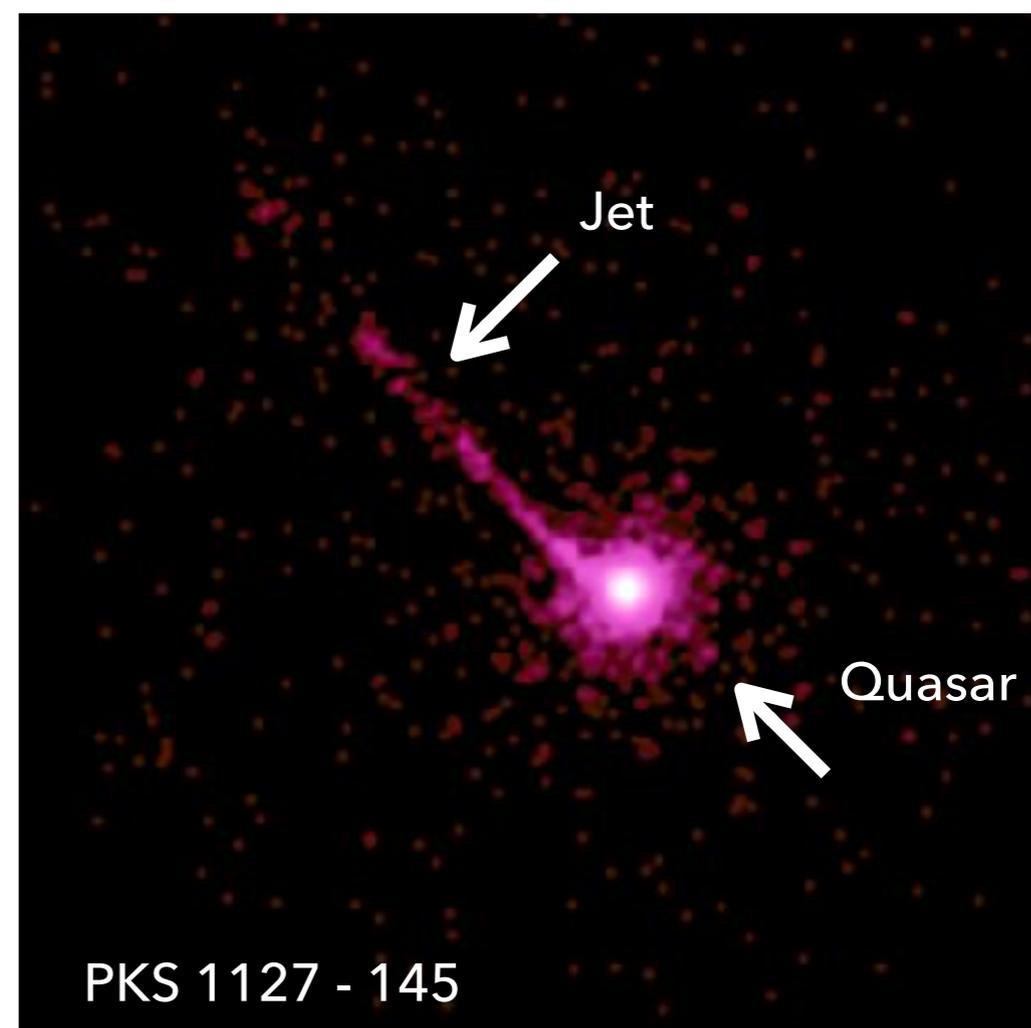


KATY MCKEOUGH & SHIHAO YANG

**DEFINING REGIONS THAT CONTAIN X-RAY JETS
IN HIGH REDSHIFT QUASARS**

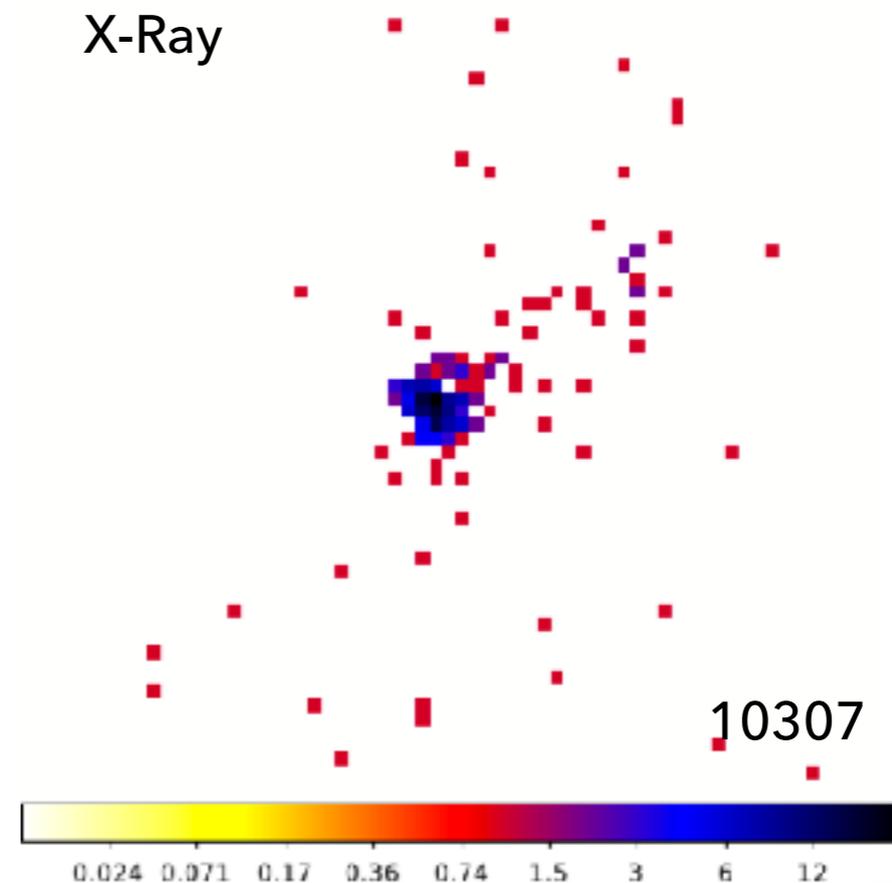
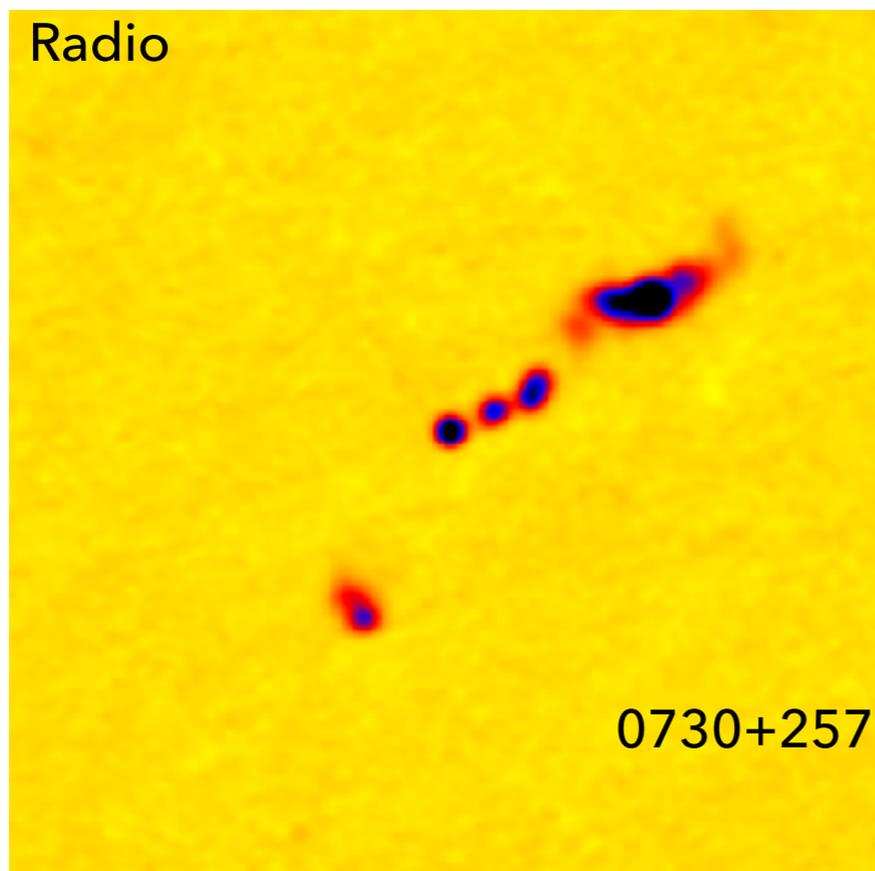
SCIENTIFIC MOTIVATION

- ▶ We are interested in defining an outline around extragalactic jets coming from quasars at high redshift ($z > 2.1$) in X-ray images
- ▶ Defining this boundary is important for accurate luminosity and flux calculations.
- ▶ Detecting jets is difficult because they are diffuse sources (no edges).
- ▶ Images of high redshift jets are of low resolution and few X-ray photons



