

# CONSTELLATION X-RAY OBSERVATORY

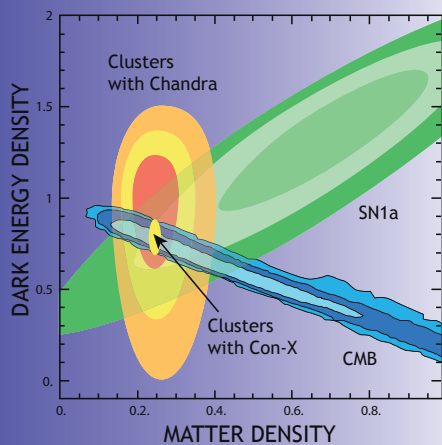
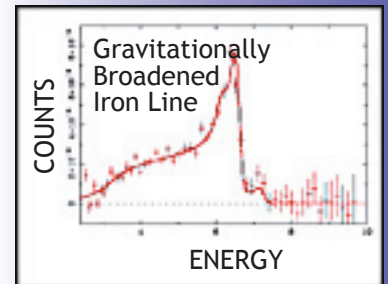
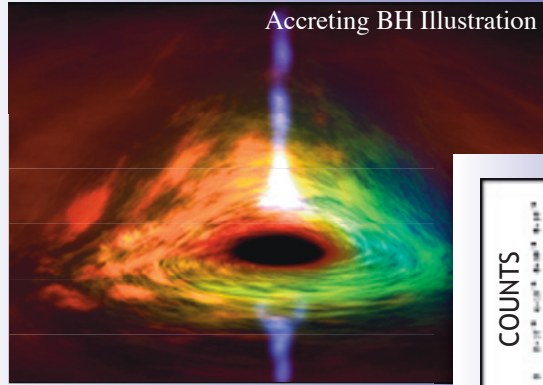
## A BEYOND EINSTEIN GREAT OBSERVATORY

TO UNDERSTAND THE GREAT MYSTERIES OF SPACE, TIME AND ENERGY

### What happens to space, time, and matter at the edge of a black hole?

- Test limits of General Relativity in extreme gravity at black hole event horizon
- Measure spin of black holes
- Make first “movies” of matter spiraling into a black hole

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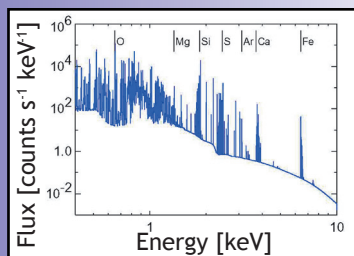
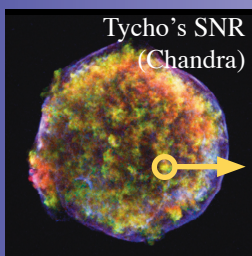
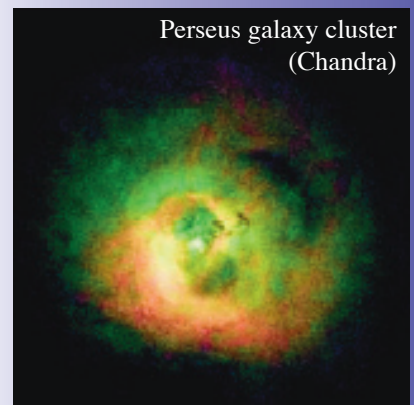
### What is the Universe made of?

- Observe clusters and growth of structure to measure expansion of Universe and determine properties of Dark Energy (DE)
- Measure Dark Matter (DM) content of galaxies, groups, and clusters.
- X-ray observations of clusters see the 4% normal matter, and constrain the nature of the 26% DM and 70% DE.

### How did the Universe come to look like it does now?

- Understand connection between growth of supermassive black holes and host galaxies. Self-regulating accretion alternating with outflows from nuclear black holes?
- Study superwind feedback mechanism in starburst galaxies

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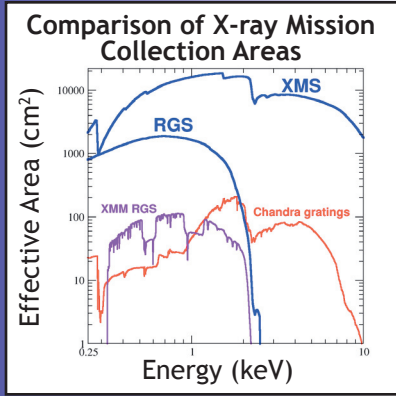
### What is the origin of atoms in stars, planets, and living organisms?

- Trace formation of individual elements in SN explosions
- Probe formation of stars and planets
- Determine nature of superdense matter in neutron stars

The 2000 NRC Decadal Survey (Astronomy and Astrophysics in the New Millennium) **ranked Con-X next in priority after JWST** among major space-based initiatives. These Decadal Survey priorities were re-affirmed by a 2005 NRC Mid-Course Review. Con-X has also received strong endorsement from the Quarks to Cosmos NRC report.



