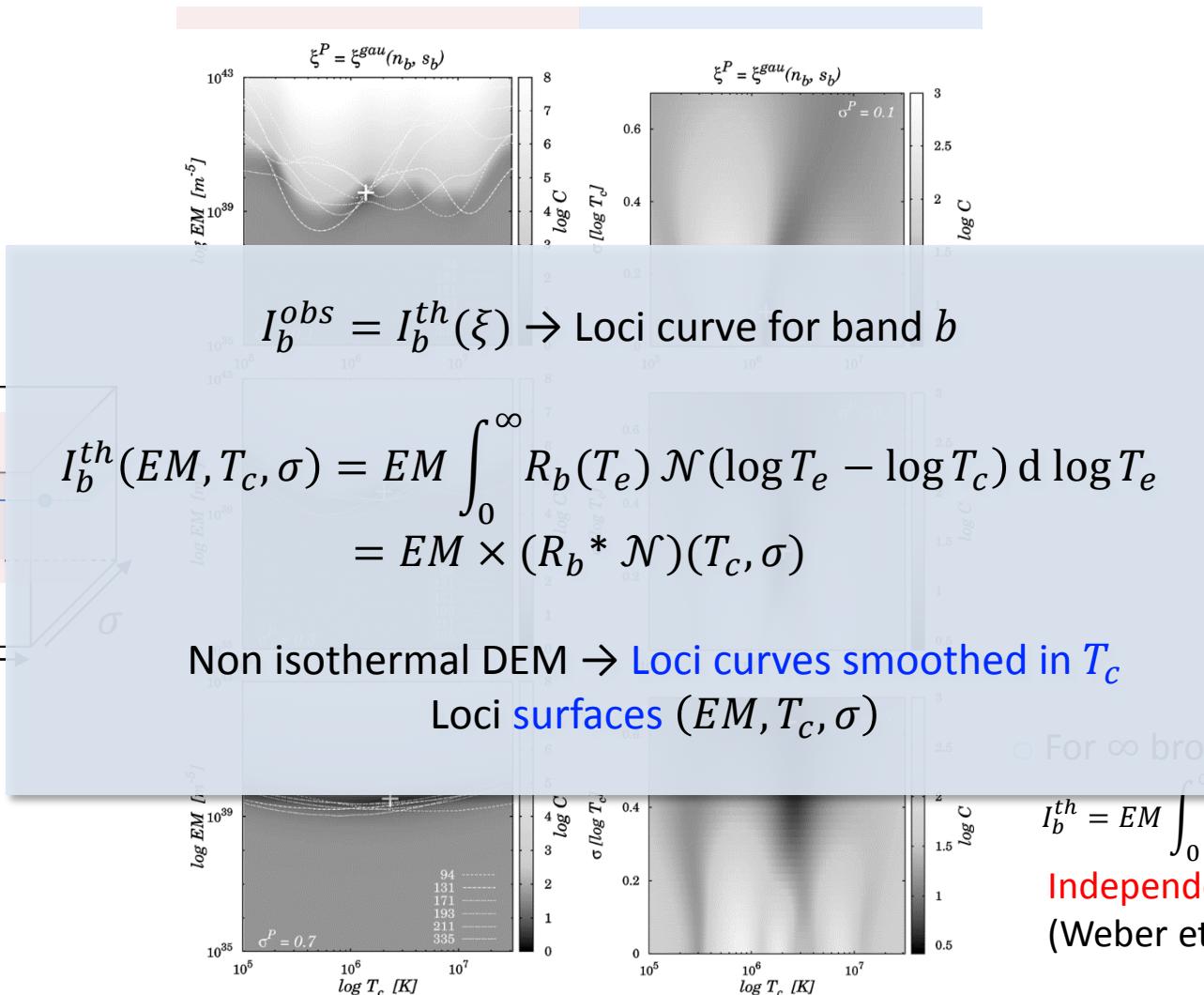


- For  $\infty$  broad DEM
- $$I_b^{th} = EM \int_0^\infty R_b(T_e) d \log T_e$$
- Independent from  $T_c$   
(Weber et al. 2005)