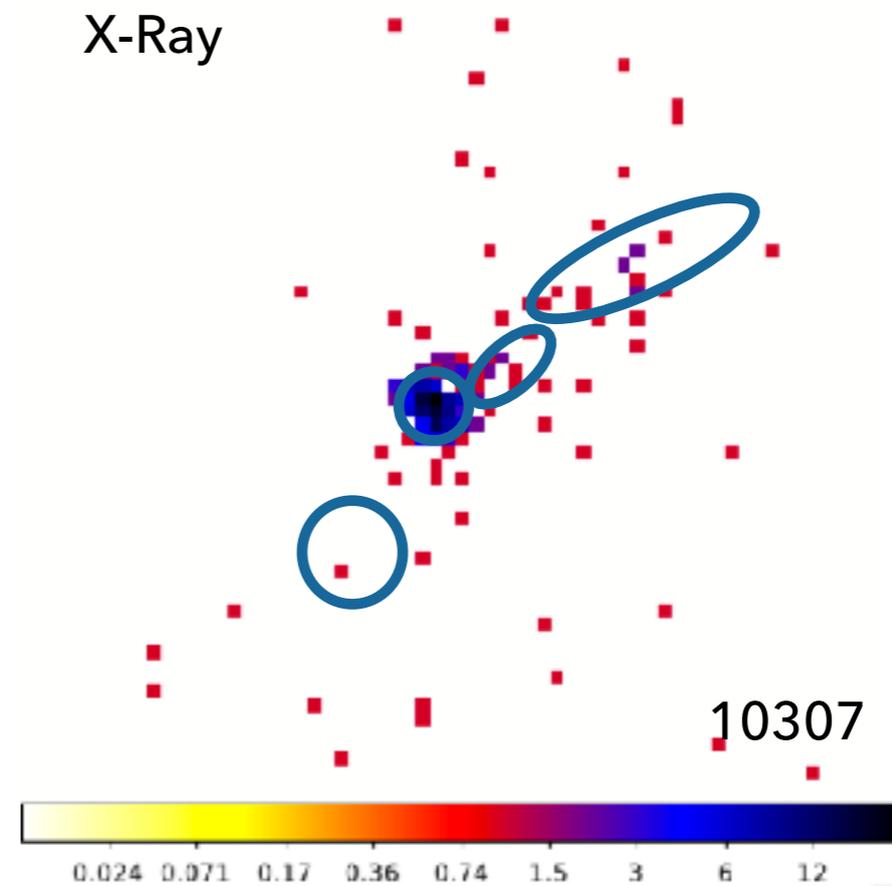
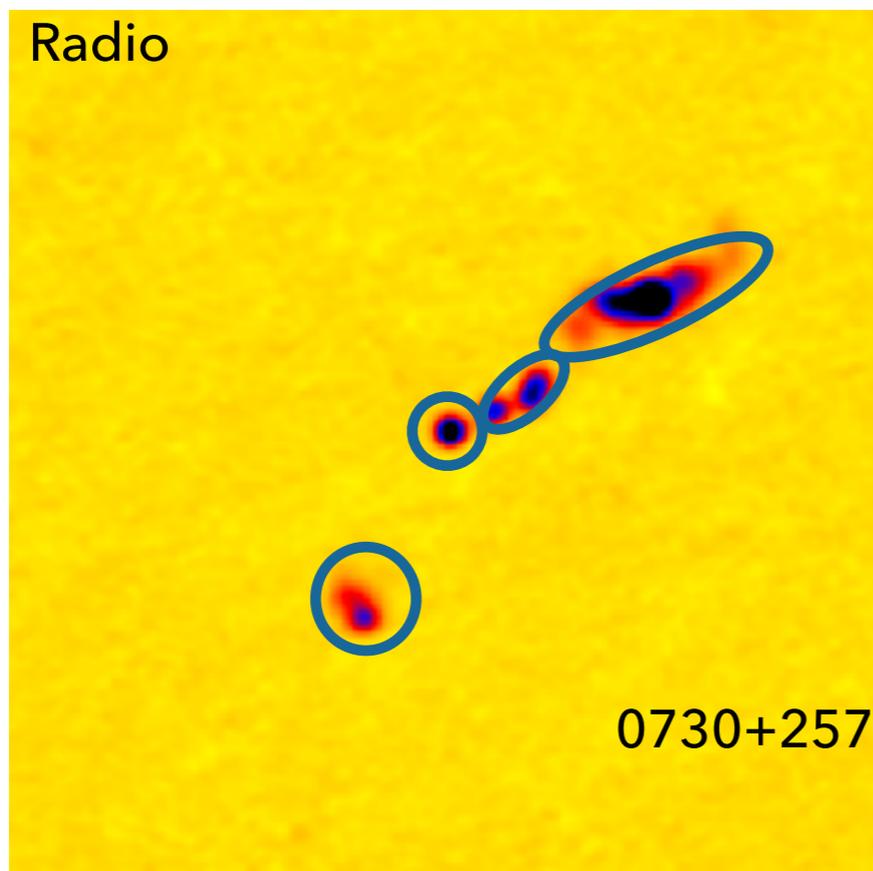
OBSERVATIONAL DATA

- ▶ Chandra X-ray Observatory - ACIS
- ▶ 64 x 64 or 128 x 128 pixel image centered on quasar
- ▶ High and intermediate redshift ($2.10 < z < 4.72$)



REGION OF INTEREST

- ▶ **Region of Interest (ROI)** - region containing the jet or a partition of the jet (e.g. node or lobe)
- ▶ Previous work tests whether or not a jet exists in a predefined ROI (McKeough et al. 2016, Stein et al. 2015)



REGION OF INTEREST

- ▶ Ability to detect jet is sensitive to fit of ROI
- ▶ Issues with previous methods:
 - ▶ Region is defined using radio imaging
 - ▶ Not always available
 - ▶ Not always aligned with X-ray imaging
 - ▶ Region definition relies on human interaction
 - ▶ Inefficient and source of potential error

GOAL

GOAL

Using only the X-ray observation of a quasar and a jet, we are interested in defining an ROI around the jet.

ROADMAP

- ▶ Pre-processing using LIRA
- ▶ Establish model for pixel assignments
- ▶ Model compatibility
- ▶ Draw assignments via Gibbs Sampler
- ▶ Results
- ▶ Future directions

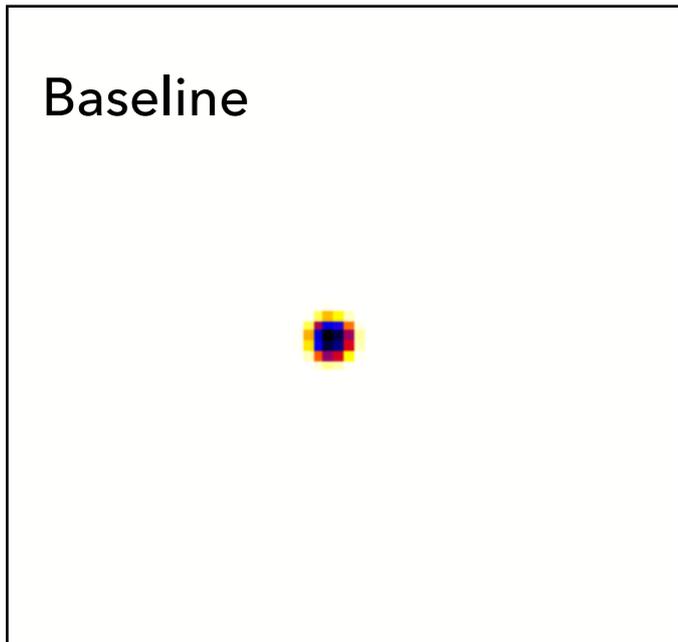
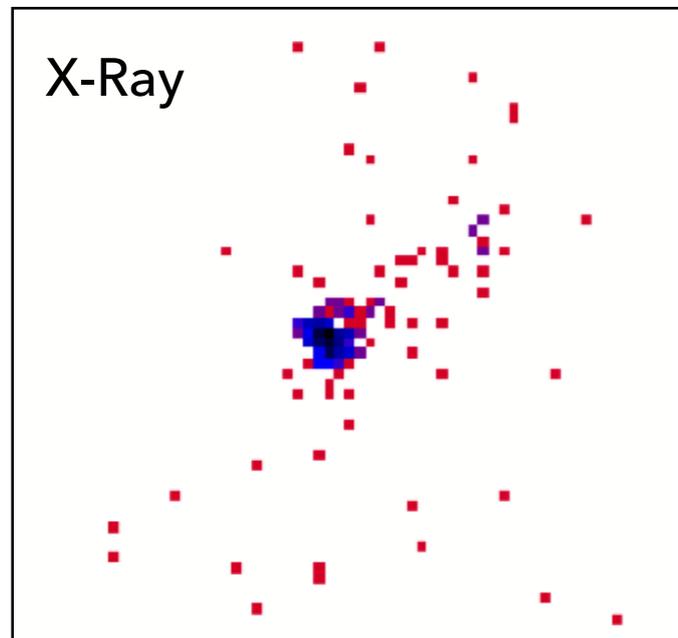
ROADMAP