# CONSTELLATION-X: MISSION AND TECHNOLOGY



## MISSION OVERVIEW

- L2 orbit for high viewing efficiency and stable thermal environment
- 5 year lifetime with 10 year goal
- Technically ready, well understood, mission with simple spacecraft

Con-X effective area vs energy compared to Chandra and XMM spectrometers. 100-fold increase in collecting area provides for breakthrough science.

## KEY REQUIREMENTS

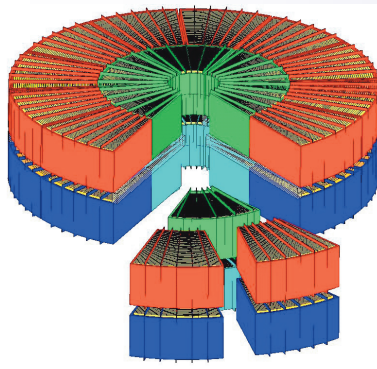
**Bandpass:** 0.25 to 40keV  
**Area:** 15000cm<sup>2</sup>@1.25keV,  
 6000cm<sup>2</sup>@6 keV,  
 1500cm<sup>2</sup>@40 keV

## Spatial Resolution:

15" HPD, 5" goal, for SXT  
 1' HPD, 30" goal, for HXT

## Spectral Revolving Power:

>300 from 0.25keV to 6keV  
 1500 from 6keV to 10keV

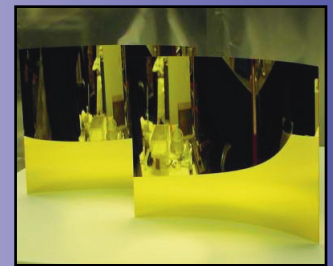


## MIRRORS

- Spectroscopy X-ray Telescope (SXT) mirror technologies derive directly from flight programs (XMM-Newton, Suzaku) but with improved figure accuracy and reduced mass.
- Assembled from 60 (30) degree wedges into circular mirror. Reflection gratings behind SXT.
- Con-X uses thermally slumped glass coated with gold.
- Hard X-ray Telescope (HXT) mirrors use similar approach with multi-layer coatings to extend to higher energies.

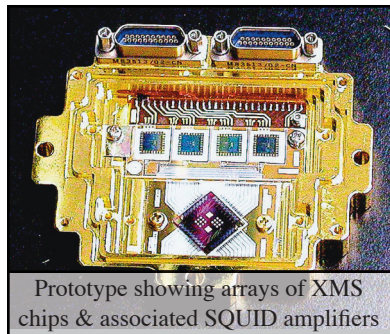
*Left:* A thermally formed glass segment being lifted off a mandrel.

*Right:* Two slumped glass segments coated with the gold reflective surface.

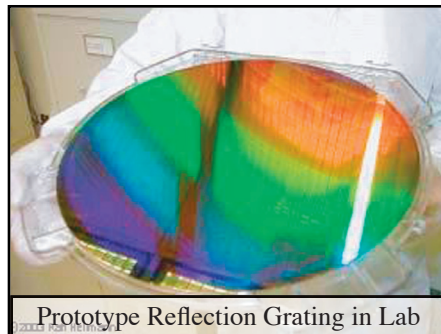


## SCIENCE INSTRUMENTS

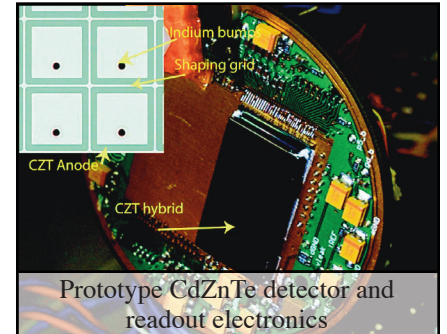
**X-ray Microcalorimeter Spectrometer (XMS).** Provides imaging, non-dispersive, high resolution spectroscopy at SXT focus. Superconducting device, closed loop cyro-cooler, no stored (expendable) cryogen.



**Reflection Grating Spectrometer (RGS).** Provides high resolution, dispersive spectroscopy at lower energies. Attached directly behind part of SXT. Readout with CCDs



**Hard X-ray Telescope (HXT)** mirrors uses segments or full shells with multi-layer coatings to extend to higher energies. Provides imaging, moderate resolution spectroscopy at highest energies using hybrid CdZnTe arrays.



X-ray spectroscopy now rivals the optical for breadth and depth of science. The technologies needed for Con-X are well understood and performance has been demonstrated.

<http://constellation.gsfc.nasa.gov>