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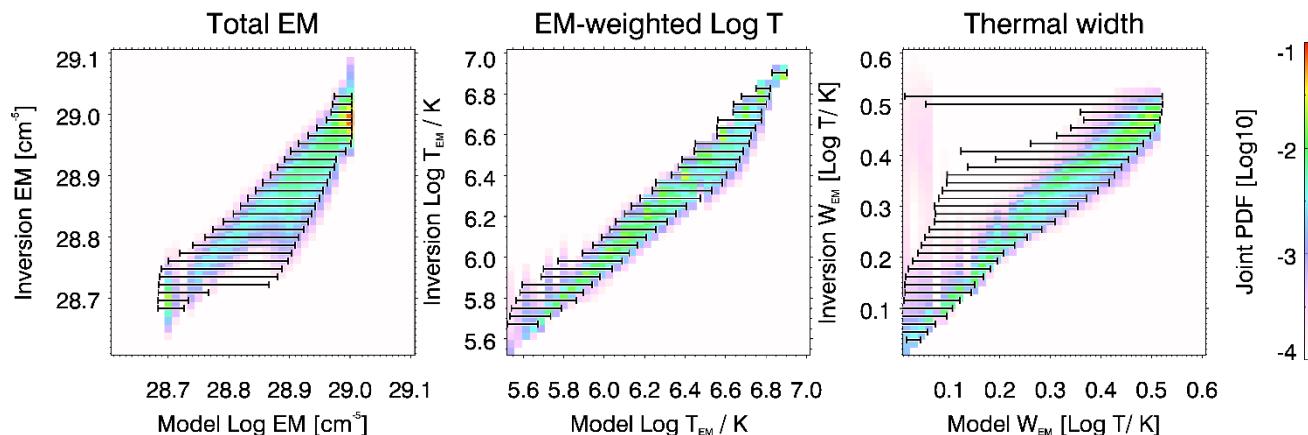
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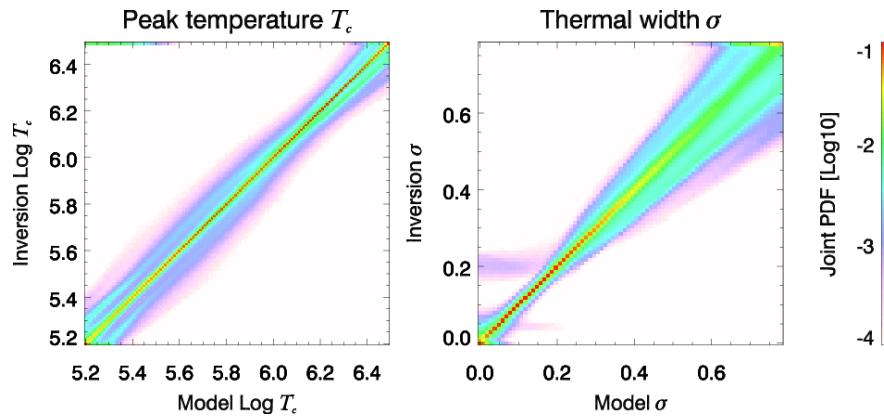


# Comparison $\chi^2$ - L1

- Cheung et al. use `aia_bp_estimate_error` to estimate the uncertainties
  - photon noise, compression and quantization round-off, error in dark subtraction
  - no atomic physics & calibration uncertainties