- ▶ **Pre-processing using LIRA**
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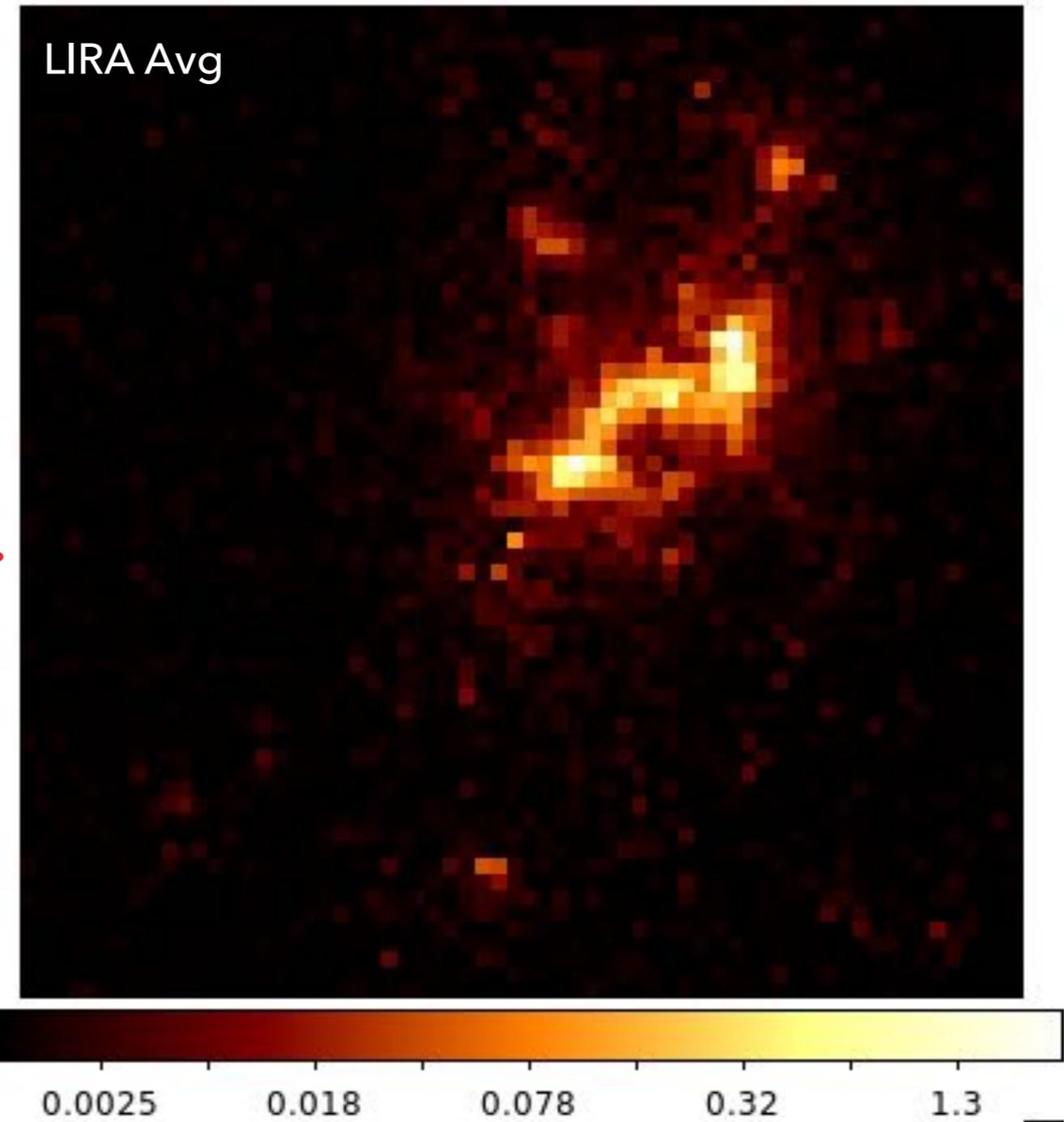
LOW COUNT IMAGE RECONSTRUCTION AND ANALYSIS (LIRA)

- ▶ Esch et al (2004), Connors & van Dyk (2007)
- ▶ Multi-scale Bayesian method
 - ▶ Intensity in “splits” of the image rather than individual pixels
- ▶ Removes quasar & deconvolve Point Spread Function (PSF)
- ▶ Creates posterior for residual pixels as a series of images that capture the emission that is present in excess of the quasar (i.e. the jet)

LOW COUNT IMAGE RECONSTRUCTION AND ANALYSIS (LIRA)



Deconvolve
PSF



ROADMAP

- ▶ Pre-processing using LIRA
- ▶ **Establish model for pixel assignment**
- ▶ Model compatibility
- ▶ Draw assignments via Gibbs Sampler
- ▶ Results
- ▶ Future work

LIKELIHOOD

$$\tilde{\lambda}_{ij}|z, \tau_0, \tau_1, \sigma^2 \sim \text{Log-Normal}(\tau_0, \sigma^2)\mathbb{I}_{z_{ij}=-1} + \text{Log-Normal}(\tau_1, \sigma^2)\mathbb{I}_{z_{ij}=+1}$$

- ▶ We are given observation Y from which we draw the LIRA output:

$$\tilde{\lambda}|Y$$

- ▶ We want to assign each pixel to either the background (-1) or the ROI (1):

$$z = \{-1, +1\}$$

- ▶ Each pixel assignment will have its own average intensity:

$$\tau_-, \tau_+$$

- ▶ For now, the assignees have the same variances:

$$\sigma^2$$

2D ISING PRIOR

$$p(z|\beta) = \frac{\exp(\beta \sum_{ij, i'j' \in |ij-i'j'|=1} z_{ij} z_{i'j'})}{\tilde{Z}(\beta)}$$

▶ Inverse temperature:

 β

▶ Higher β induces more correlation between pixels

▶ Partition function:

 $\tilde{Z}(\beta)$

▶ Estimated via Beale (1996) assuming periodic structure

▶ Commonly used in modeling ferromagnetism.

▶ Induces spatial correlation; adjacent pixels will tend to have the same assignment.

POSTERIOR

$$p(z|\tilde{\lambda}, \tau_0, \tau_1, \sigma^2, \beta) \propto f(\tilde{\lambda}|z, \sigma^2, \tau_0, \tau_1)p(z|\beta)\pi(\sigma^2, \tau_0, \tau_1, \beta)$$

ROADMAP

- ▶ Pre-processing using LIRA
- ▶ Establish model for pixel assignments
- ▶ **Model compatibility**
- ▶ Draw assignments via Gibbs Sampler
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HYPOTHETICAL IDEAL LIRA...

- ▶ Current LIRA output:

$$P(\lambda_{ij}, 1 \leq i, j \leq 64 | Y_{ij}, 1 \leq i, j \leq 64)$$

- ▶ The missing piece of LIRA is the pixel membership indicator:

$$z_{ij} = \{-1, +1\}$$

- ▶ An ideal joint model (denote using subscript \mathcal{J}) would infer λ_{ij} and z_{ij} simultaneously

$$P_{\mathcal{J}}(\lambda, \mathbf{z} | Y) \propto f(Y | \lambda, \mathbf{z}) \pi_{\mathcal{J}}(\lambda, \mathbf{z})$$

OUR APPROACH IS

- ▶ Using LIRA "as is", treating it as pre-processing, trying to infer pixel membership indicator z_{ij} from LIRA output posterior draws of λ_{ij} .
- ▶ Essentially a two-step approach:

- ▶ LIRA (model S_1)

$$P_{S_1}(\lambda|Y) \propto f(Y|\lambda)\pi_{S_1}(\lambda)$$

- ▶ Ising (model S_2) conditional on ONE draw of from S

$$P_{S_2}(z|\tilde{\lambda}) \propto P_{S_2}(\tilde{\lambda}|z)\pi_{S_2}(z)$$

- ▶ So what is this model $S=S_1+S_2$?

$$P_S(\tilde{\lambda}, z|Y) = P_{S_1}(\tilde{\lambda}|Y)P_{S_2}(z|\tilde{\lambda})$$

$$P_S(\lambda, z|Y) \propto f(Y|\lambda)\pi_{S_1}(\lambda) \frac{P_{S_2}(\lambda|z)\pi_{S_2}(z)}{P_{S_2}(\lambda)}$$

THE DIFFERENCE (COMPATIBILITY I)

$$P_{\mathcal{J}}(\lambda, \mathbf{z}|Y) \propto f(Y|\lambda, \mathbf{z})\pi_{\mathcal{J}}(\lambda, \mathbf{z})$$

$$P_S(\lambda, \mathbf{z}|Y) \propto f(Y|\lambda)\pi_{S_1}(\lambda) \frac{P_{S_2}(\lambda|\mathbf{z})\pi_{S_2}(\mathbf{z})}{P_{S_2}(\lambda)}$$

▶ $S=J$? Sufficient condition:

▶ $f(Y|\lambda) = f(Y|\lambda, \mathbf{z})$

▶ $P_{S_2}(\lambda|\mathbf{z})\pi_{S_2}(\mathbf{z}) = \pi_{\mathcal{J}}(\lambda, \mathbf{z})$, which implies
 $\pi_{S_1}(\lambda) = \int \pi_{\mathcal{J}}(\lambda, \mathbf{z}) d\mathbf{z} = \int P_{S_2}(\lambda|\mathbf{z})\pi_{S_2}(\mathbf{z}) d\mathbf{z}$

▶ How far is two-step approach from the ideal output?

