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IVOA Spectral Data Model

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<http://www.ivoa.net/Documents/WD/SpectrumDM/SpectrumDM-20061023.html>

Latest version:

<http://www.ivoa.net/Documents/latest/SpectrumDM.html>

Previous versions:

<http://hea-www.harvard.edu/~jcm/vo/docs/spec0.93.html>

Note: ivoa.net links are not yet active; the above are placeholders.

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Abstract

We present a data model describing the structure of spectrophotometric datasets with spectral and temporal coordinates and associated metadata. This data model may be used to represent spectra, time series data, and segments of SED (Spectral Energy Distributions) and other spectral or temporal associations.

Status of this document

This is a Working Draft, developed with the intention to support the Simple Spectral Access Protocol. The working group seeks confirmation that comments have been addressed to the satisfaction of the community.

The DM Working Group expects to promote this document to Proposed Recommendation status in essentially its present form once implementations are complete. We urge both data providers and developers of client applications to review this document and notify the WG of any major problems before 2003 NOVEMBER 13. We further recommend that since this document is expected to be stable except for minor amendments, it should form the basis for trial implementations. Please inform the WG if you are planning an implementation.

This IVOA Working Draft, for review by IVOA members and other interested parties, is a draft document and may be updated, replaced, or obsoleted by other documents at any time. It is inappropriate to use IVOA Working Drafts as reference materials or to cite them as other than "work in progress." A list of current IVOA Recommendations and other technical documents can be found at <http://www.ivoa.net/Documents>.

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The **Virtual Observatory (VO)** is general term for a collection of federated resources that can be used to conduct astronomical research, education, and outreach.

The **International Virtual Observatory Alliance (IVOA)** (<http://www.ivoa.net>) is a global collaboration of separately funded projects to develop standards and infrastructure that enable VO applications.

Contents

Part 1 - Spectrum Data Model	5
1 Introduction and Motivation	6
1.1 Change Log	6
2 Requirements	6
3 Spectral Data Model summary	6
3.1 Model Components	6
3.2 Units	11
3.3 UCDS	11
3.4 UTYPEs	12
3.5 Packaging model	12
3.6 Data Model Fields	12
4 Spectral data model Measurement objects	20
4.1 Spectral coordinate	20
4.2 Flux (Spectral Intensity) Object	20
4.3 BackgroundModel Object	22
4.4 Time coordinate	23
4.5 Position coordinate	23
4.6 Accuracy Fields	23
4.6.1 Coordinate bins	23
4.6.2 Uncertainties	24
4.6.3 Resolution	24
4.6.4 Quality	24
4.6.5 Calibration	25
5 Associated Metadata Fields	26
5.1 CoordSys Fields	26
5.1.1 SpaceFrame	26
5.1.2 TimeFrame	27
5.1.3 SpectralFrame	27
5.1.4 RedshiftFrame (also the Velocity frame)	27
5.1.5 STC	28
5.2 Characterization	31
5.2.1 Coverage Fields	31
5.2.2 Region definitions	32
5.3 Derived Data Fields	32
5.3.1 Signal-to-noise	32
5.3.2 Redshift measurement model	32
5.3.3 Variability amplitude	33
5.4 Curation model	33
5.5 Data Identification model	33
5.6 Target model	34
5.7 Spectrum top level object	34

6 Relationship to general VO data models	36
6.1 Extensibility	36
Part 2 - XML schema serialization	37
7 XML schema serialization	38
7.1 XML schema	38
7.2 Instance example	49
Part 3 - VOTABLE serialization	53
8 VOTABLE serialization	54
8.1 Mapping Schema to VOTABLE	54
8.2 A VOTABLE instance	54
Part 4 - FITS serialization	59
9 FITS serialization	60
9.1 Mapping Spectrum to FITS	60
9.2 Expressing the spectrum spatial coordinates in FITS	65
9.3 The SPECSYS keyword	65
9.4 An instance example	66

Part 1: Spectrum Data Model

1 Introduction and Motivation

Spectra are stored in many different ways within the astronomical community. In this document we present a proposed abstraction for spectral data and serializations in VOTABLE, FITS, and XML, for use as a standard method of spectral data interchange.

We distinguish in several places between the implementation proposed in this document, referred to as Version 1, and capabilities proposed for possible later implementation.

1.1 Change Log

Not yet added, pending discussion:

- Photometry support
- Reformat units in Tables 2,3 to OGIP convention

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- Added table numbers
- Changed some defaults in Table 1
- Added flux UCDs for transmission curves, polarized flux
- Amplified discussion of RedshiftFrame
- Added Spectral location and bounds
- Reorganized order of some sections
- Further rationalization of FITS keywords, rewrote FITS section
- Added TUCDn and TUTYPEn

2 Requirements

We need to represent a single 1-dimensional spectrum in sufficient detail to understand the differences between two spectra of the same object and between two spectra of different objects.

We need to represent time series photometry, with many photometry points of the same object at different times.

Finally, we need to represent associations of spectra, such as the segments of an echelle spectrum, or spectral energy distributions (SED) which consist of multiple spectra and photometry points, usually for a single object. The 'Spectral Associations' model will be described in a separate document which builds on the structures described here.

3 Spectral Data Model summary

3.1 Model Components

Our model for a spectrum is a set of one or more data points (photometry) each of which share the same contextual metadata (aperture, position, etc.). Specifically, a spectrum will have arrays of the following values:

- Flux value, with upper and lower statistical (uncorrelated) errors
- Spectral coordinate (e.g. wavelength), central and bin min and max
- (Optionally) Time coordinate, convertible to MJD UTC
- Optional Quality mask
- Optional spectral resolution array

and will have associated metadata including, for example,

- Data collection and Dataset ID
- Exposure time in seconds
- Position of aperture center, given as ICRS degrees (similar to J2000)
- Aperture size in degrees
- Systematic (correlated) error
- Bibcode

In section 4 we elaborate these concepts in detail, including some complications that we explicitly do not attempt to handle in this version. The data model fields and possible values are listed. We distinguish between optional and required fields in the text, as well as via a "Req" column in the tables which has values of R (Required) and O (Optional). Where appropriate we list those values of the physical units which interoperable implementations are required to recognize.

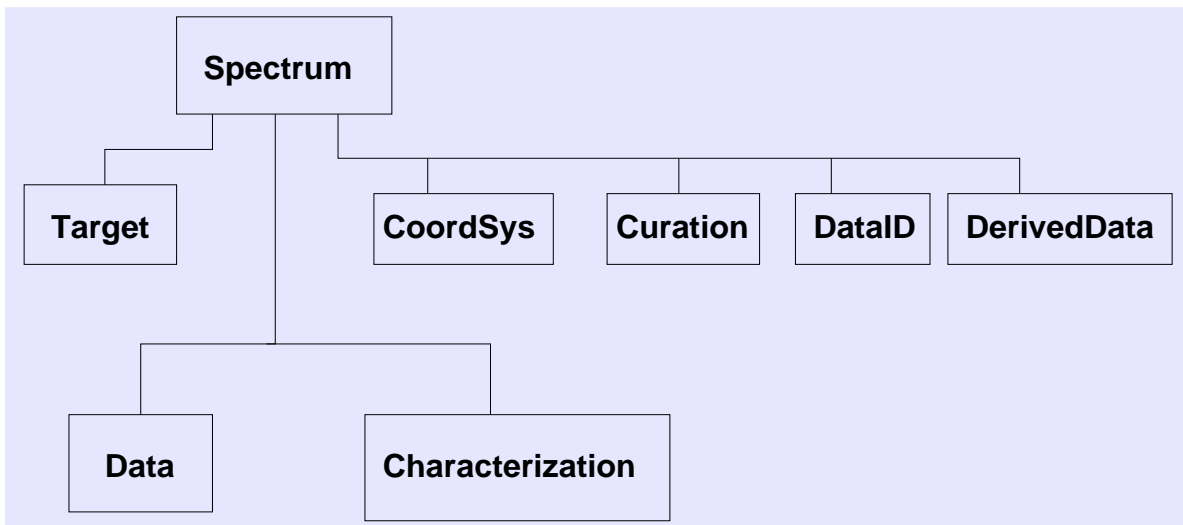


Figure 1: UML class diagram for the spectral data model.

The Coverage, Curation, DataID classes are shown in detail below in Section 4, both in diagram form and text descriptions. The text descriptions are definitive.

The minimal required content is:

- Spectrum model version
- Target name
- Characterization Coverage.Location and Coverage.Bounds (Extent or Start/Stop range) descriptions of the location and extent of the data in the RA, Dec, time and spectral domains
- the Curation.Publisher field
- the descriptions of the spectral coordinate and flux fields including UCD and units (Spectrum.Char.SpectralAxis, Spectrum.Char.FluxAxis)
- For each point: the values of the spectral coordinate and flux.