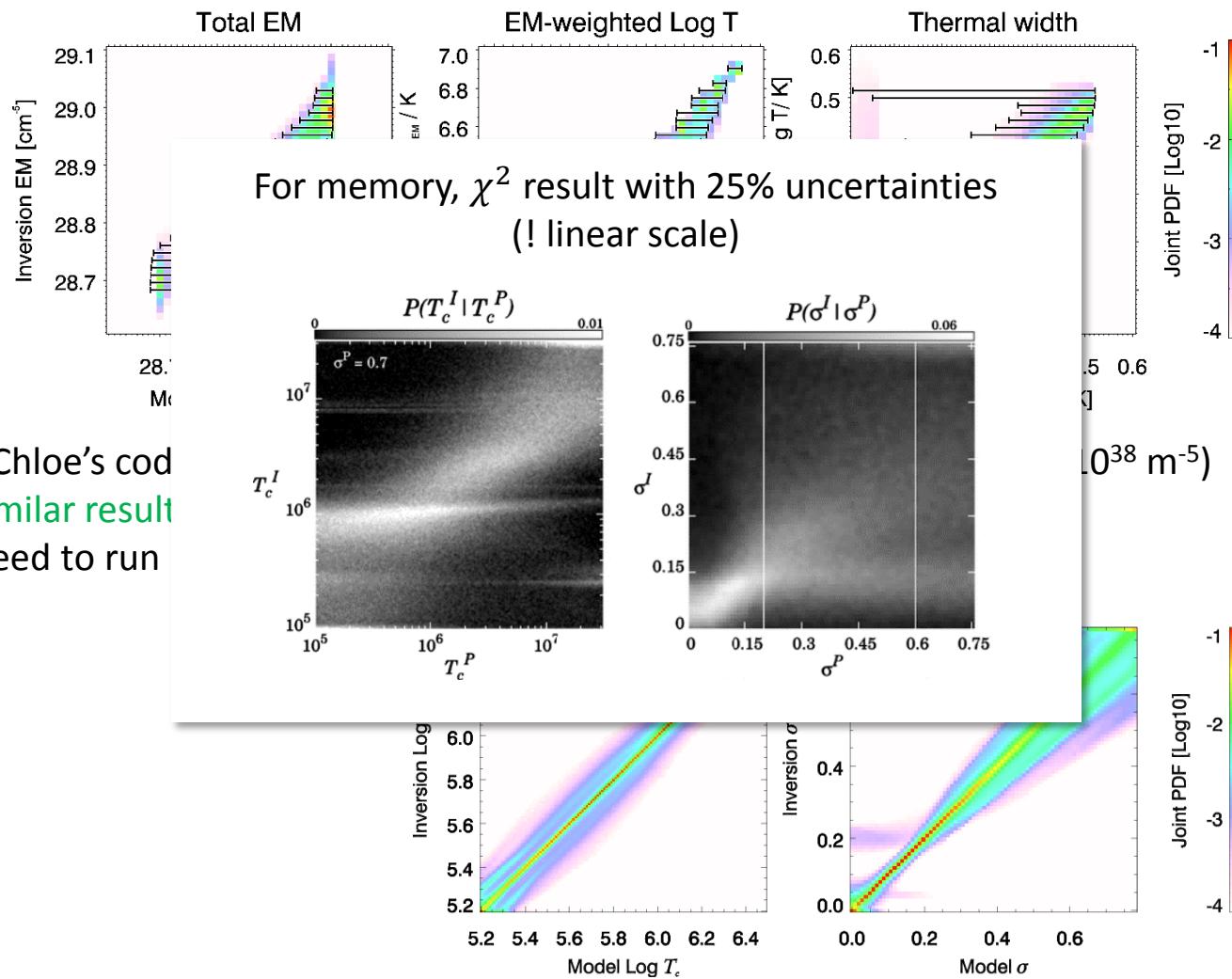


- Run of Chloe's code with the same (<<25%) uncertainties (constant  $EM = 10^{38} \text{ m}^{-5}$ )
  - Similar results !
  - Need to run Mark Cheung's code with 25% uncertainties