▶ Note that $\mathbf{z} \perp\!\!\!\perp Y|\lambda$ for both S and \mathcal{J}

▶ For λ , inference is equivalent

$$P_{\mathcal{J}}(\lambda|Y) = \int P_{\mathcal{J}}(\lambda, \mathbf{z}|Y) d\mathbf{z} \propto f(Y|\lambda) \int \pi_{\mathcal{J}}(\lambda, \mathbf{z}) d\mathbf{z}$$

$$P_S(\lambda|Y) = \int P_S(\lambda, \mathbf{z}|Y) d\mathbf{z} \propto f(Y|\lambda)\pi_{S_1}(\lambda)$$

THE DIFFERENCE (COMPATIBILITY II)

- ▶ For z , calculate K-L divergence:

$$\begin{aligned} & D_{KL}(P_{\mathcal{J}}(\lambda, z|Y) \| P_{\mathcal{S}}(\lambda, z|Y)) \\ &= \int \int P_{\mathcal{J}}(\lambda, z|Y) (\log P_{\mathcal{J}}(\lambda, z|Y) - \log P_{\mathcal{S}}(\lambda, z|Y)) d\lambda dz \\ &= \int \int P_{\mathcal{J}}(\lambda|Y) P_{\mathcal{J}}(z|\lambda, Y) (\log P_{\mathcal{J}}(z|\lambda, Y) - \log P_{\mathcal{S}}(z|\lambda, Y)) d\lambda dz \\ &= \int P_{\mathcal{J}}(\lambda|Y) D_{KL}(P_{\mathcal{J}}(z|\lambda) \| P_{\mathcal{S}}(z|\lambda)) d\lambda \end{aligned}$$

- ▶ so posterior divergence is bounded by prior divergence.

ROADMAP

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STEP 1 – LIKELIHOOD PARAMETERS

- ▶ Drawn through seeded Gibbs sampler via JAGS

- ▶ Priors:

$$\begin{aligned}\tau_-, \tau_+ &\sim N(\mu_0, \sigma^2) \\ \sigma^2 &\sim \text{Inv} - \chi^2(\nu_0, \omega_0^2)\end{aligned}$$

STEP 2 – TEMPERATURE PARAMETER

- ▶ Drawn through Metropolis Hastings

- ▶ Prior:

$$\beta \sim \text{Gamma}(a_\beta, b_\beta)$$

STEP 3- ASSIGNMENTS

- ▶ A well established way to draw the spin state given a specific temperature is **Swendsen & Wang (1987)**.
- ▶ The S-W method takes a spin system $z|\beta$ and induces a bigger system that contains the original N spin variables and M additional bond variables, denoted by d .
- ▶ Define joint distribution that couples spins to bonds:

$$p(z, d|\tilde{\lambda}, \tau_0, \tau_1, \sigma^2, \beta) \propto \prod_{m=1}^M g_m(z_m, d_m|\beta) \prod_{ij} f(\tilde{\lambda}_{ij}|z, \tau_0, \tau_1, \sigma^2)$$

- ▶ Marginal distribution of z is equal to our posterior.

$$\sum_d p(z, d|\tilde{\lambda}, \tau_0, \tau_1, \sigma^2, \beta) = p(z|\tilde{\lambda}, \tau_0, \tau_1, \sigma^2, \beta)$$

- ▶ Conditional distributions are easy to sample from.

$$p(z|d, \beta, -) \quad p(d|z, \beta, -)$$

COUPLING SPINS TO BONDS

- ▶ Bonds can be disconnected (0) or connected (1).

$$d = \{0, 1\}$$

COUPLING SPINS TO BONDS

- ▶ Bonds can be disconnected (0) or connected (1).

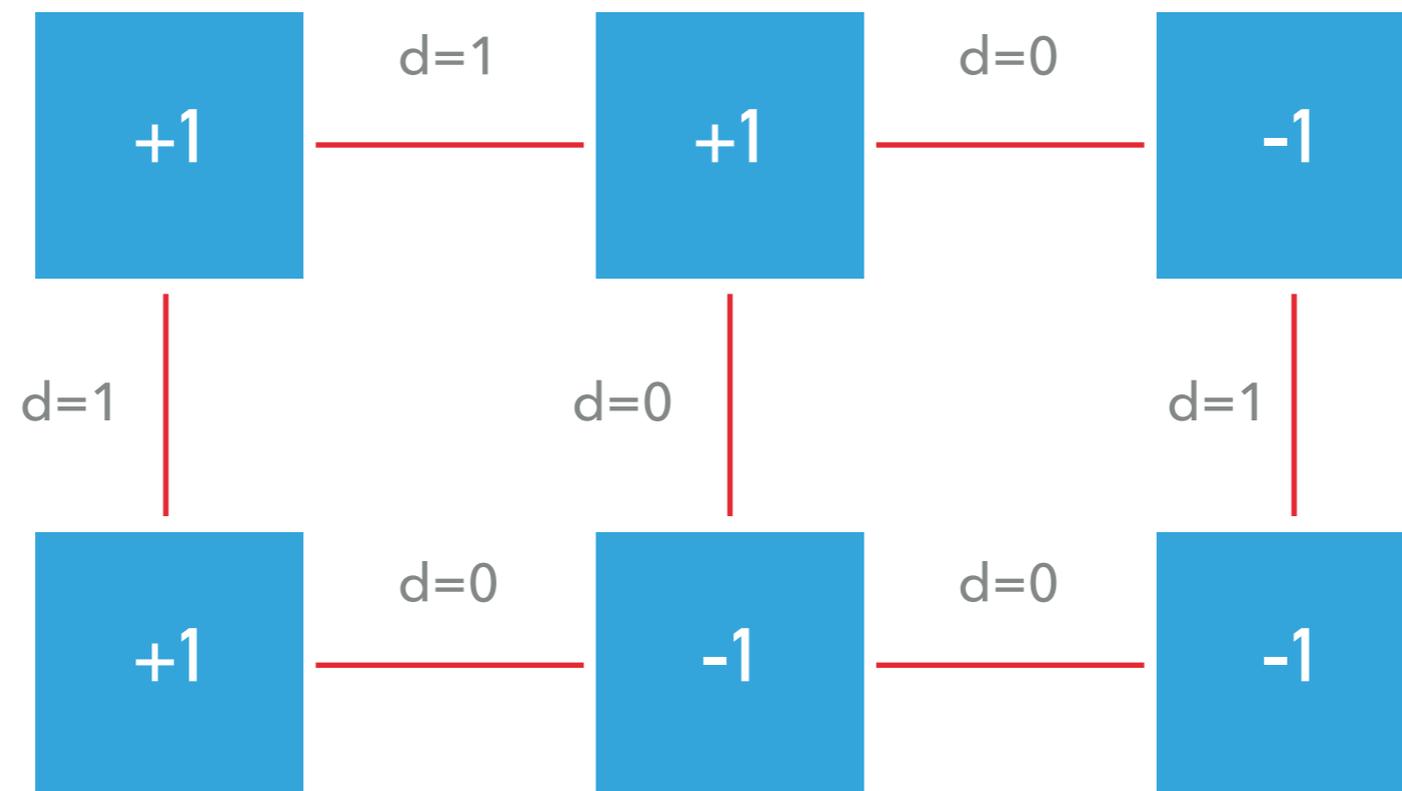
$$d = \{0, 1\}$$

+1	+1	-1
+1	-1	-1

COUPLING SPINS TO BONDS

- ▶ Bonds can be disconnected (0) or connected (1).

$$d = \{0, 1\}$$



COUPLING SPINS TO BONDS

- ▶ Factor coupling bonds and spins is:

$$g_m(z_m, d_m) = \begin{cases} & d_m = 0 & & d_m = 1 \\ z_{ij} = 0 & z_{i'j'} = 0 & z_{i'j'} = 1 & z_{i'j'} = 0 & z_{i'j'} = 1 \\ & e^{-\beta} & e^{-\beta} & e^{\beta} - e^{-\beta} & 0 \\ z_{ij} = 1 & e^{-\beta} & e^{-\beta} & 0 & e^{\beta} - e^{-\beta} \end{cases}$$