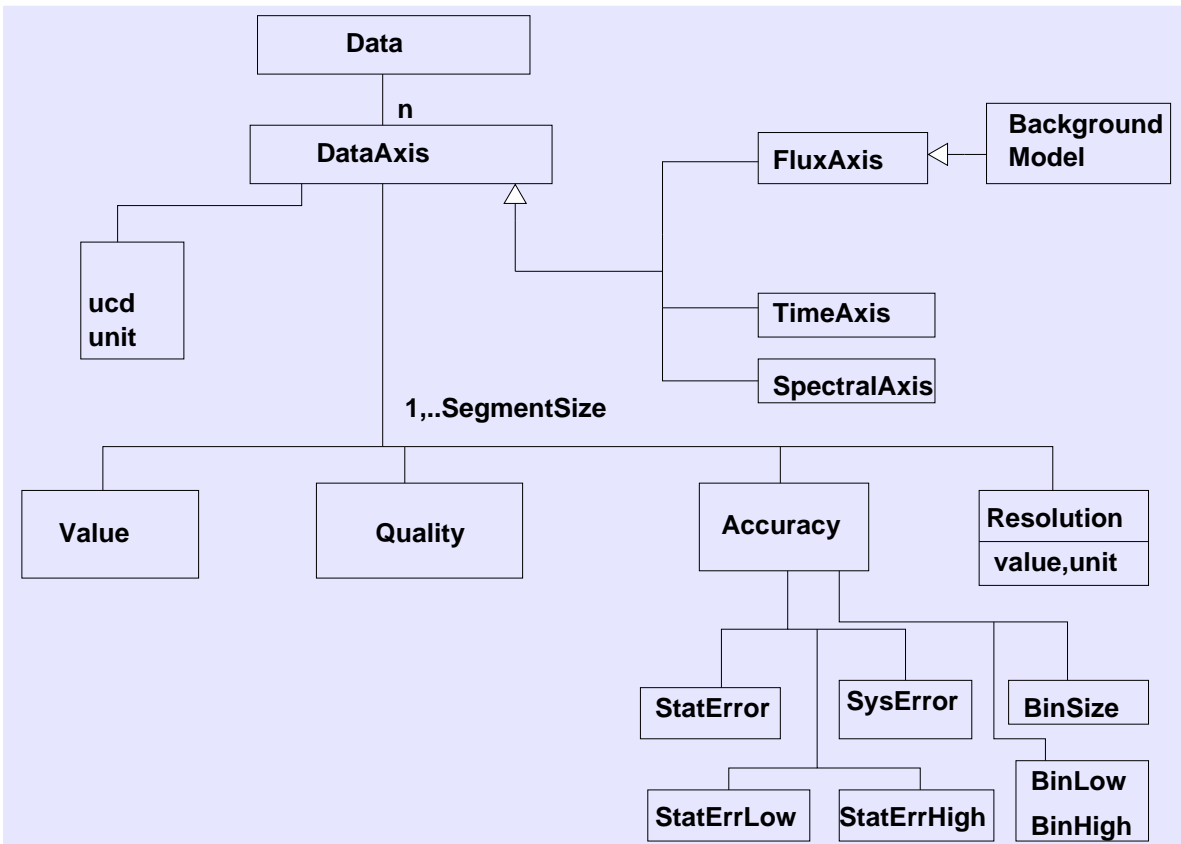


Figure 2: Diagram for Data object

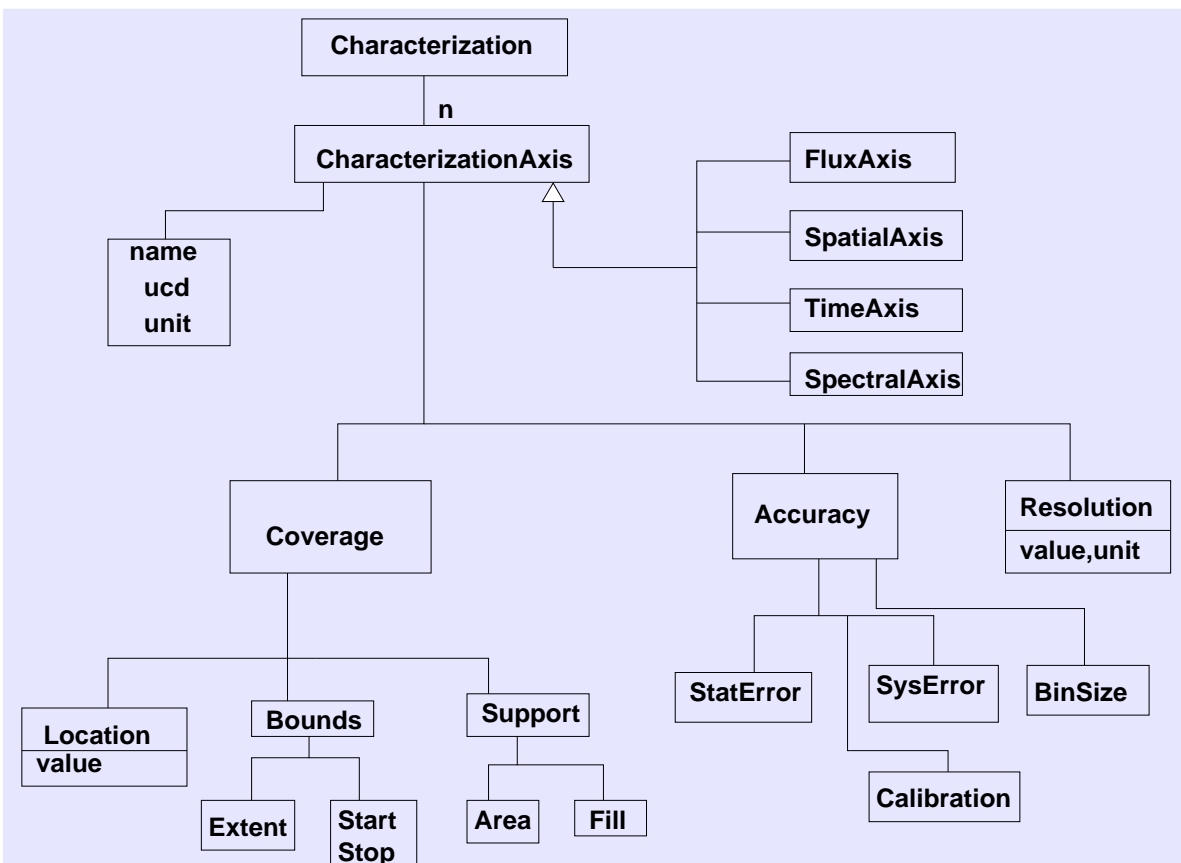


Figure 3: Diagram for Characterization object

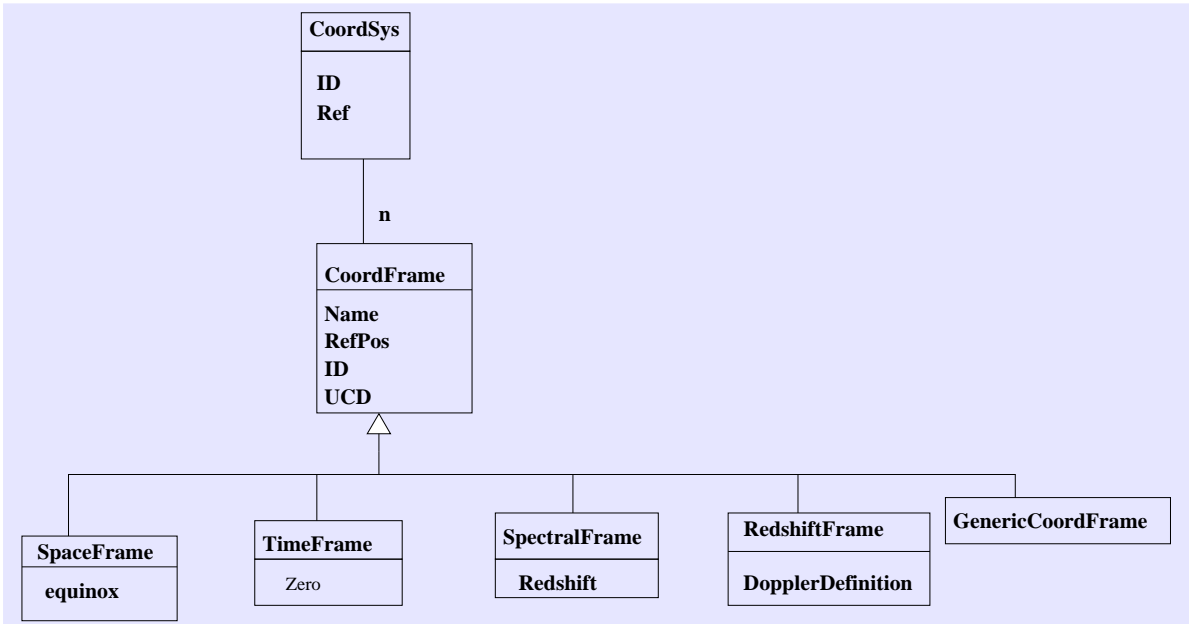


Figure 4: Diagram for CoordSys object

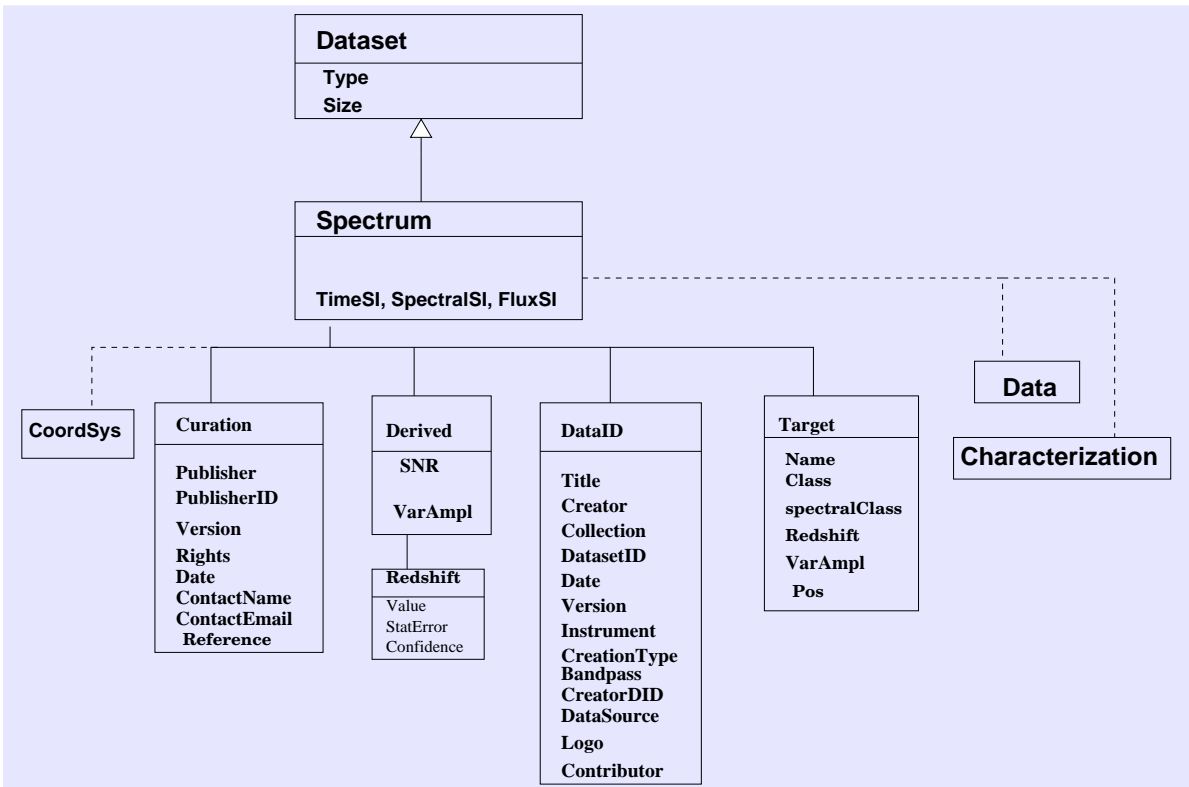


Figure 5: Diagram for remaining metadata: Curation, DataID, Derived, Target objects

3.2 Units

We adopt the WCS/OGIP convention for units: Document OGIP 93-001 (http://legacy.gsfc.nasa.gov/docs/heasarc/ofwg/docs/general/ogip_93_001/ogip_93_001.html).

Briefly, units are given in the form

$10^{(-14)} \text{ erg/cm}^2/\text{s/Hz}$, 10^3 Jy Hz

i.e. with exponents denoted by **, division by /, multiplication by a space.

This format is mostly consistent with the AAS standards for online tables in journals (<http://grumpy.as.arizona.edu/~gschwarz/unitstandards.html>) except for the use of space rather than "." for multiplication and the fact that we do not require the use of SI units.

SI prefixes for units are to be recognized; for instance, the listing of "m" as a known unit for wavelength implies that "cm", "nm", and "um" (with "u" the OGIP convention for rendering "micro") are also acceptable.

Until IVOA generic unit conversion software is mature and widely deployed, it is helpful to interoperable applications to include a representation of the units in "base SI form", including only the base units kg, m, s (and possibly A, sr) with a numeric prefix. Pedro Osuna and Jesus Salgado have proposed a representation in the spirit of dimensional analysis, using the symbols M, L, T to signify kg, m, s respectively and omitting the ** for powers, so that

10^3 Jy Hz

which is equivalent to

$10^{-23} \text{ kg s}^{-2}$

is written compactly as

10^{-23}MT^{-2}

This alternate representation is supported for the main model fields (time, spectral coordinate and flux) only.

Although the spectral model is flexible enough to permit different units for each field, as a matter of style we strongly recommend that whenever possible the same units should be used for compatible fields (e.g. flux and error on flux).

3.3 UCDs

UCDs or Uniform Content Descriptors are the IVOA's standardized vocabulary for astronomical concepts. In this document we use UCDs as field attributes (for example, element attributes in XML) to distinguish alternate physics within the same data model roles - for example, to distinguish frequency versus wavelength on the spectral coordinate 'X-axis'.

The current list of UCDs is <http://cdsweb.u-strasbg.fr/UCD/ucd1p-words.txt> with syntax defined in the UCD recommendation <http://www.ivoa.net/Documents/latest/UCD.html>.

UCDs should be case insensitive.

3.4 UTYPEs

UTYPE was a concept introduced in VOTABLE to label fields of a hierarchical data model. The word is now used generally to mean a standard identifier for a data model field. They are also case-insensitive and are of the form a.b.c.d where the dots indicate a 'has-a' hierarchy with the leftmost element being the containing model and the rightmost element being the lowest level element referred to. This is quite close to a simple XPATH in an XML schema, but we chose not to use slash instead of dot to emphasize that we are only specifying the element type, not the exact position in an instance (so no sophisticated query syntax). We use the terms 'data model field' and 'UTYPE' interchangeably.

3.5 Packaging model

The simple Packaging model for SSA describes the format of the associated dataset. Allowed values for the format (detailed serialization for formats 4 to 7 to be specified in a separate document; only 1-3 are discussed here.)

These packaging values will be part of the SSA protocol response, and are implicit in the individual serializations.

- (1) FITS (standard BINTABLE for Spectrum, defined in this document)
- (2) VOTABLE
- (3) XML (native XML for web services and XML tools)
- (4) text (simple text table with columns of data and no markup)
- (5) text/html
- (6) graphics; a JPG, GIF etc. representation of the data
- (7) metadata; only the XML metadata.

3.6 Data Model Fields

The DM fields (or UTYPEs) for the Spectrum DM are tabulated on the following pages. The field names are to be used as the UTYPE values in VOTABLE serializations.

We specify fields that are MANDATORY (MUST), RECOMMENDED (SHOULD), or OPTIONAL (MAY). MANDATORY fields are in bold. MANDATORY means that the document must provide a value; however, the value may be UNKNOWN (the value exists but is not known) or N/A (not applicable: for example, RA and DEC for a moving object or absolute time for a theory simulation). RECOMMENDED means that a data provider should try to fill the relevant fields if possible, but the document is still compliant if they are omitted. However, particular serializations (FITS, VOTABLE, etc) may amend these requirements by specifying default values for the serialization.

These requirements apply specifically for the spectrum application of this model; the VO may specify different MANDATORY/RECOMMENDED/OPTIONAL requirements for time series and other applications of the same model.

Some optional ID and UCD fields are allowed but are not listed below.

The fields are explained in more detail in the following sections.

Table 1: Spectrum metadata fields

Field	UCD1+	Meaning	Req	Default
Spectrum.DataModel		Version of this model	MAN	SPECTRUM-1.0
Spectrum.Type		Segment type	OPT	Spectrum
Spectrum.Length		Number of points	OPT	(must be derived)
Spectrum.TimeSI	time;arith.zp	SI factor and dimensions	REC	
Spectrum.SpectralSI	-	SI factor and dimensions	REC	
Spectrum.FluxSI	-	SI factor and dimensions	REC	
Spectrum.CoordSys.ID		ID string for coordinate system	OPT	
Spectrum.CoordSys.SpaceFrame.Name		ICRS or FK5	REC	ICRS
Spectrum.CoordSys.SpaceFrame.UCD	-	Space frame UCD	OPT	Char.SpatialAxis.UCD
Spectrum.CoordSys.SpaceFrame.RefPos		Origin of SpaceFrame	OPT	UNKNOWN
Spectrum.CoordSys.SpaceFrame.Equinox	time.equinox;pos.frame	Equinox	OPT	2000.0
Spectrum.CoordSys.TimeFrame.Name	time.scale	Timescale	OPT	TT
Spectrum.CoordSys.TimeFrame.UCD	-	Time frame UCD	OPT	time
Spectrum.CoordSys.TimeFrame.Zero	time;arith.zp	Zero point of timescale in MJD	OPT	0.0
Spectrum.CoordSys.TimeFrame.RefPos	time.scale	Times of photon arrival are at this location	OPT	TOPOCENTER
Spectrum.CoordSys.SpectralFrame.Name	-	Spectral frame name	OPT	(None)
Spectrum.CoordSys.SpectralFrame.UCD	-	Spectral frame UCD	OPT	Char.SpectralAxis.ucd
Spectrum.CoordSys.SpectralFrame.RefPos	?	Spectral frame origin	OPT	TOPOCENTER
Spectrum.CoordSys.SpectralFrame.Redshift		If restframe corrected	OPT	0.0
Spectrum.CoordSys.RedshiftFrame.Name	-	Redshift frame name	OPT	(None)
Spectrum.CoordSys.RedshiftFrame.DopplerDefinition	-	Opt, Radio, or Rel.	OPT	UNKNOWN
Spectrum.CoordSys.RedshiftFrame.RefPos	-	Redshift frame origin	OPT	UNKNOWN
Spectrum.Curation.Publisher	meta.curation	Publisher	MAN	
Spectrum.Curation.PublisherID	meta.ref.url;meta.curation	URI for VO Publisher	OPT	UNKNOWN
Spectrum.Curation.Date		Date curated dataset last modified	OPT	UNKNOWN
Spectrum.Curation.Version	meta.version;meta.curation	Version info	OPT	UNKNOWN
Spectrum.Curation.Rights		Restrictions: public, proprietary, mixed	REC	Public
Spectrum.Curation.Reference	meta.bib.bibcode	URL or Bibcode for documentation	REC	UNKNOWN
Spectrum.Curation.Contact.Name	meta.bib.author;meta.curation	Contact name	OPT	UNKNOWN
Spectrum.Curation.Contact.Email	meta.ref.url;meta.email	Contact email	OPT	UNKNOWN
Spectrum.Curation.PublisherDID	meta.ref.url;meta.curation	Publisher's ID for the dataset ID	OPT	DataID.DatasetID

Field	UCD1+	Meaning	Req	Default
Spectrum.DataID.Title	meta.title;meta.dataset	Dataset Title	OPT	None
Spectrum.DataID.Creator		VO Creator ID	OPT	UNKNOWN
Spectrum.DataID.Collection		Collection name(s)	OPT	None
Spectrum.DataID.DatasetID	meta.id;meta.dataset	IVOA Dataset ID	OPT	UNKNOWN
Spectrum.DataID.CreatorDID	meta.id	Creator's ID for the dataset	OPT	None
Spectrum.DataID.Date	time;meta.dataset	Data processing/creation date	OPT	UNKNOWN
Spectrum.DataID.Version	meta.version;meta.dataset	Version of dataset	OPT	UNKNOWN
Spectrum.DataID.Instrument	meta.id;instr	Instrument ID	OPT	UNKNOWN
Spectrum.DataID.Bandpass	instr.bandpass	Band, consistent with RSM Coverage.Spectral	OPT	UNKNOWN
Spectrum.DataID.CreationType		dataset creation type (archive, cutout,derived)	OPT	Archival
Spectrum.DataID.Logo	meta.ref.url	URL for creator logo	OPT	UNKNOWN
Spectrum.DataID.Contributor		Contributor (may be many)	OPT	UNKNOWN
Spectrum.DataID.DataSource		Original data type: survey, pointed, theory, artificial, composite	OPT	UNKNOWN
Spectrum.Derived.SNR	stat.snr	Signal-to-noise for spectrum	OPT	UNKNOWN
Spectrum.Derived.redshift.Value		Measured redshift for spectrum	OPT	UNKNOWN (may be undefined)
Spectrum.Derived.redshift.StatError	stat.error;src.redshift	Error on measured redshift	OPT	UNKNOWN
Spectrum.Derived.redshift.Confidence		Confidence value on redshift	OPT	UNKNOWN
Spectrum.Derived.VarAmpl	src.var.amplitude;stat.ratio	Variability amplitude as fraction of mean	OPT	UNKNOWN
Spectrum.Target.Name	meta.id;src	Target name	MAN	
Spectrum.Target.Description		Target descriptive text	OPT	UNKNOWN
Spectrum.Target.Class	src.class	Target or object class	OPT	UNKNOWN
Spectrum.Target.spectralClass	src.spType	Object spectral class	OPT	UNKNOWN
Spectrum.Target.redshift	src.redshift	Target redshift	OPT	UNKNOWN
Spectrum.Target.pos	pos.eq;src	Target RA and Dec	REC	UNKNOWN (may be variable)
Spectrum.Target.VarAmpl	src.var.amplitude	Target variability amplitude, typical	OPT	UNKNOWN

Field	UCD1+	Meaning	Req	Default
Spectrum.Char Axis definition fields				
Spectrum.Char.FluxAxis.name		name for flux	OPT	Flux
Spectrum.Char.FluxAxis.ucd		ucd for flux	MAN	
Spectrum.Char.FluxAxis.Unit		Unit for flux	MAN	
Spectrum.Char.SpectralAxis.name		name for spectral axis	OPT	SpectralCoord
Spectrum.Char.SpectralAxis.ucd		ucd for spectral coord	MAN	
Spectrum.Char.SpectralAxis.Unit		Unit for spectral coord	MAN	
Spectrum.Char.TimeAxis.name		name for time axis	OPT	Time
Spectrum.Char.TimeAxis.ucd		ucd for time	REC	time
Spectrum.Char.TimeAxis.Unit		Unit for time	REC	d
Spectrum.Char.SpatialAxis.name		name for spatial axis	OPT	Sky
Spectrum.Char.SpatialAxis.ucd		ucd for spectral coord	REC	pos.eq
Spectrum.Char.SpatialAxis.Unit		Unit for spectral coord	REC	deg
Spectrum.Char Coverage Fields				
Spectrum.Char.SpatialAxis.Coverage.Location.Value	pos.eq	Position, usually ICRS	MAN	
Spectrum.Char.SpatialAxis.Coverage.Bounds.Extent	instr.fov	Aperture angular diameter, deg	MAN	
Spectrum.Char.SpatialAxis.Coverage.Support.Area		Aperture region	REC	UNKNOWN
Spectrum.Char.SpatialAxis.Coverage.Support.Fill	stat. <i>fill</i> ;pos.eq	Sampling Filling factor	OPT	1.0
Spectrum.Char.TimeAxis.Coverage.Location.Value	time.epoch	Midpoint of exposure, on MJD scale, d	MAN	
Spectrum.Char.TimeAxis.Coverage.Bounds.Extent	time.expo	Total exposure time	MAN	
Spectrum.Char.TimeAxis.Coverage.Bounds.Start	time.expo.start	Start time	REC	UNKNOWN
Spectrum.Char.TimeAxis.Coverage.Bounds.Stop	time.expo.end	Stop time	REC	UNKNOWN
Spectrum.Char.TimeAxis.Coverage.Support.Fill	time;stat. <i>fill</i> ;time	Sampling Filling factor	OPT	UNKNOWN
Spectrum.Char.SpectralAxis.Coverage.Location.Value	instr.bandpass	Spectral coord value	MAN	
Spectrum.Char.SpectralAxis.Coverage.Bounds.Extent	instr.bandwidth	Width of spectrum in A or other spec. coord. (See text)	MAN	
Spectrum.Char.SpectralAxis.Coverage.Bounds.Start	em.*;stat.min	Start in spectral coordinate	MAN	
Spectrum.Char.SpectralAxis.Coverage.Bounds.Stop	em.*;stat.max	Stop in spectral coordinate	MAN	
Spectrum.Char.SpectralAxis.Coverage.Support.Fill	stat. <i>fill</i> ;em.*	Sampling Filling factor	OPT	1.0