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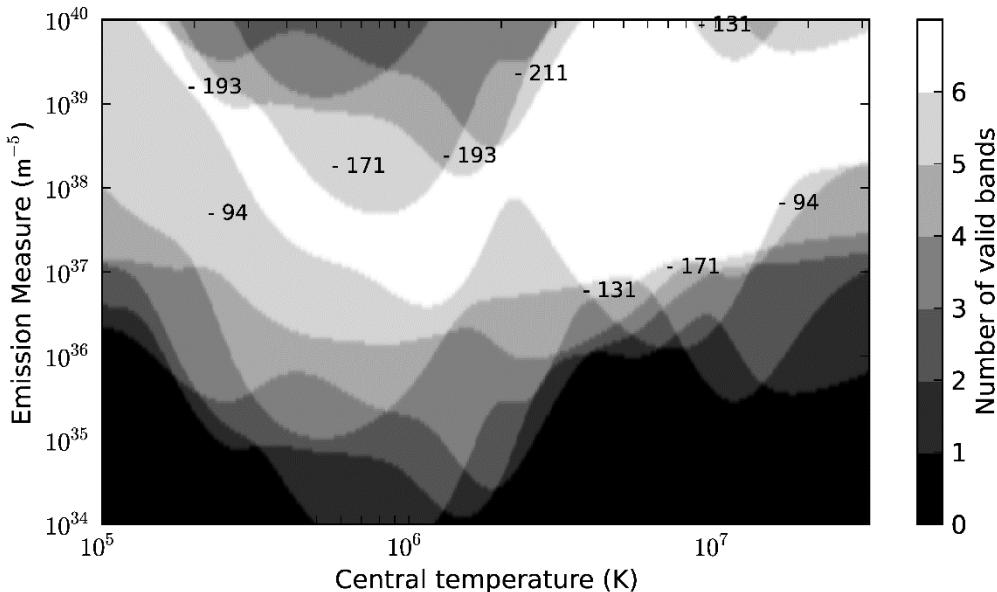
- Run of Chloe's code
  - Similar result
  - Need to run

# To be continued...

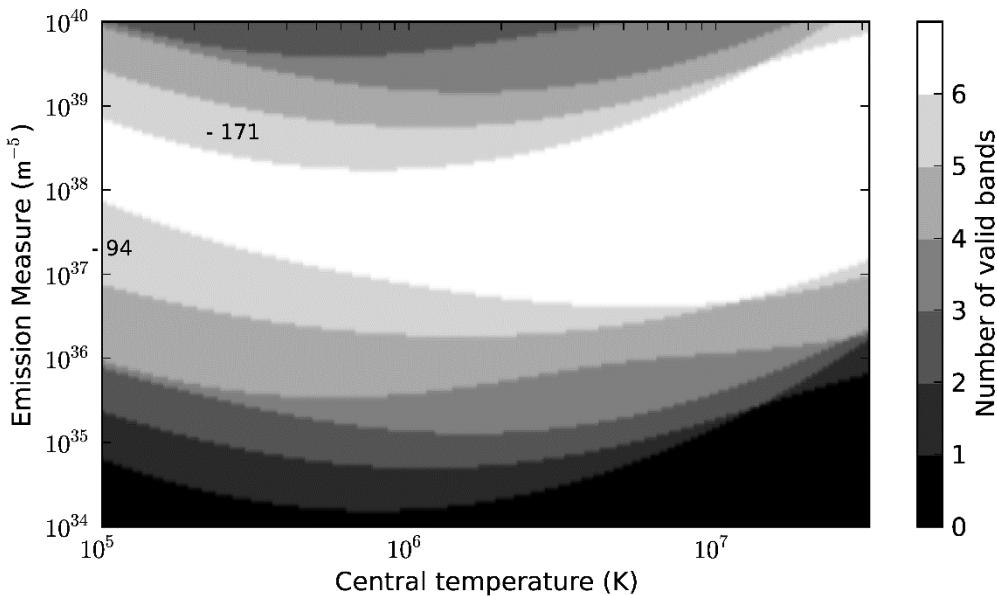
- Mark Cheung's method is all new to me
  - I don't understand yet how the L1 approach can alleviate the difficulties found for broad DEMs
  - That does not mean it's not the case :D
- Discussion started with Mark Cheung
  - Run both codes with the same input DEMs & uncertainties
  - Compare
- ISSI 2016 ...



# AIA signal vs Temperature & EM



Isothermal



$$\sigma = 0.5 \log T_e$$

# Is my plasma isothermal?