- ▶ Rescale by constant factor: $p = 1 - e^{-2\beta}$

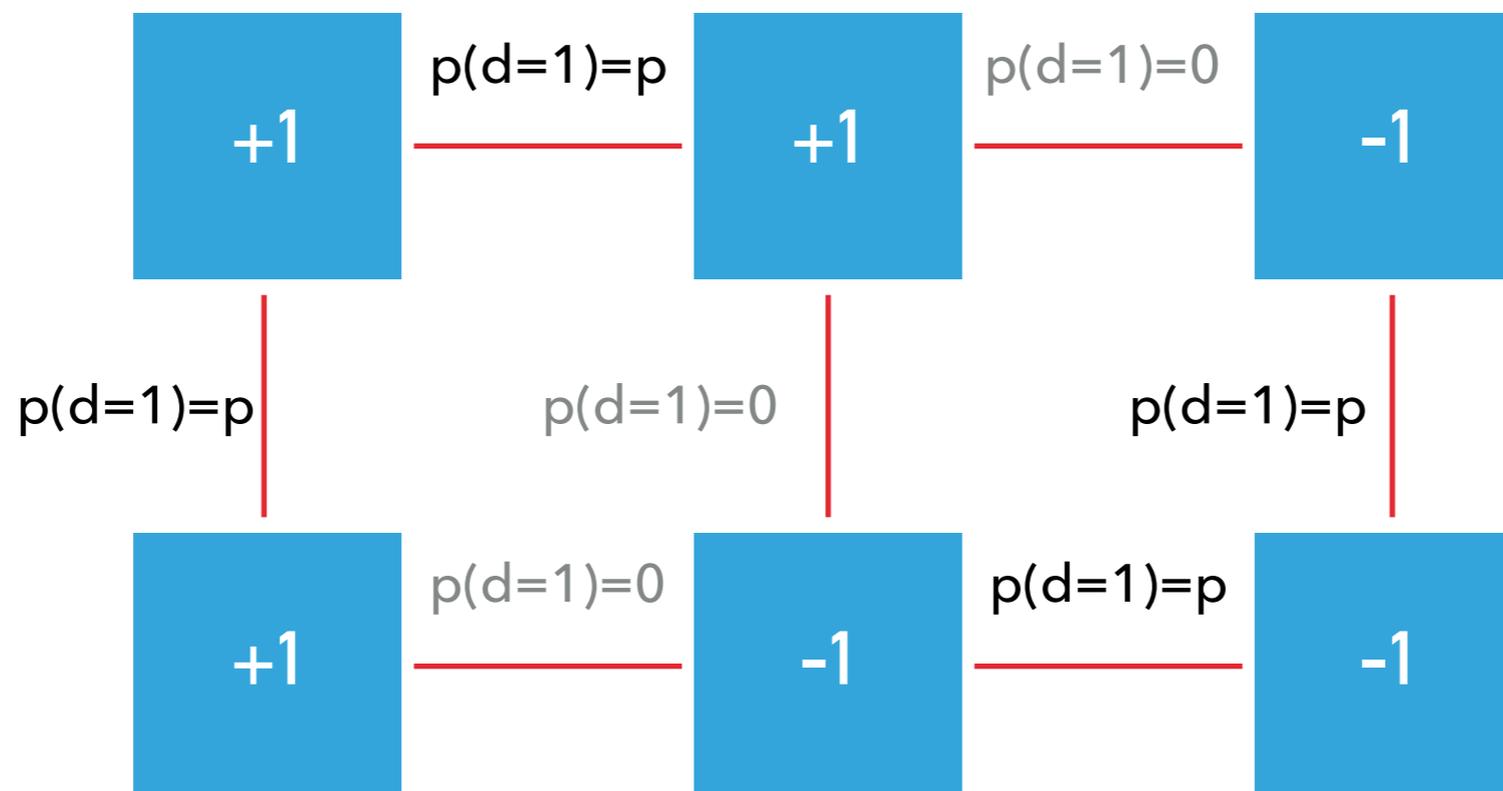
$$\tilde{g}_m(z_m, d_m) = \begin{cases} & d_m = 0 & & d_m = 1 \\ z_{ij} = 0 & z_{i'j'} = 0 & z_{i'j'} = 1 & z_{i'j'} = 0 & z_{i'j'} = 1 \\ & 1 - p & 1 - p & p & 0 \\ z_{ij} = 1 & 1 - p & 1 - p & 0 & p \end{cases}$$

- ▶ Therefore:

$$p(z, d|\beta) \propto \prod_m g_m(z_m, d_m|\beta) \prod_{ij} f(\tilde{\lambda}_{ij}|z_{ij}, b, \tau, \sigma^2) \propto \prod_m \tilde{g}_m(z_m, d_m|\beta) \prod_{ij} f(\tilde{\lambda}_{ij}|z_{ij}, b, \tau, \sigma^2)$$

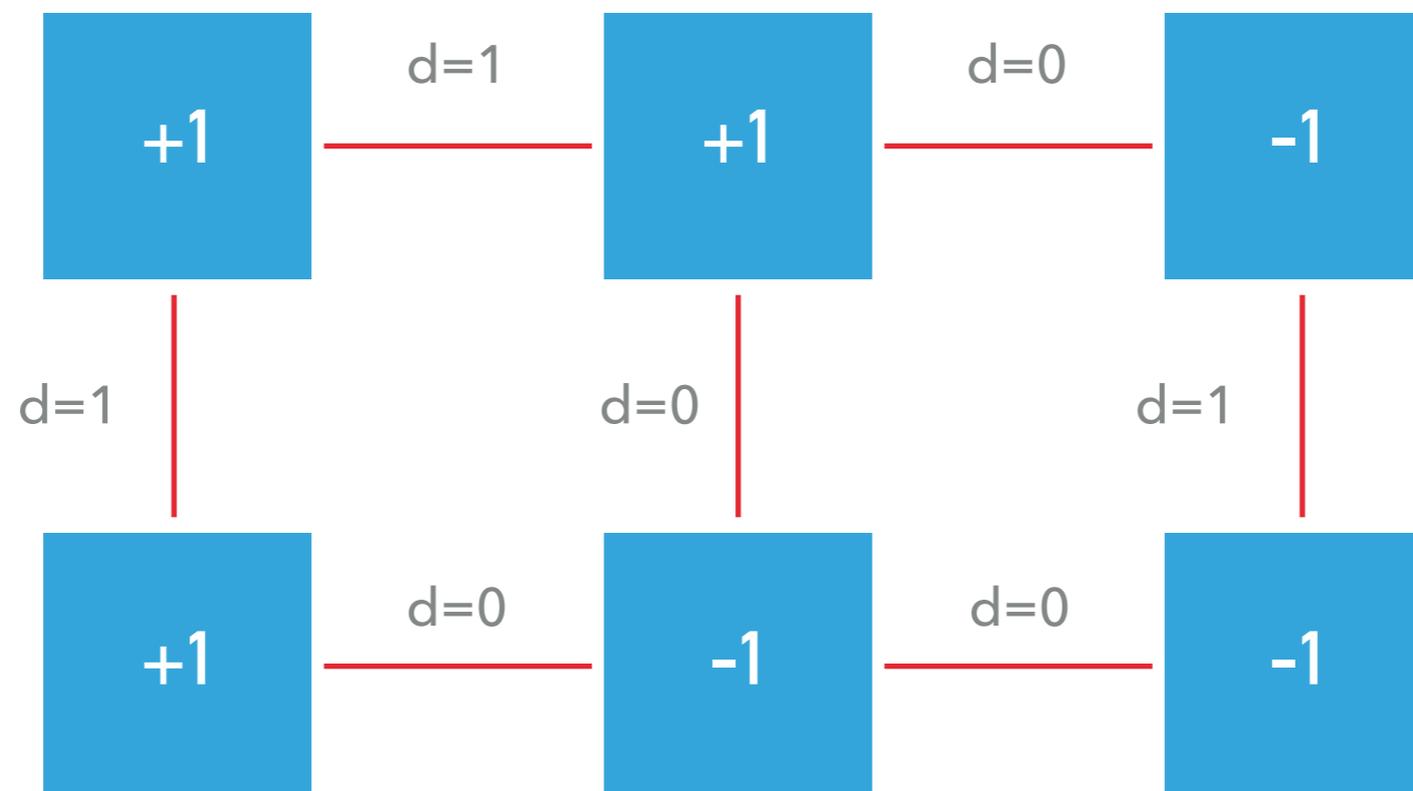
SAMPLING

- ▶ Sample from $p(d|z, \beta)$
 - ▶ If two spins connected to bond are equal, set the bond d_m equal to 1 with probability $p=1-\exp(-2\beta)$, and 0 otherwise.