Field	UCD1+	Meaning	Req	Default
Spectrum.Char Accuracy Fields - global				
Spectrum.Char.FluxAxis.Accuracy.StatError	phot.flux;em...;stat.error	error	REC	UNKNOWN
Spectrum.Char.FluxAxis.Accuracy.SysError	phot.flux;em...;stat.error.sys	Systematic error	REC	UNKNOWN
Spectrum.Char.FluxAxis.Accuracy.Calibration		Type of coord calibration	OPT	CALIBRATED
Spectrum.Char.SpectralAxis.Accuracy.BinSize	em.*;?	Wavelength bin size	OPT	UNKNOWN (may be undefined)
Spectrum.Char.SpectralAxis.Accuracy.StatError	em.*; stat.error	Spectral coord measurement error	REC	0
Spectrum.Char.SpectralAxis.Accuracy.SysError	em.*; stat.error	Spectral coord measurement error	REC	0
Spectrum.Char.SpectralAxis.Accuracy.Calibration		Type of coord calibration	OPT	CALIBRATED
Spectrum.Char.SpectralAxis.Resolution	instr.spectr.resolution	Spectral resolution FWHM	OPT	Accuracy.BinSize
Spectrum.Char.TimeAxis.Accuracy.BinSize	em.*;?	Time bin size	OPT	UNKNOWN (undefined)
Spectrum.Char.TimeAxis.Accuracy.StatError	time; stat.error	Time coord measurement statistical error	OPT	UNKNOWN
Spectrum.Char.TimeAxis.Accuracy.SysError	time; stat.error.sys	Time coord measurement systematic error	OPT	UNKNOWN
Spectrum.Char.TimeAxis.Accuracy.Calibration		Type of coord calibration	OPT	CALIBRATED
Spectrum.Char.TimeAxis.Resolution	TimeAxis.Accuracy.resolution	Temporal resolution FWHM	OPT	Accuracy.BinSize
Spectrum.Char.SpatialAxis.Accuracy.StatError	pos.eq;stat.error	Astrometric statistical error	OPT	UNKNOWN
Spectrum.Char.SpatialAxis.Accuracy.SysError	pos.eq;stat.error.sys	Systematic error	OPT	UNKNOWN
Spectrum.Char.SpatialAxis.Accuracy.Calibration		Type of coord calibration	OPT	CALIBRATED
Spectrum.Char.SpatialAxis.Resolution	instr.ang-res	Spatial resolution of data	OPT	UNKNOWN

Field	UCD1+	Meaning	Req	Default
Data Fields				
Spectrum.Data.FluxAxis.Value		Flux values for points	MAN	
Spectrum.Data.FluxAxis.ucd		ucd for flux	OPT	Char.FluxAxis.ucd
Spectrum.Data.FluxAxis.Unit		Unit for flux	OPT	Char.FluxAxis.Unit
Spectrum.Data.SpectralAxis.Value		Spectral coordinates for points	MAN	(Char.SpectralAxis.Location)
Spectrum.Data.SpectralAxis.ucd		ucd for spectral coord	OPT	Char.SpectralAxis.ucd
Spectrum.Data.SpectralAxis.Unit		Unit for spectral coord	OPT	Char.SpectralAxis.Unit
Spectrum.Data.TimeAxis.Value		Time coordinates for points	OPT	Char.TimeAxis.Location
Spectrum.Data.TimeAxis.ucd		ucd for time	OPT	Char.TimeAxis.ucd
Spectrum.Data.TimeAxis.Unit		Unit for time	OPT	Char.TimeAxis.Unit
Spectrum.Data.BackgroundModel.Value		Flux values for points	OPT	No background model
Spectrum.Data.BackgroundModel.ucd		ucd for background flux	OPT	Points.FluxAxis.ucd
Spectrum.Data.BackgroundModel.Unit		Unit for background flux	OPT	Points.FluxAxis.Unit

Field	UCD1+	Meaning	Req	Default
Accuracy Fields - per data point (default to corresponding Spectrum.Char values)				
Spectrum.Data.FluxAxis.Accuracy.StatError	phot.flux;em...;stat.error	symmetric error	OPT	(Char)
Spectrum.Data.FluxAxis.Accuracy.StatErrLow	phot.flux;em...;stat.error;stat.min	Lower error	OPT	(Char)
Spectrum.Data.FluxAxis.Accuracy.StatErrHigh	phot.flux;em...;stat.error;stat.max	Upper error	OPT	(Char)
Spectrum.Data.FluxAxis.Accuracy.SysError	phot.flux;em...;stat.error.sys	Systematic error	OPT	(Char)
Spectrum.Data.FluxAxis.Quality	meta.code.qual;phot.flux,em...	Quality mask	OPT	0
Spectrum.Data.FluxAxis.Quality.n		String value, for n = 0,1,2..; meaning of quality value	OPT	None
Spectrum.Data.SpectralAxis.Accuracy.BinSize	em.*;?	Wavelength bin size	OPT	(Char)
Spectrum.Data.SpectralAxis.Accuracy.BinLow	em.*;stat.min	Spectral coord bin lower end	OPT	Midpoint of value
Spectrum.Data.SpectralAxis.Accuracy.BinHigh	em.*;stat.max	Spectral coord bin upper end	OPT	Midpoint of value
Spectrum.Data.SpectralAxis.Accuracy.StatError	em.*; stat.error	Spectral coord measurement error	OPT	(Char)
Spectrum.Data.SpectralAxis.Accuracy.StatErrLow	em.*; stat.error; stat.min	Spectral coord measurement lower error	OPT	(Char)
Spectrum.Data.SpectralAxis.Accuracy.StatErrHigh	em.*; stat.error; stat.max	Spectral coord measurement upper error	OPT	(Char)
Spectrum.Data.SpectralAxis.Accuracy.SysError	em.*; stat.error	Spectral coord systematic error	OPT	(Char)
Spectrum.Data.SpectralAxis.Resolution	instr.spectr.resolution	Spectral resolution FWHM	OPT	(Char)
Spectrum.Data.TimeAxis.Accuracy.BinSize	em.*;?	Time bin size	OPT	(Char)
Spectrum.Data.TimeAxis.Accuracy.BinLow	em.*;stat.min	Time bin start	OPT	Midpoint of value
Spectrum.Data.TimeAxis.Accuracy.BinHigh	em.*;stat.max	Time bin stop	OPT	Midpoint of value
Spectrum.Data.TimeAxis.Accuracy.StatError	time; stat.error	Time coord measurement statistical error	OPT	(Char)
Spectrum.Data.TimeAxis.Accuracy.StatErrLow	time; stat.error, stat.min	Time coord measurement lower error	OPT	(Char)
Spectrum.Data.TimeAxis.Accuracy.StatErrHigh	time; stat.error, stat.max	Time coord measurement upper error	OPT	(Char)
Spectrum.Data.TimeAxis.Accuracy.SysError	time; stat.error.sys	Time coord measurement systematic error	OPT	(Char)
Spectrum.Data.TimeAxis.Resolution	time.resolution	Temporal resolution FWHM	OPT	(Char)
Spectrum.Data.BackgroundModel.Accuracy.StatError	phot.flux;em...;stat.error	Symmetric error	OPT	(Char)
Spectrum.Data.BackgroundModel.Accuracy.StatErrLow	phot.flux;em...;stat.error;stat.min	Lower error	OPT	(Char)
Spectrum.Data.BackgroundModel.Accuracy.StatErrHigh	phot.flux;em...;stat.error;stat.max	Upper error	OPT	(Char)
Spectrum.Data.BackgroundModel.Accuracy.SysError	phot.flux;em...;stat.error.sys	Systematic error	OPT	(Char)
Spectrum.Data.BackgroundModel.Quality	meta.code.qual;phot.flux,em...	Quality mask	OPT	0

4 Spectral data model Measurement objects

4.1 Spectral coordinate

Astronomers use a number of different spectral coordinates to label the electromagnetic spectrum. The cases enumerated by Greisen et al. (2006) are listed below with their UCDs.

REQUIRED: Exactly one `Spectrum.Char.SpectralAxis` field should be present, with units and one of the UCD values listed below. We distinguish between the VO data model field name (which might be used for VOTABLE UTYPE), the FITS WCS name (provided for comparison only), and the UCD1+ names.

Note 1: For this version, only the first four entries, Wavelength, Frequency, Energy, and spectral channel, should be used for interoperable transmission of data - implementations are not required to understand (convert) the other UCD values.

Note 2: For the velocity cases, the UCD uses a `spect.dopplerVeloc` tree rather than a `src.veloc` tree, because the velocity here is really a labelling of a spectral coordinate, and the link to the physical radial velocity of the different emission sources contributing to the spectrum is rather indirect.

Field	FITS WCS	UCD1+	Meaning	Units
PREFERRED CHOICES				
<code>SpectralAxis.ucd</code>	WAVE	<code>em.wl</code>	Wavelength	Angstrom, m
<code>SpectralAxis.ucd</code>	FREQ	<code>em.freq</code>	Frequency of photon	Hz
<code>SpectralAxis.ucd</code>	ENER	<code>em.energy</code>	Photon energy	erg, eV, J
<code>SpectralAxis.ucd</code>	-	<code>instr.pixel;em.wl</code>	Instrumental spectral bin	chan
ALTERNATIVE CHOICES				
<code>SpectralAxis.ucd</code>	WAVN	<code>em.wavenumber</code>	Wavenumber	$m^{*(-1)}$
<code>SpectralAxis.ucd</code>	AWAV	<code>em.wl; obs.atmos</code>	Air wavelength	Angstrom, m
<code>SpectralAxis.ucd</code>	WAVE-LOG	<code>em.wl</code>	Log wavelength	
<code>SpectralAxis.ucd</code>	FREQ-LOG	<code>em.freq</code>	Log frequency of photon	
<code>SpectralAxis.ucd</code>	ENER-LOG	<code>em.energy</code>	Log photon energy	
<code>SpectralAxis.ucd</code>	VELO	<code>spect.dopplerVeloc</code>	Apparent radial velocity	m/s
<code>SpectralAxis.ucd</code>	VRAD	<code>spect.dopplerVeloc.radio</code>	Radio velocity	m/s
<code>SpectralAxis.ucd</code>	VOPT	<code>spect.dopplerVeloc.opt</code>	Optical velocity	m/s
<code>SpectralAxis.ucd</code>	BETA	<code>spect.dopplerVeloc</code>	Velocity ($c=1$)	-

4.2 Flux (Spectral Intensity) Object

Two instances of the Flux object are supported: Flux and BackgroundModel. The Flux may be either the background-subtracted net flux or the total flux (the source+background), in the latter case hopefully with the BackgroundModel (see below). Net and total flux are distinguished by the 'src.net' UCD adjective.

For each of these cases, there are many slightly different physical quantities covered by the general concept of Flux; we distinguish them by their UCD.

Note in particular the distinction between the unit **count** (an instrumental value) and the unit **photon** (used in the photon number flux, i.e. the number of photons incident; photon number flux = energy flux divided by photon energy).

Note: The concept of the "nu L-nu" or "lambda L-lambda" flux, or equivalently the luminosity per logarithmic energy interval $L(\log \nu)$, is a distinct concept in the world of spectral energy distributions - and it's a different concept from the bolometric luminosity, which has the same units. The UCD board has not yet approved a UCD expressing this concept; we have to use `phys.luminosity` and infer the concept from the units. My solution for brightness temperature is also rather questionable.

Note: we propose the UCD `spect.continuum` to represent continuum flux.

Table 3: Flux Value options

Field	UCD1+	Meaning	Unit
FluxAxis.ucd	phot.flux.density;em.wl	Flux density per unit wave.	$\text{erg cm}^{-2}\text{s}^{-1}\text{\AA}^{-1}$ $\text{Wm}^{-2}\text{m}^{-1}$, $\text{keVcm}^{-2}\text{s}^{-1}\text{\AA}^{-1}$
FluxAxis.ucd	phot.flux.density;em.freq	Flux density per unit freq.	$\text{erg cm}^{-2}\text{s}^{-1}\text{Hz}^{-1}$ $\text{Jy, Wm}^{-2}\text{Hz}^{-1}$
FluxAxis.ucd	phot.flux.density;em.energy	Flux density per energy interval	$\text{keV cm}^{-2}\text{s}^{-1}\text{keV}^{-1}$
FluxAxis.ucd	phot.flux.density;em.energy; meta.number	Photons per unit area, time, energy	$\text{photon cm}^{-2}\text{s}^{-1}\text{keV}^{-1}$
FluxAxis.ucd	phot.flux.density;em.wl	Flux density per log wave interval (nu fnu)	Jy Hz
FluxAxis.ucd	phot.flux.density.sb;em.wl	Surface brightness per unit wavelength	$\text{erg cm}^{-2}\text{s}^{-1}\text{\AA}^{-1}\text{arcsec}^{-2}$
FluxAxis.ucd	phot.flux.density.sb;em.freq	Surface brightness per unit frequency	Jy sr^{-1}
FluxAxis.ucd	phot.count	Counts in spectral channel	count
FluxAxis.ucd	arith.rate;phot.count	Count rate in spectral channel	count/s
FluxAxis.ucd	arith.ratio; phot.flux.density	Flux ratio of two spectra	-
FluxAxis.ucd	phys.luminosity;em.wl	Luminosity per unit wave	$\text{ergs}^{-1}\text{\AA}^{-1}$, W/m
FluxAxis.ucd	phys.luminosity;em.freq	Luminosity per unit freq	$\text{ergs}^{-1}\text{Hz}^{-1}$, W/Hz
FluxAxis.ucd	phys.luminosity;em.energy	Luminosity per unit energy	$\text{ergs}^{-1}\text{keV}^{-1}$
FluxAxis.ucd	phys. luminosity;em.energy	Luminosity per log frequency	erg s^{-1} , W
FluxAxis.ucd	phys.energDensity	Radiation energy density per unit volume, per unit wave etc.	erg cm^{-3} , W m^{-3}
FluxAxis.ucd	phot.flux.density;em.wl; phys.polarization	Polarized flux per unit wavelength	$\text{erg cm}^{-2}\text{s}^{-1}\text{\AA}^{-1}\text{arcsec}^{-2}$
FluxAxis.ucd	phys.polarization	Polarized fraction vs spectral coord	(dimensionless)
FluxAxis.ucd	phys.luminosity; phys.angArea;em.wl	Flux per unit solid angle (at source)	$\text{erg cm}^{-2}\text{s}^{-1}\text{sr}^{-1}\text{\AA}^{-1}$
FluxAxis.ucd	instr.antenna-temp	Antenna temperature	K
FluxAxis.ucd	phot.flux.density; phys.temperature	Brightness temperature	K
FluxAxis.ucd	phot.mag	Magnitude in defined band	mag
FluxAxis.ucd	phot.mag	AB (spectrophotometric) magnitude	mag
FluxAxis.ucd	phot.flux.density;instr.beam	Flux per resolution element (e.g. Jy/beam)	Jy/beam
FluxAxis.ucd	phot.mag.sb	Surface brightness in magnitudes	mag arcsec^{-2}
FluxAxis.ucd	phys.transmission	Filter transmission	(dimensionless)
FluxAxis.ucd	phys.area;phys.transmission	Effective area	cm^{*-2}
FluxAxis.ucd	phot.flux.density;em.wl; <i>spect.continuum</i>	Continuum only	$\text{erg cm}^{-2}\text{s}^{-1}\text{\AA}^{-1}\text{arcsec}^{-2}$

4.3 BackgroundModel Object

We optionally allow a BackgroundModel value for each Flux value. We define $\text{NetFlux} = \text{TotalFlux} - \text{BackgroundModel}$. The name BackgroundModel, rather than Background, reminds us that it is an estimate: often, the BackgroundModel will be generated by taking a flux measurement at another location and rescaling it for any difference in exposure time or extraction

aperture.

The BackgroundModel array is required to have the same UCD and units as the Flux array. It represents a model for the expected flux values if the Target had zero flux.

OPTIONAL: There may be at most one BackgroundModel.Value field present. It must have the same UCD as the Flux.

4.4 Time coordinate

For data with a time-series component, whether regularly sampled or sparse photometry points, the time coordinate is given by an elapsed time in some physical units (e.g. seconds or days) relative to a reference time.

This reference time is given in MJD as the field Spectrum.Char.TimeAxis.Coverage.Location, as described in the Characterization section. For a simple spectrum with no time-resolved data, this is the time of the observation.

For time-resolved data, the time coordinate refers to the midpoint of the sample interval. See the Space-Time Coordinates document for details of time coordinate complexity.

The time unit is specified by a string, and the only valid values for this unit are 's' (seconds) and 'd' (days).

4.5 Position coordinate

In general we may consider position coordinates as part of the measurement (and possibly varying from point to point), but this capability is not included in the current document. The (celestial) position of the aperture for the spectrum is given in the spatial Spectrum.Char.SpatialAxis.Coverage.Location field. The Spectrum.Char.SpatialAxis.Coverage.Location.Value field is in the coordinates of CoordSys.SpaceFrame. The default is ICRS RA,Dec in decimal degrees.

4.6 Accuracy Fields

We include accuracy models for both the coordinates (spectral, spatial and temporal) and the fluxes. The accuracy can appear in two places: in the global characterization, where it represents typical accuracy for the dataset, and in the data points themselves, providing a way to provide per-data-point errors. All the Accuracy fields are optional, both in the per-data-point fields and in the Characterization instances; the per-data fields default to the values in Characterization.

4.6.1 Coordinate bins

We express the bandpass for each spectral bin as a low and high value for the spectral coordinate, or as a width. The same is done for photometry points, which amounts to approximating a filter by a rectangular bandpass. Time bins are also given as low and high values or as a width. Note that the width values are suitable for Spectrum.Char (the global accuracy) while the bin low/high values only have meaning for Spectrum.Data (the per-data-point values).

Only one of BinSize, or both BinLow and BinHigh, must be present (possibly as a header parameter implying a constant value for each flux point). If absent, the bin limits are assumed to be halfway between the coordinate values and bounded by the range given in Char.*.Coverage.Extent.

4.6.2 Uncertainties

In addition to the binning, we allow the model to express uncertainties (which may be larger than the bin width), both statistical and systematic. We allow one or two-sided statistical errors but only one-sided systematic errors. You can specify StatErr, or StatErrHigh/StatErrLow, but not both.

For position we have a single statistical error - a two-sided error doesn't make sense for a 2D coordinate. Eventually we may want a full error ellipse, but this is too complicated for the present model.

We also use a very simple error model for the fluxes: we include plus and minus flux errors, and a quality flag. The errors are understood as 1 sigma gaussian errors which are uncorrelated for different points in the spectrum. If the data provider has only upper limit information, it should be represented by setting the flux value and the lower error value equal to the limit, and the upper error value equal to zero (e.g. 5 (+0,-5)). In general applications may choose to render measurements as upper limits if the flux value is less than some multiple (e.g. 3) of the lower error. We also allow a systematic error value, assumed constant across a given spectrum and fully correlated (so that, e.g. it does not enter into estimating spectral slopes).

CLARIFICATION: the two-sided errors StatErrLow and StatErrHigh are the plus/minus ERRORS, not the (value+error, value-error). In other words, if Value = 10 and there is a symmetric uncertainty of 3, the ErrorLow and ErrorHigh are both +3.0, and NOT 7.0, 13.0. This is different from the sampling description BinLow and BinHigh, which give the VALUES at the low and high end of the bin. Thus if the central wavelength of the bin is 4200.0, and the bin size is 10, then the BinLow and BinHigh values are 4195.0, 4205.0 and NOT 10.0, 10.0. Note that because of this, 0.0 is NOT an acceptable default for BinLow and BinHigh, while it IS acceptable (albeit unlikely) for StatErrLow and StatErrHigh.

The StatErrLow, StatErrHigh, SysError fields for SpectralCoord, Time, Sky and Flux are optional; however, omitting these fields indicates that the errors are unknown. Data providers are STRONGLY encouraged to provide explicit error measures whenever possible.

4.6.3 Resolution

We also include a trivial resolution model: a single number nominally representing a FWHM spectral or time resolution expressed in the same units as the spectral or time coordinate. The default is to assume that the resolution is equal to the BinSize if defined. The spatial (sky) resolution may be useful to know if it exceeds the aperture size; the default is to assume it is equal to the aperture size.

4.6.4 Quality

The Quality model represents quality by an integer, with the following meanings: 0 is good data, 1 is data which is bad for an unspecified reason (e.g., no data in the sample interval), and other positive integers greater than 1 may be used to flag data which is bad or dubious for specific reasons.

The data provider may also define scalar string-valued metadata fields Quality.2, Quality.3... to define specific quality flags on a per-spectrum basis. Bitmasks, used in some archives such as SDSS, should be remapped to such independent Quality fields.

Quality defaults to zero, i.e. good data.

4.6.5 Calibration

We also introduce a Calibration field which can have the values ABSOLUTE, RELATIVE or UNCALIBRATED. This is expected to be particularly useful to describe the flux. ABSOLUTE indicates that the values in the data are expected to be correct within the given uncertainty. RELATIVE indicates that although an unknown systematic error is present, the ratio of any two values will be correct. UNCALIBRATED indicates that although the values reflect a measurement of the given UCD, they are modified by an unspecified coordinate-dependent correction. Such values may be useful in the case of a spectrum with ABSOLUTE calibration on the wavelengths but UNCALIBRATED fluxes; the wavelengths of discontinuous features such as spectral lines can be measured on the assumption that the missing calibration function has no sharp discontinuities in the region of interest.

The Calibration fields are present in the global characterization accuracy but not in the per-data-point accuracy.

5 Associated Metadata Fields

Most of the associated metadata are generic observational metadata that can be applied to future data models, and are not specific to spectra.

5.1 CoordSys Fields

The CoordSys object is a simplified instance of the STC CoordSystem object. For XML serializations, it can be replaced by an actual STC CoordSystem instance.

CoordSys consists of 1 or more CoordFrame objects, each of which defines the coordinates for a particular axis. The CoordSys has an overall ID string, which is user-defined and arbitrary. Each CoordFrame also has a type, a UCD and a ReferencePosition; the Reference Position gives the origin of the coordinate system (and thus also its rest frame).

For the space, time, and spectral axes we define specialized CoordFrames for convenience: SpaceFrame, TimeFrame and SpectralFrame. The CoordFrame names (types) for SpaceFrame and TimeFrame must be from a controlled list; for other frames, the type is an arbitrary string.

Token	Meaning	Note
UNKNOWN	Unknown origin	
RELOCATABLE	Relative origin	Suitable for simulations
CUSTOM	Origin specified wrt another system	
TOPOCENTER	Location of the observing device	(telescope)
BARYCENTER	Solar system barycenter	
HELIOCENTER	Center of the Sun	
GEOCENTER	Center of the Earth	
EMBARYCENTER	Earth-Moon barycenter	
MOON	Center of the Moon	
MERCURY	Center of Mercury	
VENUS	Center of Venus	
MARS	Center of Mars	
JUPITER	Center of Jupiter	
SATURN	Center of Saturn	
URANUS	Center of Uranus	
NEPTUNE	Center of Neptune	
PLUTO	Center of Pluto	
LSRK	Kinematic local standard of rest	Redshift frame only
LSRD	Dynamic local standard of rest	Redshift frame only
GALACTIC_CENTER	Center of the Galaxy	
LOCAL_GROUP_CENTER	Barycenter of the Local Group	

Table 4: Allowed values for CoordFrame.ReferencePosition

5.1.1 SpaceFrame

The SpaceFrame has an optional Equinox attribute which is used if the frame name is FK4 or FK5. The allowed frame names for SpaceFrame are listed below.

5.1.2 TimeFrame

The TimeFrame is defined by the frame name and the ReferencePosition. Allowed values of the name are given below.

One standard reference time in astronomy is the origin of Julian Day Number on the TT (Terrestrial Time) timescale, BC 4713 Nov 24 at 11:59:27.81 (Gregorian). Using TT is preferable to UTC because it does not contain leap seconds, so the elapsed time in days is just equal to the difference in JD values.

The ISO-8601 calendar format standard does not support dates before AD 1, so cannot express this reference time. Therefore, it is not a suitable format for internal representations of such reference times. However, non-default choices of reference time may be specified in external serializations by a date in ISO-8601 format, e.g. "2004-11-30T11:59:00.01".

In this version of the model we require use of MJD as the time type for absolute times. (ISO dates and JD are other possibilities covered by the STC document). Relative times in a time series may be in other units, relative to the TimeFrame.Zero value.

(Note that in the FITS serialization, the MJDREF keyword allows definition of reference times in decimal days relative to MJD 0.0 = JD 2400000.5.)

5.1.3 SpectralFrame

The spectral frame is defined by its ReferencePosition. Once the choice of wavelength versus frequency or energy has been made, the only free parameter is the location at which the spectrum would have the given spectral coordinates. For directly observed data this is the topocenter (location of the observation); spectra may be velocity-corrected to a given velocity frame, which may be defined by the location which is at rest in that velocity frame (e.g. the heliocenter). Strictly, the correction may not be just a velocity shift, but any kind of spectral shift including e.g. gravitational redshifts; it is still true that such a shift corresponds to a location (e.g. surface of a white dwarf star) that can be quoted as a reference position.

Since the frame is defined by its ReferencePosition, the frame name is not important, and will not be significant to software. We suggest that it may be filled by the name of the spectral coordinate, using FITS names such as 'WAVE', 'FREQ' or 'ENER'.

The spectral frame has an optional Redshift attribute to specify a rest frame; it is used only if the the frame's ReferencePosition is "CUSTOM". This redshift is measured in dimensionless units, defined as $\Delta\lambda/\lambda$ and may be negative. No specific interpretation of the shift as a cosmological or velocity shift effect is implied.

5.1.4 RedshiftFrame (also the Velocity frame)

When you convert the spectral coordinate to velocity or redshift (relative to some assumed rest-frame spectral feature) you need to record some other metadata. Our field name containing this metadata is RedshiftFrame, but we emphasize that the name redshift does not imply that blueshifts are excluded, merely that, in both galactic and extragalactic astronomy, when a shift is interpreted as a velocity a positive value indicates a shift to the red. The concept of Redshift frame includes both cosmological and local Doppler velocities.

Note that you only use RedshiftFrame if you're measuring things in velocities; a rest-frame spectrum of a redshifted quasar whose spectral axis is in Angstroms will be described by a SpectralFrame. The reason we have BOTH SpectralFrame and RedshiftFrame is to support certain data products, particularly used in spectral line radioastronomy, in which a spectrum (possibly obtained in piecewise spectral regions) is refactored into a set of separate spectral

segments centered on different spectral lines; each segment is assigned a velocity axis centered on that line (and the same pixel from the original spectrum can appear in multiple segments each with a different velocity coordinate); you then consider the data as a 2D array with a spectral axis (indexing the segments) and a velocity axis (for each segment/spectral line).

Other coordinate system information needed for velocity spectral coordinates include the observation-fixed spectral frame, the observatory location, the source redshift, and the velocity zero point (in Greisen et al, SSYSOBS, OBSGEO, VELOSYS, RESTFRQ/RESTWAV). However, we omit these in the current model. The only metadata we provide is the Doppler Definition - optical, radio or pseudo-relativistic.

5.1.5 STC

Notes on compatibility with, and differences from, STC 1.0:

- We add the extra Redshift attribute in the SpectralFrame, instead of the more complex CustomReferencePosition approach used in STC.
- In STC's XML serialization, the frame types and reference positions are enumerated defined elements. Here they are strings, and we require that the frame name is the frame type.
- We don't explicitly include the coordinate flavor.

OPTIONAL: All CoordSys values are optional , but data providers should take special care to check whether or not the defaults are appropriate for their data. The implications of the defaults are:

- Positions are given in ICRS RA,Dec in degrees and are heliocentric values (i.e. corrected for annual parallax and aberration, as normally found in source catalogs).
- Times are given in MJD days and represent times of photon arrival at the telescope.
- Spectral coordinates are as observed at the telescope, and not corrected for redshift, the motion of the Earth, etc.

Token	Meaning	Parameter(s)
UNKNOWN	Unknown frame	
CUSTOM	Custom frame	Pole, axis
AZ_EL	Azimuth and elevation	
BODY	Generic body (eg planet)	
ICRS	The ICRS frame	
FK4	FK4	Equinox
FK5	FK5	Equinox
ECLIPTIC	Ecliptic l,b	Equinox
GALACTIC_I	Old galactic LI,BI	
GALACTIC_II	Galactic LII,BII	
SUPER.GALACTIC	SGL, SGB	
MAG	Geomagnetic ref frame	
GSE	Geocentric Solar Ecliptic	
GSM	Geocentric Solar Magnetic	
SM	Solar Magnetic	
HGC	Heliographic	
HEE	Heliocentric Earth Ecliptic	
HEEQ	Heliocentric Earth Equatorial	
HCI	Heliocentric Inertial	
HCD	Heliocentric of Date	
GEO_C	Geocentric corotating	
GEO_D	Geodetic ref frame	Spheroid
MERCURY_C	Corotating planetocentric	
VENUS_C	Corotating planetocentric	
LUNA_C	Corotating planetocentric	
MARS_C	Corotating planetocentric	
JUPITER_C.III	Corotating planetocentric	
SATURN_C.III	Corotating planetocentric	
URANUS_C.III	Corotating planetocentric	
NEPTUNE_C.III	Corotating planetocentric	
PLUTO_C	Corotating planetocentric	
MERCURY_G	Corotating planetographic	
VENUS_G	Corotating planetographic	
LUNA_G	Corotating planetographic	
MARS_G	Corotating planetographic	
JUPITER_G.III	Corotating planetographic	
SATURN_G.III	Corotating planetographic	
URANUS_G.III	Corotating planetographic	
NEPTUNE_G.III	Corotating planetographic	
PLUTO_G	Corotating planetographic	

Table 5: Allowed values for CoordSys.SpaceFrame.Name

Token	Meaning	Note
LOCAL	Relocatable (simulation) time	
TT	Terrestrial Time	
UTC	Coordinated Universal Time	
ET	Ephemeris Time	
TDB	Barycentric dynamical time	
TCG	Terrestrial Coordinate Time	
TCB	Barycentric Coordinate Time	
TAI	International Atomic Time	
LST	Local Sidereal Time	

Table 6: Allowed values for CoordSys.TimeFrame.Name

5.2 Characterization

The Characterization metadata in this document are based on, but not identical to, the IVOA Note published by the Characterization data model subgroup in March 2006. The Characterization model has a set of CharacterizationAxis objects. Each CharacterizationAxis has an AxisFrame describing the axis, a Coverage describing the scope of the data, and optionally a Resolution and a Sampling object. We do not include Sampling in the current Spectrum model.

The CharacterizationAxis is identified by its UCD attribute. Spectrum instances should have Spatial, Time and Spectral characterization axes as well as FluxAxis. To simplify things for the common axes, we define SpatialAxis, SpectralAxis, TimeAxis objects as special cases of CharacterizationAxis.

The CoordSystem element in CharacterizationAxis is there for compatibility with the Characterization document and, if present, should be a simple reference to the main Spectrum CoordSystem.

The Characterization fields will have a constant value for a given spectrum.

Note: In the SSA protocol/query response, we will restrict the Char units to meters (spectral coord), days (time coord), and decimal degrees (spatial), for simplicity and consistency with other parameters. In the model and the returned serializations, the units may be as described elsewhere in this document.

5.2.1 Coverage Fields

The coverage fields will have a constant value for a given spectrum. They describe the region of space, time and spectrum from which the data were taken. In the Characterization model, we define progressively more accurate descriptions of this region: Location gives a single characteristic point, Bounds gives a range within which the data lies, and Support gives the detailed spatial field of view footprint, on/off time ranges (including gaps) and spectral ranges. (A fourth level not yet supported, Sensitivity, will provide detailed depth information: exposure map, time sensitivity variation, spectral transmission curve).

There is a field for giving the effective exposure time (useful for selecting among multiple spectra from the same instrument). The aperture is important to determine what part of an extended object is contributing to the spectrum; we allow a simple aperture description (Char.SpatialAxis.Coverage.Bounds.Extent) consisting of a single number representing the aperture size in decimal degrees. (For a slit spectrum, the effective aperture on the sky is usually the slit width). A full region polygon is allowed in the Area field.

The units of the spectral Coverage.Bounds.Extent (or Coverage.Bounds.Start/Stop) and Coverage.Support should be the same as those of SpectralCoord.

For time, the Coverage.Bounds.Start/Stop is a pair of values giving the start and stop time. We can also have Coverage.Support with a whole array of start-stop pairs indicating data accumulated over a series of intervals.

The SpatialAxis.Coverage.Location and SpatialAxis.Coverage.Bounds.Extent, TimeAxis.Coverage.Location are required, as are either TimeAxis.Coverage.Bounds.Extent or TimeAxis.Coverage.Bounds.Start and Stop. If Extent is provided, Start and Stop are defined to be (Location - 0.5* Extent, Location +0.5*Extent).

The spectral equivalents, SpectralAxis.Coverage.Location and SpectralAxis.Bounds.Start/Stop, are also required in the model; serializations may decide to omit them since they are easily derived from the data.

5.2.2 Region definitions

In the optional Char.SpatialAxis.Coverage.Support.Area we describe the detailed aperture shape in absolute coords on the sky. However, we don't allow a full STC region description. Our simplified region model allows for (1) a circle and (2) a polygon in a string representation: either

```
circle x0 y0 r
```

or

```
polygon x1 y1 x2 y2 x3 y3 ...
```

for example

```
circle 233.70 -13.32 0.00043  
polygon 233.70 -13.32 233.71 -13.30 ...
```

where the positions and radii are required to be in degrees, in the coordinate system defined by CoordSys.

5.3 Derived Data Fields

The Derived Data object has useful, and optional, summary information about the spectrum. For now, we include the option of adding signal-to-noise and variability indicators and a measurement of the redshift.

5.3.1 Signal-to-noise

The signal-to-noise is provided mainly as a way for searches to exclude data whose quality is insufficient for a particular study. Data providers may use their own definition, as we do not prescribe a uniform method to calculate it. A suitable method, used by the STScI MAST group, is define the noise by the median difference between adjacent independent flux values in the spectrum. (The MAST definition multiplies this noise value by a 1.048 correction factor for precise applications). This method describes the high-spectral-frequency noise but does not take into account low-spectral-frequency background 'noise'; projects which are background dominated may wish to include this in the noise definition. Furthermore most spectra vary in SNR across their waveband; users should therefore only use this single SNR as a crude selection parameter.

5.3.2 Redshift measurement model

One common piece of derived data for a spectrum is the source redshift. We provide fields for both the redshift measured value and statistical error. As above, we define the redshift to be $\Delta\lambda/\lambda$ and it may be positive or negative. The DerivedData field represents a measurement of the redshift from the data; a field in the Target object is available to store the redshift of the source as known from other means.

We add a further optional measure of accuracy, the Confidence, which expresses a probability between 0 and 1 that the quoted errors do apply. This measure is used in the Sloan spectral service to provide a way of describing the estimated probability that the redshift is completely in error because the lines have been misidentified. Its default value is 1.0.

In general, such a Confidence could be useful for any measurement where the error probability distribution has multiple peaks in parameter space, and could later be added to the standard Accuracy model.

Note that there are two other redshifts in our model: the Target redshift, a useful piece of metadata particularly for extragalactic objects, considered as an externally known property of the target (and so defined even if no lines are visible in the spectrum); and the SpectralFrame redshift, used only if a "rest frame" spectrum is presented and representing the assumed redshift used to shift the spectrum.

5.3.3 Variability amplitude

The variability amplitude field allows data providers to supply a characteristic amplitude (a precise value is not required). It is dimensionless; a value of 0.2 implies a 20 percent variation around the mean value.

5.4 Curation model

The Curation is an object consistent with the Curation information in the document "Resource Metadata for the Virtual Observatory Version 1.01", although some of the fields from RM curation have been moved to the DataID object, as discussed in the SSAP protocol document.