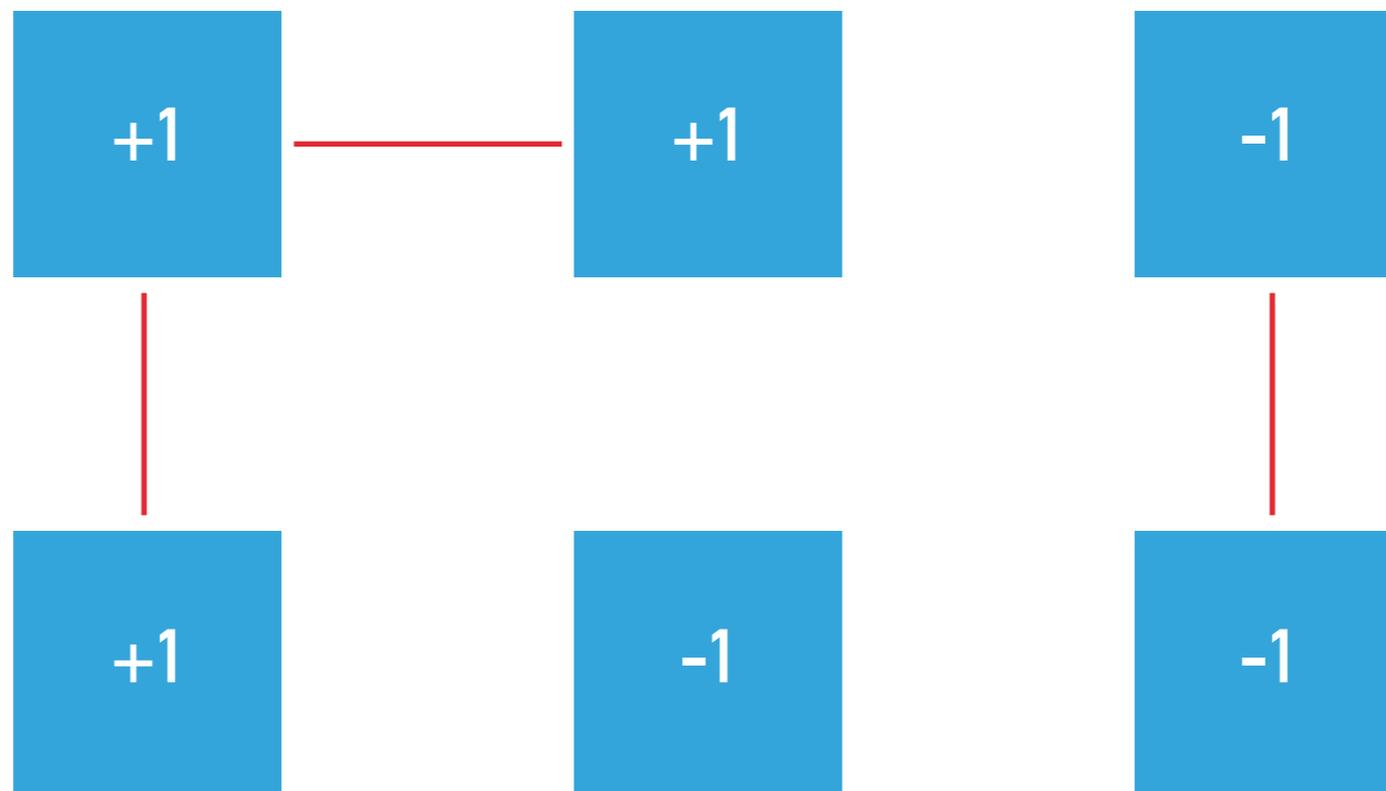
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SAMPLING

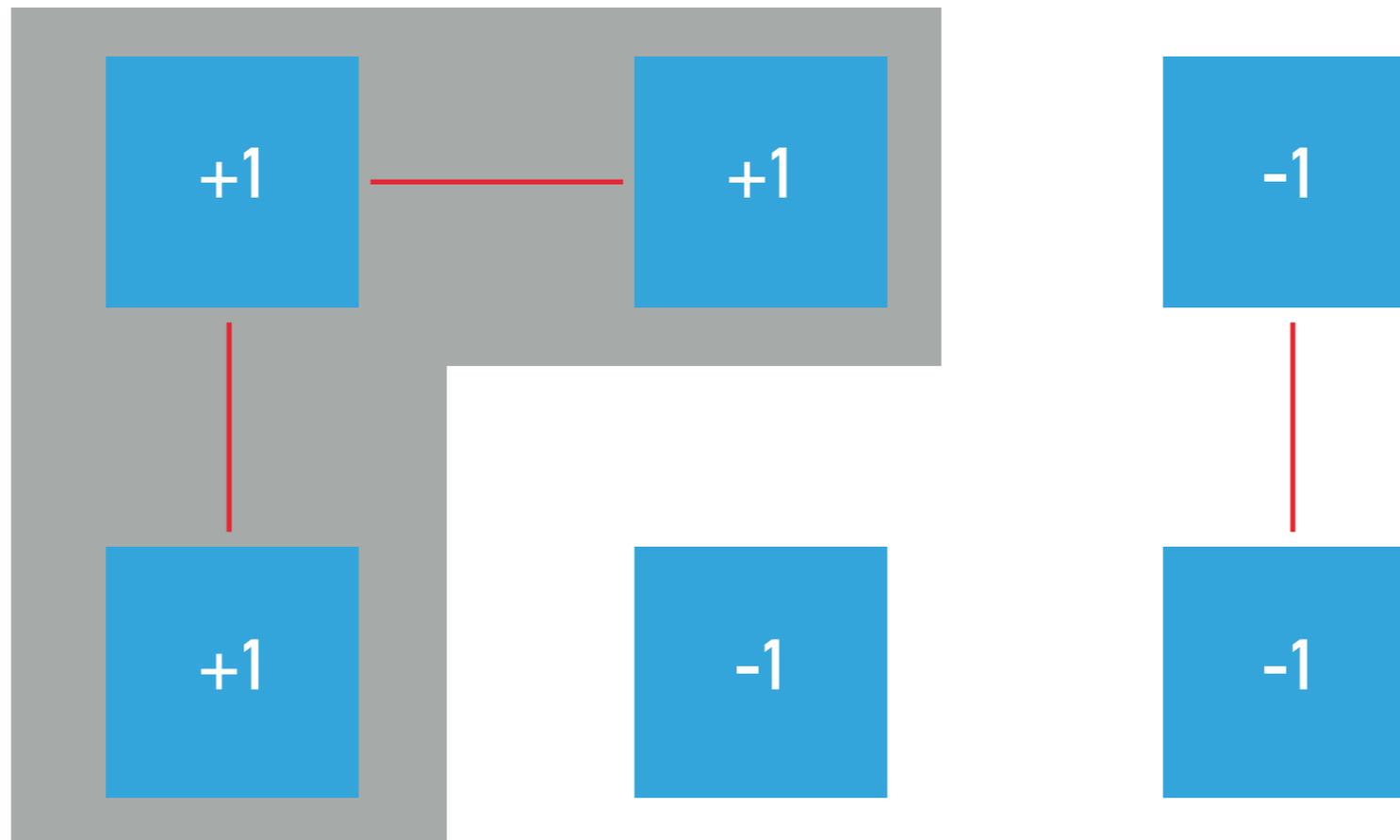
- ▶ Sample from $p(z|d, \beta)$
 - ▶ Bonds connect spins into C cluster.
 - ▶ Cluster - all pixels that are connected by a bond $d_m=1$
 - ▶ Each cluster will take spin $+1$ with probability p_+
 -1 with probability $p_-=1-p_+$

$$\frac{p_+}{p_-} = \frac{\prod_{ij \in C} f(\tilde{\lambda}_{ij} | z_{ij} = +1, \tau_1, \tau_2, \sigma^2)}{\prod_{ij \in C} f(\tilde{\lambda}_{ij} | z_{ij} = -1, \tau_1, \tau_2, \sigma^2)}$$