In Curation, we have added a Reference field for a bibliographic reference, Rights field (same as Resource.Rights) for public/proprietary, and PublisherDID for a publisher-specified IVORN to the data. The Curation.PublisherDID is the same as the Resource Metadata V1.10 Resource.Identifier.

Version is provided by the publisher or creator and may be any string.

Curation.Publisher is REQUIRED. All other fields are optional.

5.5 Data Identification model

The Data Identification model gives the dataset ID for a particular spectrum, and its membership of larger collections. All DataID fields are optional.

There are three dataset identifiers in the model: one under Curation and two here. All of them are ivo: URIs as specified by the IVOA.

The DataID.CreatorDID is the dataset ID defined internally by the creator and may be entirely different from the DatasetID described above. It is used to identify a particular original exposure in an archive and will not necessarily change even if the VO object in question is a cutout or is otherwise further processed.

The Curation.PublisherDID is a dataset ID defined by a publisher of the data.

The DataID.DatasetID may be the same as Curation.PublisherDID; for this field we recommend a journal-based URI such as the IVOA/ADEC/ADS dataset identifier. By agreement between the AAS journals, the ADS and the ADEC (NASA data centers), dataset identifiers, described in <http://vo.ads.harvard.edu/dv/>, will be used to link journal articles back to the archival datasets containing the relevant observational data. If analogous but independent systems of URI designation are later adopted by other centers (e.g. by European journals) and accepted by IVOA, will be suitable in this field.

The PublisherDID and DatasetID are currently redundant with each other; we reserve the latter for use in future IVOA standards.

We introduce the concept of an dataset creation type, which can have one of the following values, described in more detail in the SSAP protocol document.

- Archival, indicating that it is one of a collection of datasets (in this case spectra) generated in a systematic, homogeneous way and stored statically (or at least versioned). It will be possible to regenerate this dataset at a later date. The remaining types imply on-the-fly manipulation.
- Cutout, indicating that the dataset was created 'on-the-fly', by subsetting, but not by modifying values.
- Filtered, which may involve excluding data prior to binning into samples, also on the fly
- Mosaic, combining multiple original datasets on the fly
- spectral extraction, e.g. from a spectral data cube;
- catalog extraction [Not relevant for Spectrum model].

The dataset is associated with one or more Collections (instrument name, survey name, etc.) indicating some degree of compatibility with other datasets sharing the same Collection properties. Examples of possible Collection values are: "WFC", "Sloan", "BFS Spectrograph", "MSX Galactic Plane Survey".

We also include a `DataID.Bandpass`, which is a string describing the spectral range. It can be one of the strings in `Resource-Service-Metadata's Spectral.Coverage` (e.g. "Optical") or `Spectral.Coverage.Bandpass` (e.g. "B"). At the moment there is no fixed list of values for the RSM `Spectral.Coverage.Bandpass`.

For `DataSource`, see the SSAP protocol document.

5.6 Target model

In spectral data it is particularly important to be able to specify the target of the observation, which may be an astronomical source or some other target (calibration, diffuse background, etc.). By explicitly including a target model we can not only facilitate searches on particular types of target but also support archives of model spectra for which the Coverage fields may not be relevant. The `Target.Name` field is required; all other Target fields are optional.

The `Target.pos` field gives a nominal RA and Dec for the target, for example the catalog position of the source; the `Coverage.Location` fields in the spectrum indicate the actual telescope pointing position for that spectrum. (An SED might have a single Target object with a known position, but many Spectrum objects with slightly different telescope pointings). Similarly, the `Target.redshift` is the assumed actual redshift of the astronomical object, if applicable (again, usually from a catalog, NED, etc.), while the redshifts in the `DerivedData` objects in the spectrum (segment) indicates a redshift measured from that spectrum. The `Target.redshift` is normally used to store the cosmological redshift of extragalactic objects, although it may also be used to store the observed redshift of Galactic sources if that information is felt by the data provider to be useful.

At the moment there is no international standard list of valid values for Target class and spectral class. Nevertheless an initial deployment of the VO would gain some benefit from using archive-specific classes, and provide a framework for converging on a standard list.

5.7 Spectrum top level object

The Spectrum object contains the Data object with the actual data; the Target and `DerivedData` objects; and the standard dataset metadata of `CoordSys`, `Characterization`, `Curation` and `DataID`.

We also add a CustomParams field to allow for propagation of unmodelled application-specific metadata.

In addition, we add an SIDim field for each axis giving the SI units of the values in the Osuna-Salgado dimensional format.

In spectral associations (such as SED applications), the Spectrum model is reused for both Spectrum and TimeSeries and is renamed Segment. The Spectrum object is expected to be generalized to a higher level Dataset object.

Each Spectrum (or Segment) may have a Length attribute giving the number of flux points in the data (in some serializations this value is deduced from the size of the data arrays, while in others it is made explicit).

Each Spectrum (or Segment) may also have a Type attribute indicating whether the data is intended as a TimeSeries (data are same spectral coord, varying times), Photometry (data are different spectral coords with irregular gaps), Spectrum (data are different spectral coords in contiguous bins), or Mixed (some mixture of the above).

This attribute is optional and defaults to Spectrum.

Segments are discussed in more detail in the Spectral Associations document which describes SEDs and other groupings.

6 Relationship to general VO data models

The Spectrum model involves objects addressed by the proposed VO Observation and Quantity data models. Although these models have not yet been fully worked out, we may note that a single Spectrum maps to the Observation model, which will include the Curation and Characterization objects. The Flux and the spectral coordinate entries together with their associated errors and quality will be special cases of the Quantity model, as will the simpler individual parameters. The field structure presented here is consistent with current drafts of the models.

6.1 Extensibility

The model and serializations defined in this document are extensible in the following sense:

- Future versions of the abstract (UML) model may add attributes or fields, and may deprecate the 'optional' property of existing fields.
- The Characterization object may have extra 'generic' axes, but these are not considered to be part of the Spectrum model. See Characterization specification for more.
- For the FITS serialization, implementors may add arbitrary additional keywords or table columns; readers must be able to handle files containing extra keywords and columns, and are encouraged to propagate such extra information when copying files. This permits local conventions to be layered on the basic definition.
- For the VOTABLE serialization, implementors may add arbitrary additional GROUP, PARAM or FIELDref/FIELD elements, with the restriction that the layering of existing elements should not be changed. (e.g. within the spectrum:DataID GROUP one may add a new GROUP containing newly defined PARAMs, but one may not move the existing Title PARAM inside the new group because that would change its indentation level). Readers must be able to handle files containing extra elements and are encouraged to propagate such extra information when copying files. This permits local conventions to be layered on the basic definition.
- For the XML object-based serialization, the CustomParams element at the top level of Spectrum is intended to allow extensibility and is equivalent to the ability to add a new GROUP at the top level. Future versions of the schema could use type extension to back-compatibly include the current schema as a special case, but apart from the CustomParams we have not currently provided for local extensibility within the current schema. (We could improve the schema by allowing the Group element to have arbitrary extra Param elements.)

References

Greisen, EW, Valdes F G, Calabretta M R and Allen S L 2006, A&A 446, 747.

Hanisch, R., (ed)., Resource Metadata for the VO, Version 1.01, 2004 Apr 26.
<http://www.ivoa.net/Documents/latest/RM.html>

Derriere, S. et al (eds.), UCD, Moving to UCD 1+, 2004 Oct 26.
<http://www.ivoa.net/Documents/latest/UCD.html>

Part 2: XML schema serialization

7 XML schema serialization

7.1 XML schema

In the following XML schema, we implement the model fairly directly.

Within a spectrum the data points are kept together in objects called Point, with their general structure specified in a Fields object. Also, we have included a CustomParams element to allow site-specific metadata to be added.

The Coverage.Location fields have been collapsed to simple values rather than SEDCoord elements; this should perhaps be extended in a future version.

The Flux object is defined as an example of a more general SEDQuantity object, which is also used for the Sloan spectral service's redshift information.

A SED aggregation model is also included in the schema, as the top level element. This may be ignored until the SED model has been approved by IVOA.

The old Fields element is subsumed in the Char.Axis elements.

```

<?xml version="1.0" encoding="utf-8"?>
<xs:schema xmlns="http://www.ivoa.net/xml/Spectrum/Spectrum-0.98b.xsd"
xmlns:xs="http://www.w3.org/2001/XMLSchema"
xmlns:jxb="http://java.sun.com/xml/ns/jaxb"
xmlns:xlink="http://www.w3.org/1999/xlink"
targetNamespace="http://www.ivoa.net/xml/Spectrum/Spectrum-0.98b.xsd"
elementFormDefault="qualified" jxb:version="1.0">
<xs:import namespace="http://www.w3.org/1999/xlink" schemaLocation="http://www.ivoa.net/xml/Xlink/xlink.xsd"/>

<!-- A single segment corresponding to a spectrum or single point -->

<xs:element name="BaseSegment" type="segmentType"/>
<xs:element name="Spectrum" type="spectrumType" substitutionGroup="BaseSegment"/>
<xs:element name="Segment" type="spectrumType" substitutionGroup="BaseSegment"/>
<xs:element name="TimeSeries" type="timeSeriesType" substitutionGroup="BaseSegment"/>

<xs:complexType name="spectrumType">
<xs:complexContent mixed="false">
<xs:extension base="segmentType"/>
</xs:complexContent>
</xs:complexType>
<xs:complexType name="timeSeriesType">
<xs:complexContent mixed="false">
<xs:extension base="segmentType"/>
</xs:complexContent>
</xs:complexType>

<xs:complexType name="segmentType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="Target" type="targetType" />
<xs:element minOccurs="0" maxOccurs="1" ref="Data"/>
<xs:element minOccurs="0" maxOccurs="1" name="Char" type="characterizationType" />
<xs:element minOccurs="0" maxOccurs="1" name="CoordSys" type="coordSysType" />
<xs:element minOccurs="0" maxOccurs="1" name="Curation" type="curationType" />
<xs:element minOccurs="0" maxOccurs="1" name="DataID" type="dataIDType" />
<xs:element minOccurs="0" maxOccurs="1" name="Derived" type="derivedDataType" />
<xs:element minOccurs="0" maxOccurs="1" name="CustomParams" type="arrayOfParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Type" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Length" type="intParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="TimeSI" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="SpectralSI" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="FluxSI" type="textParamType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

```

```

<!-- The top level element: an SED with one target and many segments -->
<xs:element name="SED" nillable="true" type="sedType" />

<xs:complexType name="sedType">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="Date" type="timeParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Target" type="targetType" />
<xs:element minOccurs="0" maxOccurs="1" name="CustomParams" type="arrayOfParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Type" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="NSegments" type="intParamType" />
<xs:element minOccurs="0" maxOccurs="unbounded" ref="BaseSegment"/>
<xs:element minOccurs="0" maxOccurs="1" name="Creator" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="CreatorDID" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="SpectralMinWavelength" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="SpectralMaxWavelength" type="doubleParamType" />
</xs:sequence>
</xs:complexType>

<!-- Define the UCDS etc for the SED coordinate and the flux coordinate,
and include a global to specify accuracy etc which happens to be
constant for the entire segment (note that in SEDCoord,
value has minOccurs=0 so it can be omitted) -->
<!-- A single SEDCoord (time or spectral coord) value, or two values if it is spatial. -->
<xs:complexType name="sedCoordType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="2" name="Value" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Accuracy" type="accuracyType" />
<xs:element minOccurs="0" maxOccurs="1" name="Resolution" type="doubleParamType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

<xs:complexType name="sedQuantityType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="Value" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Accuracy" type="accuracyType" />
<xs:element minOccurs="0" maxOccurs="1" name="Resolution" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Quality" type="intParamType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

```

```

<!-- A set of useful types to add UCDS and units to base types; like BasicQuantity -->
<xs:complexType name="Group" >
<xs:attribute name="id" type="xs:ID" use="optional"/>
<xs:attribute name="idref" type="xs:IDREF" use="optional"/>
</xs:complexType>

<xs:complexType name="textParamType">
<xs:simpleContent>
<xs:extension base="paramType" />
</xs:simpleContent>
</xs:complexType>

<xs:complexType name="paramType">
<xs:simpleContent>
<xs:extension base="xs:string">
<xs:attribute name="name" type="xs:string" />
<xs:attribute name="ucd" type="xs:string" />
</xs:extension>
</xs:simpleContent>
</xs:complexType>

<xs:complexType name="dateParamType">
<xs:simpleContent>
<xs:extension base="paramType" />
</xs:simpleContent>
</xs:complexType>

<xs:complexType name="positionParamType">
<xs:complexContent mixed="false">
<xs:extension base="xs:anyType">
<xs:element minOccurs="2" maxOccurs="2" name="value" type="doubleParamType" />
</xs:complexContent>
</xs:complexType>

<xs:complexType name="doubleParamType">
<xs:simpleContent>
<xs:extension base="paramType">
<xs:attribute name="unit" type="xs:string" />
</xs:extension>
</xs:simpleContent>
</xs:complexType>

<xs:complexType name="timeParamType">
<xs:simpleContent>
<xs:extension base="paramType">
<xs:attribute name="unit" type="xs:string" />
</xs:extension>
</xs:simpleContent>
</xs:complexType>

<xs:complexType name="intParamType">
<xs:simpleContent>
<xs:extension base="paramType">
<xs:attribute name="unit" type="xs:string" />
</xs:extension>
</xs:simpleContent>
</xs:complexType>

```

```

<!-- The error model. Bin entries will usually be omitted for the flux coordinate -->
<xs:complexType name="accuracyType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="BinLow" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="BinHigh" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="BinSize" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="StatError" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="StatErrLow" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="StatErrHigh" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="SysError" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Confidence" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Calibration" type="textParamType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

<!-- The Field type allows us to define what our axes are -->
<xs:complexType name="arrayOfFieldType">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="unbounded" name="Field" nillable="true" type="fieldType" />
</xs:sequence>
</xs:complexType>

<xs:complexType name="fieldType">
<xs:attribute name="name" type="xs:string" />
<xs:attribute name="unit" type="xs:string" />
<xs:attribute name="ucd" type="xs:string" />
</xs:complexType>

<!-- The Point type groups a single set of time, spectral and flux values -->
<xs:element name="Data" type="arrayOfGenPointType"/>

<xs:complexType name="arrayOfGenPointType"/>

<xs:complexType name="arrayOfPointType">
<xs:complexContent mixed="false">
<xs:extension base="arrayOfGenPointType">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="unbounded" name="Point" nillable="true" type="pointType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

<xs:element name="ArrayOfPoint" type="arrayOfPointType" substitutionGroup="Data"/>

<xs:complexType name="pointType">
<xs:sequence>
<xs:element name="TimeAxis" minOccurs="0" maxOccurs="1" type="sedCoordType" />
<xs:element name="SpectralAxis" minOccurs="0" maxOccurs="1" type="sedCoordType" />
<xs:element name="FluxAxis" minOccurs="0" maxOccurs="1" type="sedQuantityType" />
<xs:element name="BackgroundModel" minOccurs="0" maxOccurs="1" type="sedQuantityType" />
</xs:sequence>
</xs:complexType>

```

```

<xs:element name="ArrayOfFlatPoint" type="ArrayOfFlatPointType" substitutionGroup="Data"/>

<xs:complexType name="ArrayOfFlatPointType">
<xs:complexContent mixed="false">
<xs:extension base="ArrayOfGenPointType">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="unbounded" name="Point" nillable="true" type="FlatPointType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

<xs:complexType name="FlatPointType">

<xs:attribute name="T" type="xs:double"/>
<xs:attribute name="T_BinL" type="xs:double" />
<xs:attribute name="T_BinH" type="xs:double" />
<xs:attribute name="T_Size" type="xs:double" />
<xs:attribute name="T_Res" type="xs:double" />
<xs:attribute name="SP" type="xs:double" />
<xs:attribute name="SP_BinL" type="xs:double" />
<xs:attribute name="SP_BinH" type="xs:double" />
<xs:attribute name="SP_Size" type="xs:double" />
<xs:attribute name="SP_Res" type="xs:double" />
<xs:attribute name="F" type="xs:double" />
<xs:attribute name="F_ErrL" type="xs:double" />
<xs:attribute name="F_ErrH" type="xs:double" />
<xs:attribute name="F_Sys" type="xs:double" />
<xs:attribute name="F_Qual" type="xs:int" />
<xs:attribute name="BG" type="xs:double" />
<xs:attribute name="BG_ErrL" type="xs:double" />
<xs:attribute name="BG_ErrH" type="xs:double" />
<xs:attribute name="BG_Sys" type="xs:double" />
<xs:attribute name="BG_Qual" type="xs:int" />
</xs:complexType>

```

```

<!-- Now we define the higher level metadata -->
<xs:complexType name="curationType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="Publisher" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="PublisherID" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Reference" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Version" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Contact" type="contactType" />
<xs:element minOccurs="0" maxOccurs="1" name="Rights" type="contactType" />
<xs:element minOccurs="0" maxOccurs="1" name="Date" type="contactType" />
<xs:element minOccurs="0" maxOccurs="1" name="PublisherDID" type="contactType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

<xs:complexType name="contactType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="Name" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Email" type="textParamType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

<xs:complexType name="characterizationType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element name="SpatialAxis" type="characterizationAxisType" minOccurs="0" maxOccurs="1"/>
<xs:element name="TimeAxis" type="characterizationAxisType" minOccurs="0" maxOccurs="1"/>
<xs:element name="SpectralAxis" type="characterizationAxisType" minOccurs="0" maxOccurs="1"/>
<xs:element name="FluxAxis" type="characterizationAxisType" minOccurs="0" maxOccurs="1"/>
<xs:element minOccurs="0" maxOccurs="unbounded" name="CharacterizationAxis"
type="characterizationAxisType"/>
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

<xs:complexType name="characterizationAxisType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="CoordSystem" type="coordSysType" />
<xs:element minOccurs="0" maxOccurs="1" name="Coverage" type="coverageType" />
<xs:element minOccurs="0" maxOccurs="1" name="Resolution" type="doubleParamType"/>
<xs:element minOccurs="0" maxOccurs="1" name="Accuracy" type="accuracyType" />
</xs:sequence>
<xs:attribute name="name" type="xs:string" />
<xs:attribute name="ucd" type="xs:string" />
<xs:attribute name="unit" type="xs:string" />
</xs:extension>
</xs:complexContent>
</xs:complexType>