SAMPLING

- ▶ Sample from $p(z|d, \beta)$

$$p(z=+1) = p_+$$

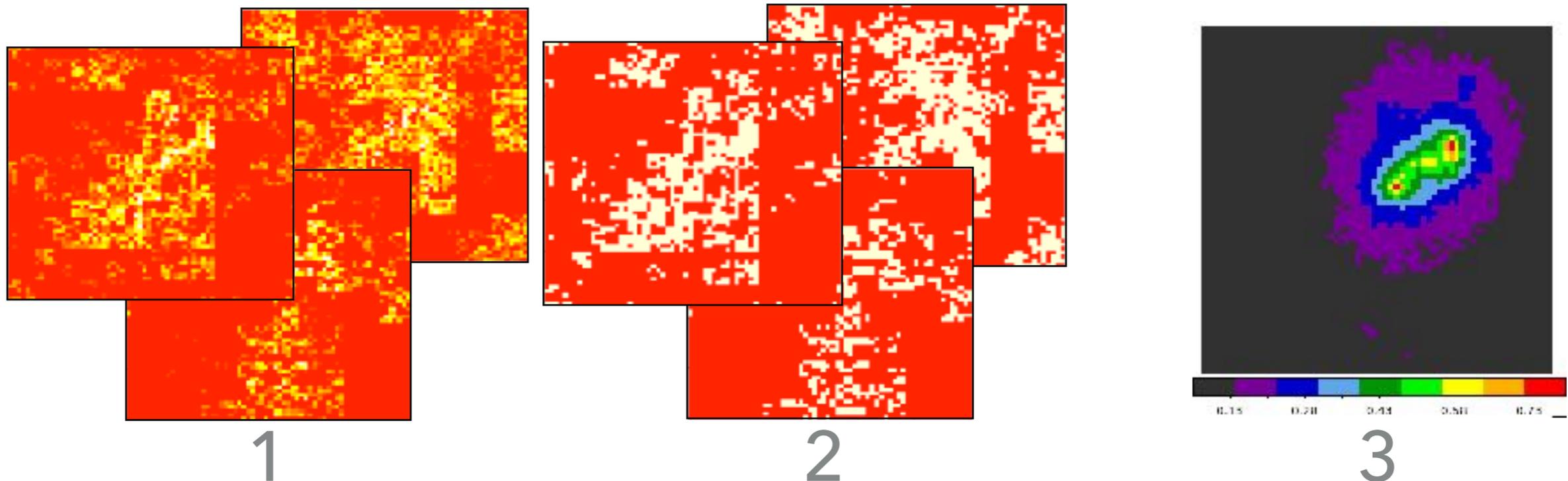


ROADMAP

- ▶ Pre-processing using LIRA
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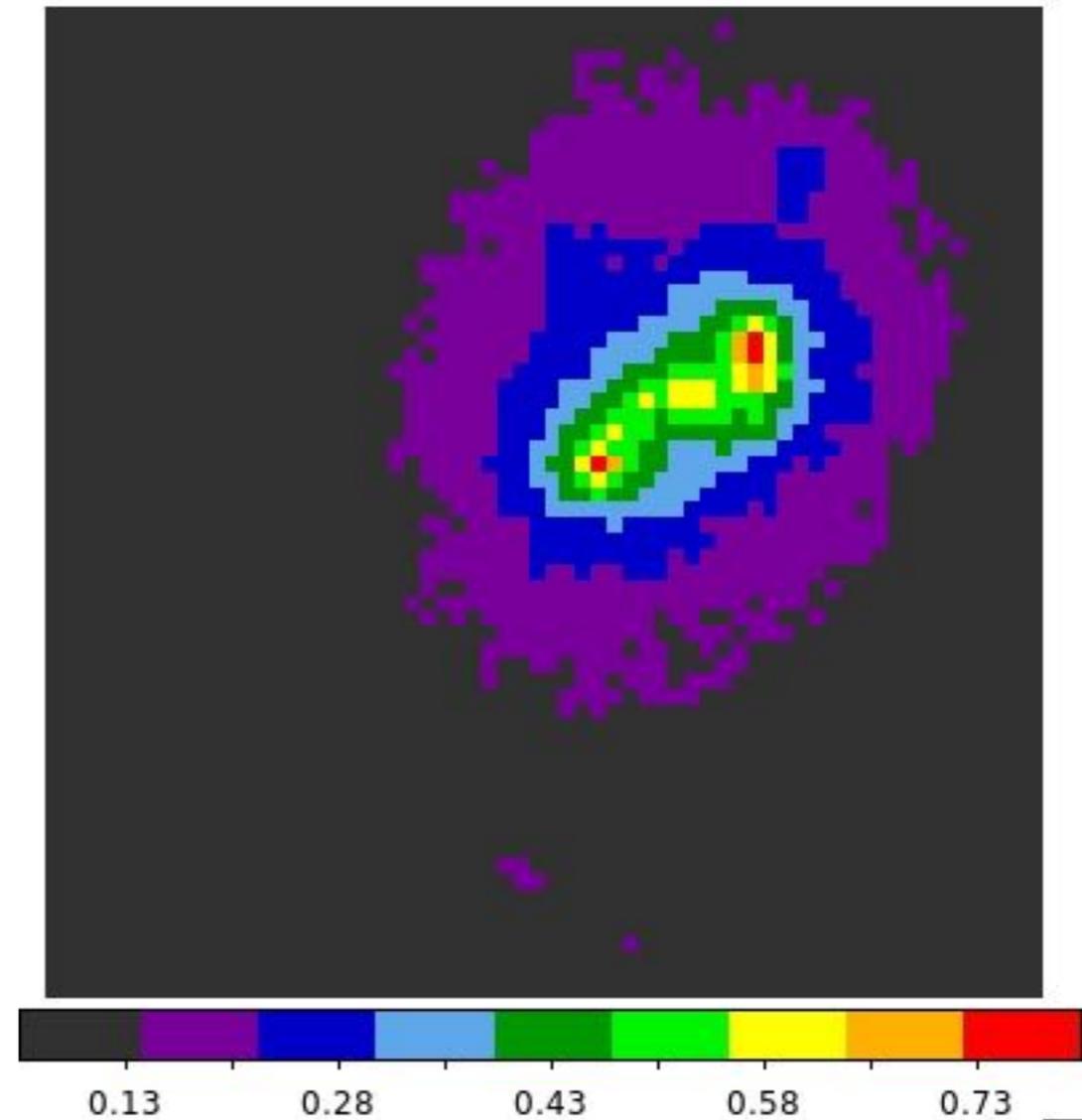
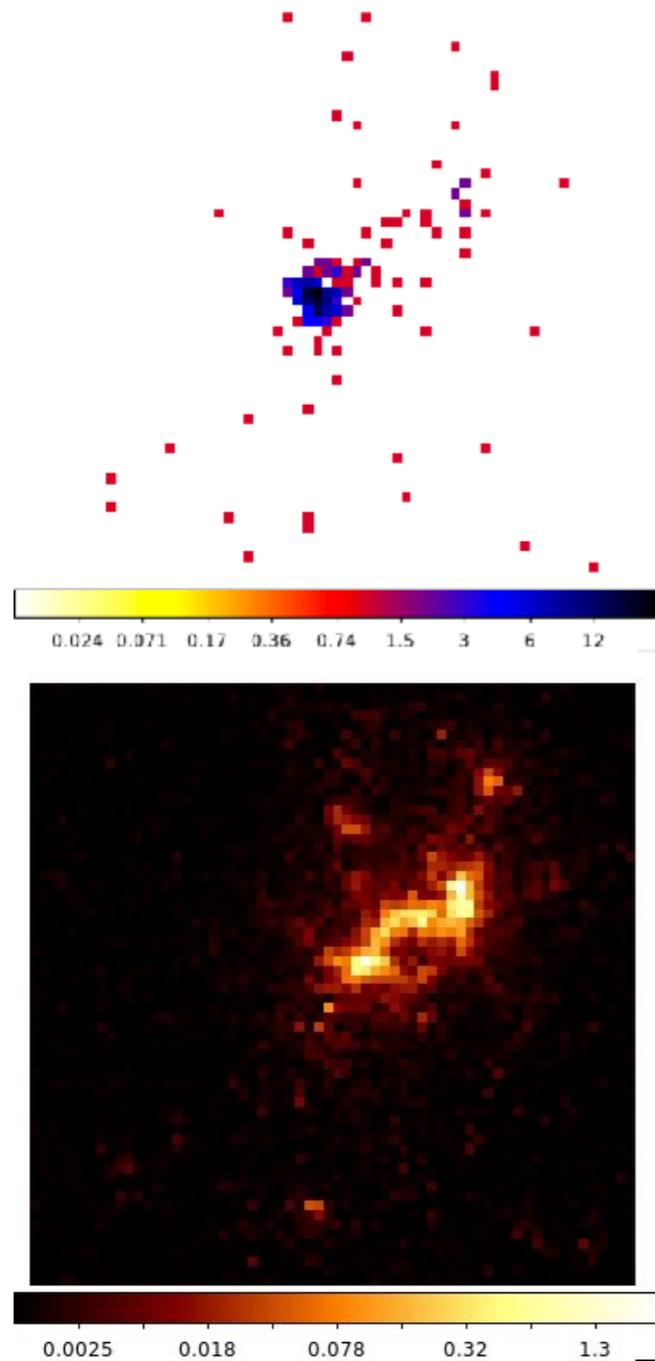
ISING-LIRA ITERATIONS

1. Get many posterior draws from LIRA
2. Apply new method to each LIRA draw
3. Average across LIRA-Ising iterations to get probability map.



PROBABILITY MAP

- ▶ “Probability” each pixel is a member of the ROI:

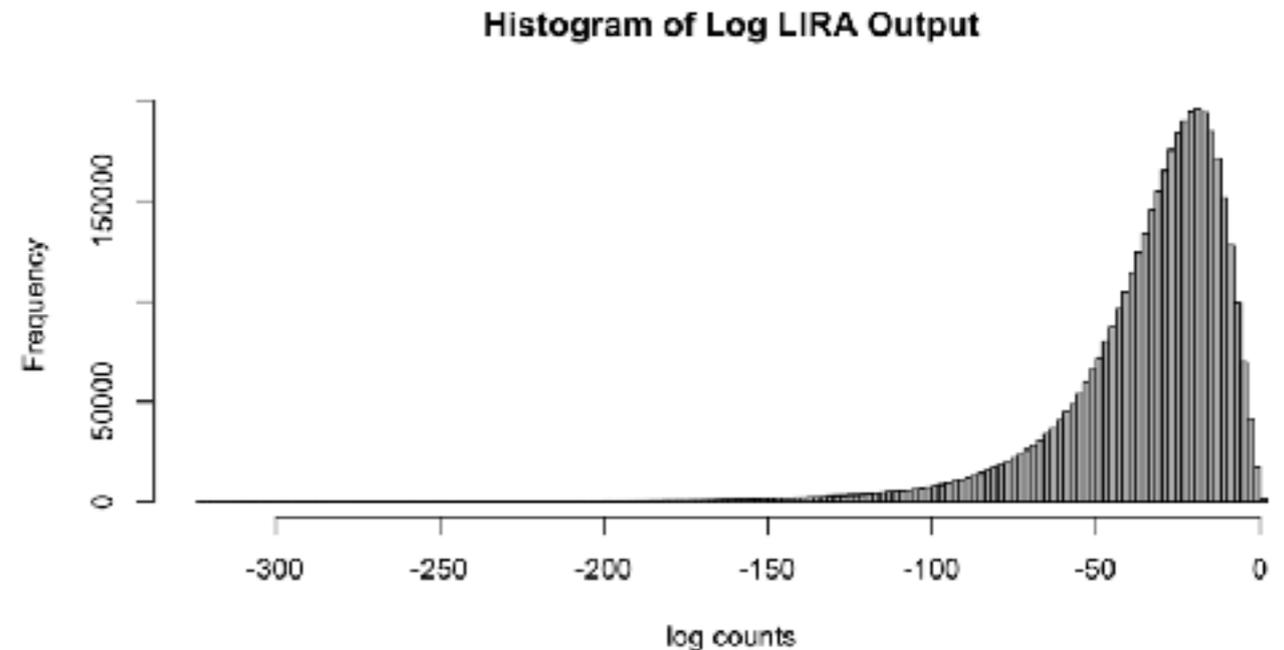


ROADMAP

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- ▶ **Future work**

“UNDERFLOW” ISSUE

- ▶ Majority of LIRA output are very small counts ranging as low as $<10^{-300}$.
- ▶ This doesn't make sense physically.
- ▶ Our model is sensitive to lower ranges induced by multi-scale structure of LIRA algorithm, which overwhelms signal from jet.
- ▶ One solution: Machine limitation is 10^{-16} and anything smaller is induced by underflow error.
- ▶ After looking at distribution of concatenated pixels from all 1000 iterations, it is clear this is not an underflow issue.
- ▶ Prior for LIRA intensity is Gamma weighted for very, very small values.