```

```

<xs:element name="CharacterizationAxis" type="characterizationAxisType"/>

<xs:complexType name="coverageType">
  <xs:complexContent mixed="false">
    <xs:extension base="Group">
      <xs:sequence>
        <xs:element minOccurs="0" maxOccurs="1" name="Location" type="coverageLocationType" />
        <xs:element minOccurs="0" maxOccurs="1" name="Bounds" type="coverageBoundsType" />
        <xs:element minOccurs="0" maxOccurs="1" name="Support" type="coverageSupportType" />
      </xs:sequence>
    </xs:extension>
  </xs:complexContent>
</xs:complexType>

<xs:complexType name="coverageLocationType">
  <xs:complexContent mixed="false">
    <xs:extension base="sedCoordType"/>
  </xs:complexContent>
</xs:complexType>

<xs:complexType name="coverageBoundsType">
  <xs:complexContent mixed="false">
    <xs:extension base="Group">
      <xs:sequence>
        <xs:element minOccurs="0" maxOccurs="1" name="Extent" type="doubleParamType" />
        <xs:element minOccurs="0" maxOccurs="unbounded" name="Range" type="intervalType" />
      </xs:sequence>
    </xs:extension>
  </xs:complexContent>
</xs:complexType>

<xs:complexType name="coverageSupportType">
  <xs:complexContent mixed="false">
    <xs:extension base="Group">
      <xs:sequence>
        <xs:element minOccurs="0" maxOccurs="1" name="Area" type="skyRegionType" />
        <xs:element minOccurs="0" maxOccurs="unbounded" name="Range" type="intervalType" />
        <xs:element minOccurs="0" maxOccurs="1" name="Fill" type="doubleParamType" />
      </xs:sequence>
    </xs:extension>
  </xs:complexContent>
</xs:complexType>

<xs:complexType name="intervalType">
  <xs:complexContent mixed="false">
    <xs:extension base="Group">
      <xs:sequence>
        <xs:element minOccurs="0" maxOccurs="1" name="Min" type="doubleParamType" />
        <xs:element minOccurs="0" maxOccurs="1" name="Max" type="doubleParamType" />
      </xs:sequence>
    </xs:extension>
  </xs:complexContent>
</xs:complexType>

```

```

<xs:complexType name="skyRegionType">
<xs:complexContent mixed="false">
<xs:extension base="textParamType"/>
</xs:complexContent>
</xs:complexType>

<xs:complexType name="derivedDataType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="SNR" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="VarAmpl" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Redshift" type="sedQuantityType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

<xs:complexType name="dataIDType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="Title" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Creator" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="unbounded" name="Collection" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="DatasetID" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Date" type="dateParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Version" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Instrument" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="CreationType" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="unbounded" name="Contributor" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Logo" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="DataSource" type="textParamType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

<xs:complexType name="targetType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="Name" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Description" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="TargetClass" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="SpectralClass" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Redshift" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Pos" type="positionParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="VarAmpl" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="CustomParams" type="arrayOfParamType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>

```

```

<xs:complexType name="arrayOfParamType">
  <xs:sequence>
    <xs:element minOccurs="0" maxOccurs="unbounded" name="Param" nillable="true" type="paramType" />
  </xs:sequence>
</xs:complexType>

<xs:attributeGroup name="STCReference">
  <xs:annotation>
    <xs:documentation>These four attributes represent the standard IVOA referencing system: internal (within th
  </xs:annotation>
  <xs:attribute name="id" type="xs:ID" use="optional"/>
  <xs:attribute name="idref" type="xs:IDREF" use="optional"/>
  <xs:attribute name="ucd" type="xs:string" use="optional"/>
  <xs:attribute ref="xlink:type" use="optional" default="simple"/>
  <xs:attribute ref="xlink:href" use="optional"/>
</xs:attributeGroup>

<xs:complexType name="coordSysType">
  <!--<xs:complexContent> -->
    <xs:sequence maxOccurs="unbounded">
      <xs:element ref="CoordFrame" minOccurs="0" maxOccurs="unbounded"/>
    </xs:sequence>
    <xs:attributeGroup ref="STCReference"/>
  <!-- </xs:complexContent>-->
</xs:complexType>

<xs:element name="CoordFrame" type="coordFrameType" abstract="true"/>
<xs:element name="SpaceFrame" type="spaceFrameType" substitutionGroup="CoordFrame"/>
<xs:element name="RedshiftFrame" type="redshiftFrameType" substitutionGroup="CoordFrame"/>
<xs:element name="SpectralFrame" type="spectralFrameType" substitutionGroup="CoordFrame"/>
<xs:element name="GenericCoordFrame" type="coordFrameType" substitutionGroup="CoordFrame"/>
<xs:element name="TimeFrame" type="timeFrameType" substitutionGroup="CoordFrame"/>

<xs:complexType name="coordFrameType">
  <xs:annotation>
    <xs:documentation>Simplification of STC version: RefPos is string</xs:documentation>
  </xs:annotation>
  <xs:sequence>
    <xs:element name="Name" type="xs:string" minOccurs="0"/>
    <xs:element name="ReferencePosition" type="xs:string" minOccurs="0"/>
  </xs:sequence>
  <xs:attribute name="id" type="xs:ID"/>
  <xs:attribute name="ucd" type="xs:string" use="optional"/>
</xs:complexType>

<xs:complexType name="spectralFrameType">
  <xs:complexContent>
    <xs:extension base="coordFrameType">
      <xs:element minOccurs="0" maxOccurs="1" name="redshift" type="doubleParamType" />
    </xs:extension>
  </xs:complexContent>
</xs:complexType>

```

```

<xs:complexType name="timeFrameType">
  <xs:complexContent>
    <xs:extension base="coordFrameType">
      <xs:element minOccurs="0" maxOccurs="1" name="Zero" type="doubleParamType" />
    </xs:extension>
  </xs:complexContent>
</xs:complexType>

<xs:complexType name="redshiftFrameType">
  <xs:complexContent>
    <xs:extension base="coordFrameType">
      <xs:sequence>
        <xs:element name="DopplerDefinition" type="xs:string" nillable="true"/>
      </xs:sequence>
    </xs:extension>
  </xs:complexContent>
</xs:complexType>

<xs:complexType name="spaceFrameType">
  <xs:complexContent>
    <xs:extension base="coordFrameType">
      <xs:sequence>
        <xs:element minOccurs="0" maxOccurs="1" name="Equinox" type="doubleParamType" />
      </xs:sequence>
    </xs:extension>
  </xs:complexContent>
</xs:complexType>

</xs:schema>

```

7.2 Instance example

```
<?xml version="1.0" encoding="UTF-8"?>
<Spectrum xmlns="http://www.ivoa.net/xml/Spectrum/Spectrum-0.98.xsd"
xmlns:xsi="http://www.w3.org/2001/XMLSchema-instance"
xsi:schemaLocation="http://www.ivoa.net/xml/Spectrum/Spectrum-0.98.xsd">

<!-- xml instance example -->
  <Target>
    <name>Arp 220 </name>
    <pos><value>233.737917</value><value>23.503330</value></pos>
    <redshift>0.0018</redshift>
  </Target>
  <SegmentType ucd="meta.code">Spectrum</SegmentType>
<!-- STC metadata, with references to defined fields -->
  <CoordSys id="ID000001">
    <TimeFrame>
      <Name>UTC</Name>
    <ReferencePosition>BARYCENTER</ReferencePosition>
    </TimeFrame>
    <SpaceFrame>
      <Name>ICRS</Name>
    <ReferencePosition>BARYCENTER</ReferencePosition>
    </SpaceFrame>
    <SpectralFrame ucd="em.wavelength">
      <Name>Wavelength</Name>
    <ReferencePosition>HELIOCENTER</ReferencePosition>
    </SpectralFrame>
    <GenericCoordFrame id="FD" ucd="phot.fludens;em.wl">
      <Name>Flux density</Name>
    <GenericCoordFrame>
  </CoordSys>
```

```

<Characterization>
  <FluxAxis>
    <name>Flux density</name>
    <ucd>phot.flux;em.wavelength</ucd>
    <unit>erg cm**(-2) s**(-1) Angstrom**(-1)"</unit>
    <Accuracy>
      <StatErrLow unit="erg cm**(-2) s**(-1) Angstrom**(-1)" id="Field5"/>
      <StatErrHigh unit="erg cm**(-2) s**(-1) Angstrom**(-1)" id="Field6"/>
      <SysErr>0.05</SysErr>
    </Accuracy>
  </FluxAxis>
  <SpatialAxis>
    <name>Sky</name>
    <ucd>pos.eq</ucd>
    <unit>deg</unit>
    <coordsystem ID="ID000001">
    <Coverage>
      <Location><Value ucd="pos.eq.ra" unit="deg">132.4210</Value>
        <Value ucd="pos.eq.dec">12.1232</Value> </Location>
    <Bounds>
      <Extent ucd="pos.region.diameter" unit="arcsec">20 </Extent>
    </Bounds>
    </Coverage>
    <Accuracy>
      <Calibration>CALIBRATED</Calibration>
    </Accuracy>
  </SpatialAxis>
  <TimeAxis>
    <name>Time</name>
    <ucd>time</ucd>
    <unit>d</unit>
    <coordsystem ID="ID000001">
    <Coverage>
      <Location><value ucd="time.obs" unit="d">52148.3252</value></Location>
    <Bounds>
      <Extent ucd="time.expo;phot.spectrum" unit="s">1500.0</Extent>
    </Bounds>
    </Coverage>
    <Accuracy>
      <Calibration>CALIBRATED</Calibration>
    </Accuracy>
  </TimeAxis>
  <SpectralAxis>
    <name>SpectralCoord</name>
    <ucd>em.wl</ucd>
    <unit>Angstrom</unit>
    <coordsystem ID="ID000001">
    <Coverage>
      <Bounds>
        <Extent ucd="instr.bandwidth" unit="Angstrom">3000.0</Extent>
      </Bounds>
    </Coverage>
    <Accuracy>
      <BinLow ucd="stat.min;em.wavelength" unit="Angstrom"/>
      <BinHigh ucd="stat.max;em.wavelength" unit="Angstrom"/>
      <Calibration>CALIBRATED</Calibration>
    </Accuracy>
  </SpectralAxis>
</Characterization>

```

```
<Curation>
  <Publisher ucd="meta.organization;meta.curation">SAO</Publisher>
  <PubID ucd="meta.curation.pubid">ivoa://cfa.harvard.edu</PubID>
  <Logo ucd="meta.curation.logo">http://cfa-www.harvard.edu/nvo/cfalogo.jpg</Logo>
  <Contact>
    <Name ucd="meta.human;meta.curation">Jonathan McDowell</Name>
    <Email ucd="meta.email">jcm@cfa.harvard.edu</Email>
  </Contact>
</Curation>
<DataID>
  <Title>Arp 220 SED</Title>
  <Creator ucd="meta.curation.creator">SAO/FLWO</Creator>
  <Date ucd="time;soft.dataset;meta.curation">2003-12-31T14:00:02Z</Date>
  <Version ucd="soft.dataset.version;meta.curation">1</Version>
  <Instrument ucd="inst.id">BCS</Instrument>
  <Collection>Archival</Collection>
</DataID>
<Derived>
  <SNR>3.0</SNR>
</Derived>
```

```

<!-- Use table structure -->
<Data>
  <Point>
    <SpectralCoord>
      <Value>3200.0</Value>
      <Accuracy><BinLow>3195.0</BinLow><BinHigh>3205.0</BinHigh></Accuracy>
    </SpectralCoord>
    <Flux>
      <Value>1.38E-12</Value>
      <Accuracy><StatErrLow>5.2E-14</StatErrLow><StatErrHigh>6.2E-14</StatErrHigh></Accuracy>
      <Quality>0</Quality>
    </Flux>
  </Point>

  <Point>
    <SpectralCoord>
      <Value>3210.5</Value>
      <Accuracy><BinLow>3205.0</BinLow><BinHigh>3216.0</BinHigh></Accuracy>
    </SpectralCoord>
    <Flux>
      <Value>1.12E-12</Value>
      <Accuracy><StatErrLow>1.12E-12</StatErrLow><StatErrHigh>0</StatErrHigh></Accuracy>
      <Quality>0</Quality>
    </Flux>
  </Point>

  <Point>
    <SpectralCoord>
      <Value>3222.0</Value>
      <Accuracy><BinLow>3216.0</BinLow><BinHigh>3228.0</BinHigh></Accuracy>
    </SpectralCoord>
    <Flux>
      <Value>1.42E-12</Value>
      <Accuracy><StatErrLow>1.3E-14</StatErrLow><StatErrHigh>0.2E-14</StatErrHigh></Accuracy>
      <Quality>3</Quality>
    </Flux>
  </Point>
  ...
</Data>

```

OR

```

<Data>
<FlatPoint SP=3200.0 SP_BinL=3195.0 SP_BinH=3205.0
          F=1.48E-12 F_ErrL=2.0E-14 F_ErrH=2.0E-14/>
<FlatPoint SP=3210.0 SP_BinL=3205.0 SP_BinH=3215.0
          F=1.48E-12 F_ErrL=3.2E-14 F_ErrH=3.8E-14/>
<FlatPoint SP=3220.0 SP_BinL=3215.0 SP_BinH=3225.0
          F=1.48E-12 F_ErrL=1.48E-12 F_ErrH=0.0 />
  ...
</Data>
</Spectrum>

```

Part 3: VOTABLE serialization

8 VOTABLE serialization

8.1 Mapping Schema to VOTABLE

We reproduce below the XML schema instance example as a VOTABLE instance example. To go from the XML instance to the VOTABLE instance, we:

- - map the top level element to a RESOURCE
- - map all elements with simple content to PARAM
- - map all elements with complex content to GROUP
- - map the element names (with appropriate path) to values of the utype attribute,
- - but, handle the FIELDS and Data elements in a special way. The FIELDS element is used to define the table fields and the Data element is used to define the table data.
- - but, also, all the second level elements below RESOURCE except SPECTRUM map to an initial TABLE, while we map SPECTRUM to a second TABLE.
- - most of the elements extend the Param element, to which I have added an optional name attribute that I have not used in the instance. If this attribute is used, it can hold the name attributes of the PARAM and FIELD; otherwise the relevant attributes could be filled with the same value as the utype (without namespace prefix).

How can this be generalized to mapping an arbitrary data model schema to VOTABLE? The only tricky parts are

- **Spotting where the tabledata parts are.** We could require any DM schema that maps to VOTABLE to include elements called FIELDS and Data (perhaps ROWS would be a better name), otherwise you would get a VOTABLE with no data section.
- **Spotting where to start the main TABLE (i.e. the fact that SPECTRUM is special).** We could change the schema to have an explicit attribute, annotation or other marker to tell us this.

These issues will require further discussion for future models.

8.2 A VOTABLE instance

The VOTable version of Spectrum uses a single VOTable <TABLE>. (Note that this may appear as one of many tables within an SED VOTable). The data model fields described above as arrays map to VOTable FIELDS, while the remaining fields map to PARAM.

We use nested GROUP constructs to delimit data model objects within the main object, and PARAM and FIELD tags for attributes. Names of fields and parameters are left to the data provider. The utype and ucd attributes are used to denote data model and UCD tags. The schema and namespace for the utypes is the XML schema given in section 8.4. I have made up arbitrary NAME attributes for the PARAM and these are not to be considered standard; the name fields are free to be whatever the data provider wants, allowing compatibility with local archive nomenclature. The NAME attributes for the FIELD elements are also not standardized (of course they must be the same as in the matching FIELDrefs); it is the utype attribute which is standardized.

The one departure from the XML schema below is that the 'ArrayOfPoints' element and the individual 'Point' elements are implicitly represented by the table structure itself. Perhaps a UTYPE attribute to the TABLEDATA element could be used to make this explicit.