DIFFERENT VARIANCES

- ▶ Pixels in the ROI should have a higher variance than the background since we expect all background pixels to be close to zero.

$$\tilde{\lambda}_{ij} | z, \tau_0, \tau_1, \sigma_0^2, \sigma_1^2 \sim \text{Log-Normal}(\tau_0, \sigma_0^2) \mathbb{I}_{z_{ij}=-1} + \text{Log-Normal}(\tau_1, \sigma_1^2) \mathbb{I}_{z_{ij}=+1}$$

HURDLE MODEL

- ▶ Computational limits only produce reliable estimates of the LIRA posterior on the order of 10^{-16}
- ▶ About roughly 70% of the data lies below this
- ▶ Hurdle model accounts for this truncation:

$$f(\lambda_{ij} | z_{ij} = +1, \tau_0, \tau_1, \sigma_0^2, \sigma_1^2) = \begin{cases} \Phi(\log x; \tau_1, \sigma_1^2) & ; \lambda_{ij} = x \\ g(\lambda_{ij} | \tau_1, \sigma_1^2) & ; \lambda_{ij} > x \end{cases}$$

ADJACENT PIXEL DEFINITION

- ▶ Could be modified to the 8 nearest pixels instead of 4.
- ▶ Modified to include pixels beyond just the adjacent pixels
- ▶ Correlation as a function of distance

POTTS MODEL

- ▶ Want to identify multiple partitions of the jet (e.g. nodes)
- ▶ Potts is a more generalized version of the Ising model allows for more than two spin assignments:

$$z_{ij} = \{0, 1, 2, 3, \dots\}$$

REFERENCES

- ▶ McKeough et al., *Detecting Relativistic X-ray Jets in High-Redshift Quasars*, The Astrophysical Journal (2016)
- ▶ Stein et al., *Detecting Unspecified Structure in Low-Count Images*, The Astrophysical Journal (2017)
- ▶ Connors & van Dyk, *How To Win With Non-Gaussian Data: Poisson Goodness-of-Fit*, SCMA IV (2007)
- ▶ Esch et al. , *An Image Restoration Technique with Error Estimates*, The Astrophysical Journal (2004)
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- ▶ Swendsen & Wang, *Nonuniversal Critical Dynamics in Monte Carlo Simulations*, Physical Review Letters (1987)

LIRA

MULTI-SCALE IMAGE REPRESENTATION

- ▶ Stores total intensities and series of four way split proportions such that the product recovers original pixel intensities

- ▶ Pixel Intensity

$$\Lambda = \{\Lambda_i, I = 1 \dots N\}$$

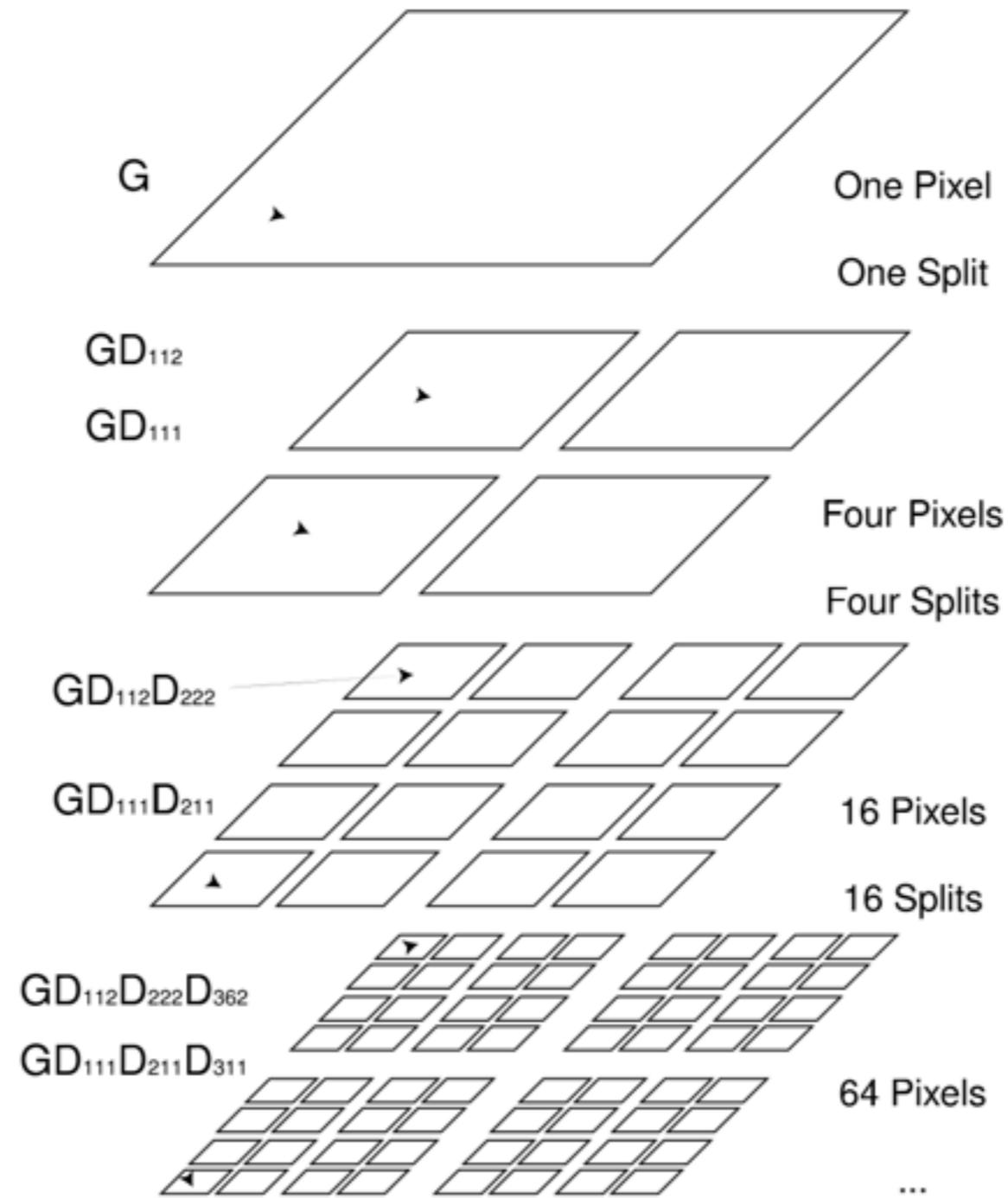
- ▶ Splits

$$D_{k, l_{k(i)}, m_{k(i)}}$$

- ▶ Split proportion at scale k corresponding to group i

$$\Lambda_i = G \prod_{k=1}^K D_{k, l_{k(i)}, m_{k(i)}}$$

MULTI-SCALE IMAGE REPRESENTATION



LIKELIHOOD

- ▶ Probability photon originating in pixel i , is observed in pixel j (*PSF*):

$$P_i = \{P_{ij}, j = 1, \dots, N\}$$

- ▶ Observed pixel counts:

$$Y = \{Y_i, i = 1, \dots, N\}$$

- ▶ Distribution of Y :

$$Y_j | \Lambda, \Lambda^B \underset{\sim}{\sim} \text{Poisson} \left[\left(\sum_{i \in \mathcal{I}} P_{ij} \Lambda_i \right) + \Lambda_j^B \right]$$

- ▶ Suppress background to obtain likelihood:

$$L(\Lambda, \Lambda^B | \mathbf{Y}) \equiv L(\Lambda | \mathbf{Y}) \propto \prod_{j \in \mathcal{I}} p(Y_j | \Lambda).$$

PRIOR

- ▶ Prior on total intensity:

$$G \sim \text{Gamma}(\gamma_0, \gamma_1)$$

- ▶ Prior on splits:

$$\mathbf{D}_{kl} \equiv \{D_{klm}, m = 1, \dots, 4\} \stackrel{d}{\sim} \text{Dirichlet}(\alpha_k, \alpha_k, \alpha_k, \alpha_k) \\ k = 1, \dots, K, \quad l = 1, \dots, 4^{k-1}$$

- ▶ Hyperprior favors smoother image:

$$p(\alpha_k) \propto \exp(-\delta \alpha^3 / 3)$$

CYCLE SPINNING

- ▶ Multiscale format produces checkerboard-like patterns
- ▶ Solution:
 - ▶ Shift center of image randomly before making splits
 - ▶ Splits wrap around edges of image to induce translation invariance