The example below describes a single SPECTRUM.

```
<?xml version="1.0" encoding="UTF-8"?>
<VOTABLE version="1.1"
  xmlns:xsi="http://www.w3.org/2001/XMLSchema-instance"
  xsi:noNamespaceSchemaLocation="http://www.ivoa.net/xml/VOTable/v1.1"
  xmlns:spec="http://www.ivoa.net/xml/SpectrumModel/v0.98"
  xmlns="http://www.ivoa.net/xml/VOTable/v1.1">
<RESOURCE utype="spec:Spectrum">

<TABLE utype="spec:Spectrum">
<GROUP utype="spec:Target">
<PARAM name="Target" utype="spec:Target.Name" datatype="char" arraysize="*" value="Arp 220"/>
<PARAM name="TargetPos" utype="spec:Target.pos" unit="deg" datatype="double"
  arraysize="2" value="233.737917 23.503330"/>
<PARAM name="z" utype="spec:Target.redshift" datatype="float" value="0.0018"/>
</GROUP>

<!-- SegmentType can be Photometry, TimeSeries or Spectrum -->
<PARAM name="Segtype" utype="spec:Spectrum.SegmentType" datatype="char" arraysize="*"
  value="Photometry" ucd="meta.code"/>
<GROUP name="CoordSys" utype="spec:Spectrum.CoordSys">
<GROUP utype="spec:Spectrum.CoordSys.SpaceFrame">
  <PARAM name="System" utype="spec:Spectrum.CoordSys.SpaceFrame.Name" ucd="frame.pos.system"
    datatype="char" arraysize="*" value="ICRS"/>
  <PARAM name="Equinox" utype="spec:Spectrum.CoordSys.SpaceFrame.Equinox" ucd="time.equinox;pos.eq"
    datatype="float" value="2000.0" />
</GROUP>
<GROUP utype="spec:Spectrum.CoordSys.TimeFrame">
  <PARAM name="TimeFrame" utype="spec:Spectrum.CoordSys.TimeFrame.Name" ucd="frame.time.scale" datatype="char"
    arraysize="*" value="UTC"/>
</GROUP>
<GROUP utype="spec:Spectrum.CoordSys.SpectralFrame">
  <PARAM name="SpectralFrame" utype="spec:Spectrum.CoordSys.SpectralFrame.RefPos" ucd="frame.em.system"
    datatype="char" arraysize="*" value="BARYCENTER"/>
</GROUP>
</GROUP>
</TABLE>
</RESOURCE>
</VOTABLE>
```

```

<GROUP utype="spec:Char">
  <GROUP utype="spec:Char.SpatialAxis">
    <PARAM name="SpatialAxisName" utype="name" ucd="pos.eq" unit="deg" value="Sky"/>
    <GROUP utype="spec:Char.Coverage">
      <GROUP utype="spec:Char.Coverage.Location">
        <PARAM name="SkyPos" utype="Value"
          ucd="pos.eq" unit="deg"
          datatype="double" arraysize="2" value="132.4210 12.1232"/>
      </GROUP>
    <GROUP utype="Bounds">
      <PARAM name="SkyExtent" utype="Extent" ucd="pos.region.diameter"
        datatype="double" unit="arcsec" value="20"/>
    </GROUP>
  </GROUP>
  <GROUP utype="spec:Char.TimeAxis">
    <PARAM name="TimeAxisName" utype="name" ucd="time" unit="d" value="Time"/>
    <GROUP utype="Coverage">
      <GROUP utype="Location">
        <PARAM name="TimeObs" utype="Value" ucd="time.obs"
          datatype="double" value="52148.3252"/>
      </GROUP>
      <GROUP utype="Bounds">
        <PARAM name="TimeExtent" utype="Extent" ucd="time.expo;phot.spectrum"
          unit="s" datatype="double" value="1500.0" />
        <GROUP utype="Range">
          <PARAM name="TimeStart" utype="Start" ucd="time" unit="s"
            datatype="double" value="52100.000" />
          <PARAM name="TimeStop" utype="Stop" ucd="time" unit="s"
            datatype="double" value="52300.000" />
        </GROUP>
      </GROUP>
    </GROUP>
  </GROUP>
  <GROUP utype="spec:Char.SpectralAxis">
    <PARAM name="SpectralAxisName" utype="name" ucd="em.wl" unit="Angstrom" value="Wavelength"/>
    <GROUP utype="Coverage">
      <GROUP utype="Bounds">
        <PARAM name="SpectralExtent" utype="Extent" ucd="instr.bandwidth"
          unit="Angstrom" datatype="double" value="3000.0"/>
      </GROUP>
    </GROUP>
  </GROUP>
</GROUP>

```

```

<GROUP utype="spec:Spectrum.Curation">
  <PARAM name="Publisher" utype="spec:Spectrum.Curation.Publisher" ucd="meta.organization;meta.curation"
    datatype="char" arraysize="*" value="SA0"/>
  <PARAM name="PubID" utype="spec:Spectrum.Curation.PubID" ucd="meta.curation.pubid" datatype="char"
    arraysize="*" value="ivoa://cfa.harvard.edu"/>
  <PARAM name="Logo" utype="spec:Spectrum.Curation.Logo" ucd="meta.curation.logo" datatype="char"
    arraysize="*" value="http://cfa-www.harvard.edu/nvo/cfalogo.jpg"/>
  <PARAM name="Contact" utype="spec:Spectrum.Curation.ContactName" ucd="meta.human;meta.curation"
    datatype="char" arraysize="*" value="Jonathan McDowell"/>
  <PARAM name="email" utype="spec:Spectrum.Curation.ContactEmail" ucd="meta.email" datatype="char"
    arraysize="*" value="jcm@cfa.harvard.edu"/>
</GROUP>

<GROUP utype="spec:Spectrum.DataID">
  <PARAM name="Title" utype="spec:Spectrum.DataID.Title" datatype="char" arraysize="*" value="Arp 220 SED"/>
  <PARAM name="Creator" utype="spec:Segment.DataID.Creator" ucd="meta.curation.creator" datatype="char"
    arraysize="*" value="SA0/FLWO"/>
  <PARAM name="DataDate" utype="spec:Spectrum.DataID.Date" ucd="time;soft.dataset;meta.curation"
    datatype="char" arraysize="*" value="2003-12-31T14:00:02Z"/>
  <PARAM name="Version" utype="spec:Spectrum.DataID.Version" ucd="soft.dataset.version;meta.curation"
    datatype="char" arraysize="*" value="1"/>
  <PARAM name="Instrument" utype="spec:Spectrum.DataID.Instrument" ucd="inst.id" datatype="char"
    arraysize="*" value="BCS"/>
  <PARAM name="Filter" utype="spec:Spectrum.DataID.Collection" ucd="inst.filter.id" datatype="char"
    arraysize="*" value="G300"/>
  <PARAM name="CreationType" utype="spec:Spectrum.DataID.CreationType" datatype="char" arraysize="*" value="Archiv
</GROUP>

<GROUP utype="spec:Spectrum.Derived">
  <PARAM name="SNR" utype="spec:Spectrum.Derived.SNR" datatype="float" value="3.0"/>
</GROUP>

<GROUP utype="spec:Spectrum.Data">

<GROUP utype="spec:Spectrum.Data.SpectralAxis">
  <FIELDref ref="Coord"/>

  <GROUP utype="spec:Spectrum.Data.SpectralAxis.Accuracy">
    <FIELDref ref="BinLow"/>
    <FIELDref ref="BinHigh"/>
  </GROUP>
  <!-- In this case Resolution is demoted from Field to Param since it is constant -->
  <PARAM name="Resolution" utype="spec:Spectrum.Data.SpectralAxis.Resolution"
    unit="Angstrom" datatype="float" value="14.2"/>
</GROUP>

```

```

<GROUP utype="spec:Spectrum.Data.FluxAxis">
  <FIELDref ref="Flux1"/>
  <GROUP utype="spec:Spectrum.Data.FluxAxis.Accuracy">
    <FIELDref ref="ErrorLow"/>
    <FIELDref ref="ErrorHigh"/>
    <PARAM name="SysErr" utype="SysErr" unit="" datatype="float" value="0.05"/>
  </GROUP>
  <FIELDref ref="Quality"/>
</GROUP>
</GROUP>

<FIELD name="Coord" ID="Coord" utype="spec:Spectrum.Data.SpectralAxis.Value" ucd="em.wavelength"
  datatype="double" unit="Angstrom"/>
<FIELD name="BinLow" ID="BinLow" utype="spec:Spectrum.Data.SpectralAxis.BinLow"
  ucd="stat.min;em.wavelength"
  datatype="double" unit="Angstrom"/>
<FIELD name="BinHigh" ID="BinHigh" utype="spec:Spectrum.Data.SpectralAxis.BinHigh"
  ucd="stat.max;em.wavelength"
  datatype="double" unit="Angstrom"/>
<FIELD name="Flux" ID="Flux1" utype="spec:Spectrum.Data.FluxAxis.value" ucd="phot.flux;em.wavelength"
  datatype="double" unit="erg cm**(-2) s**(-1) Angstrom**(-1)"/>
<FIELD name="ErrorLow" ID="ErrorLow" utype="spec:Spectrum.Data.FluxAxis.Accuracy.StatErrLow"
  datatype="double" unit="erg cm**(-2) s**(-1) Angstrom**(-1)"/>
<FIELD name="ErrorHigh" ID="ErrorHigh" utype="spec:Spectrum.Data.FluxAxis.Accuracy.StatErrHigh"
  datatype="double" unit="erg cm**(-2) s**(-1) Angstrom**(-1)"/>
<FIELD name="Quality" ID="Quality" datatype="int" utype="spec:Spectrum.Data.FluxAxis.Quality"/>
<DATA>
<TABLEDATA>
<!-- Note slightly nonlinear wavelength solution -->
<!-- Second row is upper limit -->
<!-- Third row has quality mask set -->
<TR><TD>3200.0</TD><TD>3195.0</TD><TD>3205.0</TD><TD>1.38E-12</TD><TD>5.2E-14</TD><TD>6.2E-14</TD>
  <TD>0</TD></TR>
<TR><TD>3210.5</TD><TD>3205.0</TD><TD>3216.0</TD><TD>1.12E-12</TD><TD>1.12E-12</TD>
  <TD>0</TD><TD>0</TD></TR>
<TR><TD>3222.0</TD><TD>3216.0</TD><TD>3228.0</TD><TD>1.42E-12</TD><TD>1.3E-14</TD>
  <TD>0.2E-14</TD><TD>3</TD></TR>
</TABLEDATA>
</DATA>
</TABLE>
</RESOURCE>
</VOTABLE>

```

Part 4: FITS serialization

9 FITS serialization

9.1 Mapping Spectrum to FITS

FITS serialization design: We define a reference serialization of this data model as a FITS binary table. The table represents a spectrum or photometry point as a single row of a table.

This serialization is a special case of an SED serialization which uses one row per spectral segment; in that case, variable-length arrays may be used to contain the array quantities. In each case below where a 'variable length array' is specified, fixed length arrays are suitable for a single spectrum or for multiple spectra where all the arrays are the same length, but readers should be prepared to handle the variable length case.

For SEDs, another approach would be to have one FITS HDU per spectrum or photometry point. However this was rejected as unworkable, as the overhead of 5760 bytes (2 FITS blocks) per photometry point would inflate the data for the photometry-only SED case by factors of around 50-100.

Standard table keywords: In table F.1 we give the mapping of data model fields to FITS columns and keywords. For each column, the standard keywords TTYPE_n, TUNIT_n, TFORM_n should be provided. Order of keywords and columns is not significant, except that it is strongly recommended that RA and Dec be in adjacent columns or keywords. Additional keywords and columns which are not part of the model (including other conventions such as e.g. TDMIN_n) are allowed to be present, but are not guaranteed to be propagated by VO software.

Keywords and columns: 'Greenbank' convention In Table F.1 we give single metadata items as keywords; arrays of data (members of the Spectrum.Data classes) are stored as columns, and Table F.1 gives the column name, i.e. the value of the keyword TTYPE_n. The 'Source' column in Table F.1 indicates if the name is a FITS standard (S), an existing convention (C) such as one of the HEA conventions, or is newly invented (N).

In some cases, the column data arrays may have the same value for each data point. In this case we may use the 'Greenbank' convention in which the column is omitted and replaced by a keyword whose name is the same as the column. Further, in SED applications when multiple spectrum data lines are present, some metadata may differ from line to line and be promoted from keyword to column. Therefore, implementors should check both keywords and column names for the appropriate tokens.

TUTYP and TUCD keywords: We map the FITS columns to the model by using TUTYP_n keywords. TUTYP_n (string-valued) gives the data model field name (UTYPE string) for the data in column n. Thus, the x and y axes (i.e. spectral coordinate and flux-like axes) of the spectrum have TUTYP_n value of Spectrum.Data.SpectralAxis.Value and Spectrum.Data.FluxAxis.Value respectively.

Different kinds of x and y axis are identified by the Spectrum.Data.SpectralAxis.UCD and Spectrum.Data.FluxAxis.UCD data model fields, which are mapped to TUCD_n keywords. TUCD_n (string valued) gives the UCD corresponding to the data in column n. Both TUTYP_n and TUCD_n should be present for any column which corresponds to a Spectrum data model field; they are optional for any additional data columns which are not part of the Spectrum model. The units of spectral coordinate and flux are given in the TUNIT_n keys of the corresponding data columns. There is no separate provision for units of Char.SpectralAxis or Char.FluxAxis; these are required to be the same as for the data.

The TTYPE_n keywords for the x and y columns are free, but it is strongly recommended that (for consistency of style with WCS Paper 3) the values for the x axis have for their first 4 characters 'WAVE', 'FREQ' and 'ENER' for the case of wavelength, frequency and energy respectively. We also recommend the value 'FLUX' for the y axis, where appropriate. Nevertheless, it is the TUTYP_n and TUCD_n keywords that should be used to interpret the semantics of the file.

WAVE, ENER, and FREQ In the header metadata, such as the Spectrum.Char entries, we use SPEC_ keywords to denote the spectral axis generically, but in the table columns (Spectrum.Data entries) we use the terms WAVE_, ENER_, and FREQ_ as appropriate. Thus if the Spectrum.Data.SpectralAxis.Value field is WAVE, the SpectralAxis.Accuracy.BinLow field should be WAVE_LO; if Value is FREQ, BinLow should be FREQ_LO. We believe the small extra parsing overhead is worth it for the readability and interoperability (since these names have been used in existing FITS files) of the crucial main data table.

Char and Data keywords: The model contains both Characterization metadata, giving over-all typical values for quantities such as spectral resolution, and the Data object, which can include such quantities on a per-pixel basis. In some cases, the FITS serialization allows the same token for both Char (as a keyword) and Data (as a column name). The name, unit and UCD fields for Char.FluxAxis and Char.SpectralAxis are required to be the same as for Data.FluxAxis and Data.SpectralAxis. The case of TimeAxis is a little different, since there may be no Data.TimeAxis present, and there exist already some HEA conventions for recording TimeAxis characterization, notably the TIMEUNIT keyword.

VOCLASS keyword: We add a new keyword VOCLASS to describe the VO object represented by the FITS table. The value of VOCLASS should be 'SPECTRUM 1.00'.

WCS table keywords: The spectral coordinate may also be identified by optional 1S_n and 1CTYP_n keywords as per WCS Paper 3. Table 9 of that paper implies that each data column which is a function of the spectral coord needs a pair of such keywords. Applications which implement the spectrum data model may ignore the WCS keys and interpret the file by recognizing 'by spec' (using TUTYP_n) which column is the spectral coordinate and that FLUX, etc. are functions of it, but the WCS keys give a general FITS application a chance at making sense of the file. In the example, TTYPE5='ERR_LO' and TUTYP5='Spectrum.Data.FluxAxis.Accuracy.StatErrLow'; the WCS keyword 1CTYPE5='WAVE-TAB' indicates that the data in column 5 is a function of wavelength, and that the wavelengths are in a lookup table. The WCS keyword 1S5.1='WAVE' indicates that the lookup table for the x-axis of column 5 (in this case, the wavelengths that the ERR_LO values correspond to) is in the column with TTYPE_n='WAVE', in this case column 1.

Table F.1: FITS keywords for VO Spectrum

Data model field	FITS keyword	Source	Value if fixed
Spectrum.DataModel	VOCLASS	N	SPECTRUM 1.0
Spectrum.Length	SIZE	N	
Spectrum.Type	VOSEGT	N	
Spectrum.CoordSys.ID	VOCSID	N	
Spectrum.CoordSys.SpaceFrame.Name	RADECSYS	S	e.g. ICRS or FK5
Spectrum.CoordSys.SpaceFrame.Equinox	EQUINOX	S	e.g. 2000.0
Spectrum.CoordSys.TimeFrame.Name	TIMESYS	C	TT (required)
Spectrum.CoordSys.TimeFrame.Zero	MJDREF	C	default 0.0
Spectrum.CoordSys.TimeFrame.RefPos			(not used)
Spectrum.CoordSys.SpectralFrame.RefPos	SPECSYS	S	(see below)
Spectrum.Curation.Publisher	VOPUB	N	
Spectrum.Curation.Reference	VOREF	N	
Spectrum.Curation.PublisherID	VOPUBID	N	
Spectrum.Curation.Version	VOVER	N	
Spectrum.Curation.ContactName	CONTACT	N	
Spectrum.Curation.ContactEmail	EMAIL	N	
Spectrum.Curation.Rights	VORIGHTS	N	
Spectrum.Curation.Date	VODATE	N	
Spectrum.Curation.PublisherDID	DS.IDPUB	N	
Spectrum.Target.Name	OBJECT	S	
Spectrum.Target.Description	OBJDESC	N	
Spectrum.Target.Class	SRCCLASS	N	
Spectrum.Target.spectralClass	SPECTYPE	N	
Spectrum.Target.redshift	REDSHIFT	C	
Spectrum.Target.pos	RA_TARG, DEC_TARG	C	
Spectrum.Target.VarAmpl	TARGVAR	N	
Spectrum.DataID.Title	TITLE	C	
Spectrum.DataID.Creator	CREATOR	C	
Spectrum.DataID.Collection	COLLECTn	N	
Spectrum.DataID.DatasetID	DS.IDENT	N	
Spectrum.DataID.CreatorDID	CR.IDENT	N	
Spectrum.DataID.Date	DATE	S	
Spectrum.DataID.Version	VERSION	C	
Spectrum.DataID.Instrument	INSTRUME	S	
Spectrum.DataID.CreationType	CRETYPE	N	
Spectrum.DataID.Logo	VOLOGO	N	
Spectrum.DataID.Contributor	CONTRIBn	N	
Spectrum.DataID.DataSource	DSSOURCE	N	
Spectrum.DataID.Bandpass	SPECBAND	N	
Spectrum.Derived.SNR	DER_SNR	N	
Spectrum.Derived.redshift.value	DER_Z	N	
Spectrum.Derived.redshift.statError	DER_ERR	N	
Spectrum.Derived.redshift.Confidence	DER_CONF	N	
Spectrum.Derived.VarAmpl	DER_VAR	N	
Spectrum.TimeSI	TIMESDIM	N	(required)
Spectrum.SpectralSI	SPECSDIM	N	(required)
Spectrum.FluxSI	FLUXSDIM	N	(required)

Data model field	FITS keyword	Source	Value if fixed
Omitted Char fields, values inherited from Spectrum.Data			
Spectrum.Char.FluxAxis.Name	-	-	TTYPE _n for FLUX
Spectrum.Char.FluxAxis.Unit	-	-	Same as Data
Spectrum.Char.FluxAxis.ucd	-	-	Same as Data
Spectrum.Char.SpectralAxis.Name	-	-	Same as Data
Spectrum.Char.SpectralAxis.Unit	-	-	Same as Data
Spectrum.Char.SpectralAxis.ucd	-	-	Same as Data
Spectrum.Char.TimeAxis.Name	-	-	TIME
Spectrum.Char.TimeAxis.ucd	-	-	time
Spectrum.Char.SpatialAxis.Name	-	-	(not used)
Spectrum.Char.SpatialAxis.Unit	-	-	deg
Char Fields which are the same as for Spectrum.Data			
Spectrum.Char.FluxAxis.Accuracy.StatError	STAT_ERR	C	
Spectrum.Char.FluxAxis.Accuracy.SysError	SYS_ERR	C	
Spectrum.Char.FluxAxis.Quality	QUALITY	C	
Spectrum.Char.TimeAxis.Accuracy.StatError	TIME_ERR	N	
Spectrum.Char.TimeAxis.Accuracy.SysError	TIME_SYE	N	
Spectrum.Char.TimeAxis.Resolution	TIME_RES	N	
Char Fields which are only present in Char			
Spectrum.Char.FluxAxis.Accuracy.Calibration	FLUX_CAL	N	
Spectrum.Char.SpectralAxis.Accuracy.Calibration	SPEC_CAL	N	
Spectrum.Char.SpectralAxis.Coverage.Location.Value	SPEC_VAL	N	
Spectrum.Char.SpectralAxis.Coverage.Bounds.Extent	SPEC_BW	N	
Spectrum.Char.SpectralAxis.Coverage.Support.Fill	SPEC_FIL	N	
Spectrum.Char.SpectralAxis.Accuracy.BinSize	SPEC_BIN	N	
Spectrum.Char.SpectralAxis.Accuracy.StatError	SPEC_ERR	N	
Spectrum.Char.SpectralAxis.Accuracy.SysError	SPEC_SYE	N	
Spectrum.Char.SpectralAxis.Resolution	SPEC_RES	N	
Spectrum.Char.TimeAxis.Unit	TIMEUNIT	C	
Spectrum.Char.TimeAxis.Accuracy.BinSize	TIMEDEL	C	
Spectrum.Char.TimeAxis.Accuracy.Calibration	TIME_CAL	N	
Spectrum.Char.TimeAxis.Coverage.Location.Value	DATE-OBS	S	
Spectrum.Char.TimeAxis.Coverage.Bounds.Extent	EXPOSURE	C	
Spectrum.Char.TimeAxis.Coverage.Bounds.Start	TSTART	C	
Spectrum.Char.TimeAxis.Coverage.Bounds.Stop	TSTOP	C	
Spectrum.Char.TimeAxis.Coverage.Support.Fill	DTCOR	C	
Spectrum.Char.SpatialAxis.ucd	SKY_UCD	N	pos.eq
Spectrum.Char.SpatialAxis.Accuracy.StatErr	SKY_ERR	N	
Spectrum.Char.SpatialAxis.Accuracy.SysError	SKY_SYE	N	
Spectrum.Char.SpatialAxis.Accuracy.Calibration	SKY_CAL	N	
Spectrum.Char.SpatialAxis.Resolution	SKY_RES	N	
Spectrum.Char.SpatialAxis.Coverage.Location.Value	RA, DEC, etc.	C	
Spectrum.Char.SpatialAxis.Coverage.Bounds.Extent	APERTURE	C	
Spectrum.Char.SpatialAxis.Coverage.Support.Area	REGION	N	String value
Spectrum.Char.SpatialAxis.Coverage.Support.Fill	SKY_FILL	N	

Data model field	FITS keyword	Source	Value if fixed
Per-data-point values			
Spectrum.Data.FluxAxis.Value	TTYPE _n	S	FLUX
Spectrum.Data.FluxAxis.Unit	TUNIT _n	S	
Spectrum.Data.FluxAxis.ucd	TUCD _n	N	(same as Char)
Spectrum.Data.FluxAxis.Accuracy.StatError	TTYPE _n	N	ERR
Spectrum.Data.FluxAxis.Accuracy.StatErrLow	TTYPE _n	C	ERR_LO
Spectrum.Data.FluxAxis.Accuracy.StatErrHigh	TTYPE _n	C	ERR_HI
Spectrum.Data.FluxAxis.Accuracy.SysError	TTYPE _n	C	SYS_ERR
Spectrum.Data.FluxAxis.Quality	TTYPE _n	C	QUALITY
Spectrum.Data.SpectralAxis.Value	TTYPE _n	S	WAVE,ENER,FREQ
Spectrum.Data.SpectralAxis.Unit	TUNIT _n	S	(same as Char)
Spectrum.Data.SpectralAxis.ucd	TUCD _n	N	(same as Char)
Spectrum.Data.SpectralAxis.Accuracy.BinSize	TTYPE _n	N	WAVE_BIN,ENER_BIN, FREQ_BIN
Spectrum.Data.SpectralAxis.Accuracy.BinLow	TTYPE _n	N	WAVE_LO,ENER_LO, FREQ_LO
Spectrum.Data.SpectralAxis.Accuracy.BinHigh	TTYPE _n	N	WAVE_HI,ENER_HI, FREQ_HI
Spectrum.Data.SpectralAxis.Accuracy.StatError	TTYPE _n	N	WAVE_ERR,ENER_ERR, FREQ_ERR
Spectrum.Data.SpectralAxis.Accuracy.StatErrLow	TTYPE _n	N	WAVE_ELO,ENER_ELO, FREQ_ELO
Spectrum.Data.SpectralAxis.Accuracy.StatErrHigh	TTYPE _n	N	WAVE_EHI,ENER_EHI, FREQ_EHI
Spectrum.Data.SpectralAxis.Accuracy.SysError	TTYPE _n	N	WAVE_SYE,ENER_SYE, FREQ_SYE
Spectrum.Data.SpectralAxis.Resolution	TTYPE _n	N	WAVE_RES,ENER_RES, FREQ_RES
Spectrum.Data.TimeAxis.Value	TTYPE _n	C	TIME
Spectrum.Data.TimeAxis.Unit	TUNIT _n	S	(same as Char)
Spectrum.Data.TimeAxis.ucd	TUCD _n	N	time
Spectrum.Data.TimeAxis.Accuracy.BinLow	TTYPE _n	N	TIME_LO
Spectrum.Data.TimeAxis.Accuracy.BinHigh	TTYPE _n	N	TIME_HI
Spectrum.Data.TimeAxis.Resolution	TTYPE _n	N	TIME_RES
Spectrum.Data.TimeAxis.Accuracy.StatError	TTYPE _n	N	TIME_ERR
Spectrum.Data.TimeAxis.Accuracy.StatErrLow	TTYPE _n	N	TIME_ELO
Spectrum.Data.TimeAxis.Accuracy.StatErrHigh	TTYPE _n	N	TIME_EHI
Spectrum.Data.TimeAxis.Accuracy.SysError	TTYPE _n	N	TIME_SYE
Spectrum.Data.BackgroundModel.Value	TTYPE _n	N	BGFLUX
Spectrum.Data.BackgroundModel.Unit	TUNIT _n	S	(same as FluxAxis)
Spectrum.Data.BackgroundModel.ucd	TUCD _n	N	(same as FluxAxis)
Spectrum.Data.BackgroundModel.Accuracy.StatErrLow	TTYPE _n	N	BG_ELO
Spectrum.Data.BackgroundModel.Accuracy.StatErrHigh	TTYPE _n	N	BG_EHI
Spectrum.Data.BackgroundModel.Accuracy.SysError	TTYPE _n	N	BG_SYE
Spectrum.Data.BackgroundModel.Quality	TTYPE _n	N	BGQUAL

9.2 Expressing the spectrum spatial coordinates in FITS

FITS has a sophisticated mechanism for expressing celestial coordinates. However, it applies only to image axes or table columns. If you want to express a single celestial position in the header of a FITS binary table, the WCS conventions do not apply. You could add an extra pair of columns to the table giving the same position in each row, but that would be wasteful.

Here we propose a local convention leveraging the existing WCS conventions:

- The keyword names for the coordinates are those used in the first four characters of the CTYPE values for the WCS paper: e.g. RA, DEC, GLON, GLAT.
- The coordinate system is identified by the keyword SKY_UCD with values such as pos.eq, etc.
- The RADECSYS and EQUINOX keywords should be used when appropriate.
- Values are always in degrees.

9.3 The SPECSYS keyword

We note the allowed values of the SPECSYS keyword from Greisen et al and the corresponding values from the VO STC:

FITS	STC	Meaning
TOPOCENT	TOPOCENTER	Topocenter
GEOCENTR	GEOCENTER	Geocenter
BARYCENT	BARYCENTER	Solar System Barycenter
HELIOCEN	HELIOCENTER	Heliocenter
LSRK	LSRK	Kinematic local standard of rest
LSRD	LSRD	Dynamic local standard of rest
GALACTOC	GALACTIC_CENTER	Galactic center
LOCALGRP	LOCAL_GROUP_CENTER	Local group barycenter
CMBDIPOL	-	Frame of the Cosmic Microwave Background dipole
SOURCE	-	Source rest frame

9.4 An instance example

We summarize this with a sample FITS extension header.

```
XTENSION= 'BINTABLE'           / binary table extension
BITPIX  =                8 / 8-bit bytes
NAXIS   =                2 / 2-dimensional binary table
NAXIS1  =            57344 / width of table in bytes
NAXIS2  =                1 / number of rows in table
PCOUNT  =                0 / size of special data area
GCOUNT  =                1 / one data group (required keyword)
TFIELDS =                7 / number of fields in each row
EXTNAME = 'SPECTRUM '         / name of this binary table extension
VOCLASS = 'Spectrum V1.0'     / VO Data Model
SIZE    =            180     / Segment size
VOSEGT  = 'Spectrum'         / Segment type
VOCSID  = 'MY-ICRS-TOPO'     / Coord sys ID
RADECSYS= 'FK5 '            / Not default - usually ICRS
EQUINOX = 2.0000000000000E+03 / default
TIMESYS = 'TT '             / Time system
MJDREF  = 0.0               / MJD zero point for times
SPECSYS = 'TOPOCENT'        / Wavelengths are as observed
VOPUB   = 'Cfa Archive'      / VO Publisher authority
VOREF   = '2006ApJ...999...99X' / Bibcode for citation
VOPUBID = 'ivo://cfa.harvard.edu' / VO Publisher ID URI
VOVER   = '1.0'             / VO Curation version
CONTACT = 'Jonathan McDowell, Cfa' /
EMAIL   = 'jcm@cfa.harvard.edu' /
VORIGHTS= 'public' /
VODATE  = '2004-08-30' /
DS_IDPUB= 'ivo://cfa.harvard.edu/spec/10304' / Publisher DID for dataset
COMMENT DS_IDPUB usually the same as DS_IDENT?
OBJECT  = 'ARP 220 '         / Source name
OBJDESC = 'Merging galaxy Arp 220' / Source desc
SRCCLASS= 'Galaxy' /
SPECTYPE= 'ULIRG' /
REDSHIFT=            0.01812 / Emission redshift
RA_TARG  =            233.73791700 / Observer's specified target RA
DEC_TARG  =            23.50333300 / Observer's specified target Dec
TARGVAR  = '1.2' / 20 percent variability amplitude
TITLE    = 'Observations of Merging Galaxies' /
CREATOR  = 'MMT Archive' / VO Creator
COLLECT1= 'Misc Pointed Observations' / Collection
DS_IDENT= 'ivo://cfa.harvard.edu/spec/10304' / Publisher DID for dataset
CR_IDENT= 'MMT4302-102' / Creator internal ID for dataset
DATE     = '2004-08-30T14:18:17' / Date and time of file creation
VERSION  = 2 / Reprocessed 2004 Aug
TELESCOP= 'MMT ' / Telescope [Not part of Spectrum DM]
INSTRUME= 'MMT/BCS ' / Instrument
FILTER   = 'G220 ' / Grating [Not part of Spectrum DM]
CRETYPE  = 'Archival' / Not an on-the-fly dataset
VOLOGO   = 'http://cfa.harvard.edu/vo/cfalogo.jpg' / VO Creator logo
CONTRIB1= 'Jonathan McDowell' / Contributor
CONTRIB2= 'Wilhelm Herschel' / Contributor
CONTRIB3= 'Harlow Shapley' / Contributor
DSSOURCE= 'Pointed' / Survey or pointed, etc
DER_SNR  =            5.0 / Estimate of signal-to-noise
DER_Z    =            0.01845 / Redshift measured in this spectrum
DER_ERR  =            0.00010 / Error in DER_Z
```

```

TIMESDIM= 'T' / Time SIDim
SPECSDIM= '10-10 L' / Spectral SIDim
FLUXSDIM= '10+7 ML-1T-3' / Flux SDim
SYS_ERR = 0.05 / Fractional systematic error in flux
FLUX_CAL= 'Calibrated' /
SPEC_STE= 0.01 / Stat error in spec coord, in SPEC units
SPEC_SYE= 0.001 / Frac sys error in spec coord
SPEC_CAL= 'Calibrated'
SPEC_RES= 5.0 / [angstrom] Spectral resolution
SPECBAND= 'Optical' / SED.Bandpass
SPEC_VAL= 4100.0 / [angstrom] Characteristic spec coord
SPEC_BW = 1800.0 / [angstrom] Width of spectrum
SPEC_FIL= 1.0 / No gaps between channels
TIMEUNIT= 's' / Time unit
TIME_CAL = 'Calibrated' /
DATE-OBS= '2004-06-03T21:18:17' / Date and time of observation
EXPOSURE = 1500.015 / [s] Effective exposure time
TSTART = 52984.301203 / [d] MJD
TSTOP = 52984.318564 / [d] MJD
SKY_CAL = 'Calibrated' /
SKY_RES = 1.0 / [arcsec] Spatial.Resolution
RA = 233.73791 / [deg] Pointing position
DEC = 23.50333 / [deg] Pointing position
RA_NOM = 233.73791 / [deg] Nominal position - not in model
DEC_NOM = 23.50333 / [deg] Nominal position - not in model
APERTURE= 2.0 / [arcsec] Aperture diameter/Slit width
TIME = 52984.301203 / [d] MJD

```

```

COMMENT -----
COMMENT WCS Paper 3 Keywords
1S4_1 = 'WAVE' / Column name with spectral coord
1CTYP4 = 'WAVE-TAB' / Spectral coord is WAVE
1S5_1 = 'WAVE' / Column name with spectral coord
1CTYP5 = 'WAVE-TAB' / Spectral coord is WAVE
1S6_1 = 'WAVE' / Column name with spectral coord
1CTYP6 = 'WAVE-TAB' / Spectral coord is WAVE
1S7_1 = 'WAVE' / Column name with spectral coord
1CTYP7 = 'WAVE-TAB' / Spectral coord is WAVE
COMMENT -----
TTYPE1 = 'WAVE' / Wavelength
TFORM1 = '180E'
TUNIT1 = 'angstrom'
TUCD1 = 'em.wl' /
TUTYP1 = 'Spectrum.Data.SpectralAxis.Value'
TTYPE2 = 'WAVE_LO' /
TFORM2 = '180E'
TUNIT2 = 'angstrom'
TTYPE3 = 'WAVE_HI' /
TFORM3 = '180E'
TUNIT3 = 'angstrom'
TTYPE4 = 'FLUX' /
TFORM4 = '180E'
TUNIT4 = 'erg cm**(-2) s**(-1) angstrom**(-1)'
TUCD4 = 'phot.fluDens;em.wl' / Type of Y axis: F-lambda
TTYPE5 = 'ERR_LO' /
TFORM5 = '180E'
TUNIT5 = 'erg cm**(-2) s**(-1) angstrom**(-1)'
TTYPE6 = 'ERR_HI' /

```

```
TFORM6 = '180E'  
TUNIT6 = 'erg cm**(-2) s**(-1) angstrom**(-1)'  
TTYPE7 = 'QUALITY' /  
TFORM7 = '180I'
```

The data would look like

```
WAVE  WAVE_LO WAVE_HI FLUX  ERR_LO  ERR_HI  QUALITY  
3200.0 3195.0 3205.0 1.48E-12 2.0E-14 2.0E-14 0  
3210.0 3205.0 3215.0 1.52E-12 3.0E-14 3.0E-14 0  
3220.0 3215.0 3225.0 0.38E-12 0.38E-12 0.0 0  
3230.0 3225.0 3235.0 1.62E-12 3.0E-14 3.0E-14 0  
...  
5000.0 4995.0 5005.0 1.33E-11 3.0E-13 3.0E-13 1
```