DASCH Variables in the Kepler Field

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Harvard-Smithsonian Center for Astrophysics And the DASCH team:

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Alison Doane - Plate Curator

Jaimie Pepper – Plate Cleaning

Bob Simcoe - Hardware engineer

Sumin Tang – Photometry, variable search Mathieu Servillat – Astrometry

Doug Mink – Astrometry & software

Numerous volunteers – Assist in Plate Cleaning/Scanning

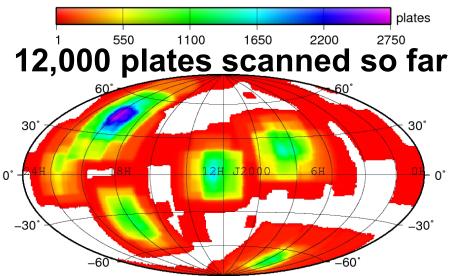
AAS 217th Meeting, Seattle, Jan. 11, 2011

DASCH (Digital Access to a Sky Century @ Harvard) to Digitize and Measure the Harvard Plates to

Open the ~100yr Time Domain Window

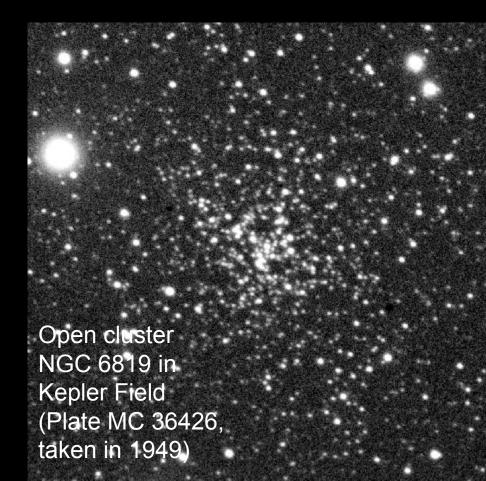
- ~530,000 photographic plates between 1880s-1980s (Grindlay et al. 2009).
- 500-1500 measurements for each object with B~10-14 (up to 18 mag in some regions)
- study temporal variations of a wide variety of objects (stars to AGNs)
- Astrometry: 0.8-3 arcsec (but up to 6" near edge of low-res plates)
 Photometry: 0.1-0.15 mag (Laycock et al. 2010).



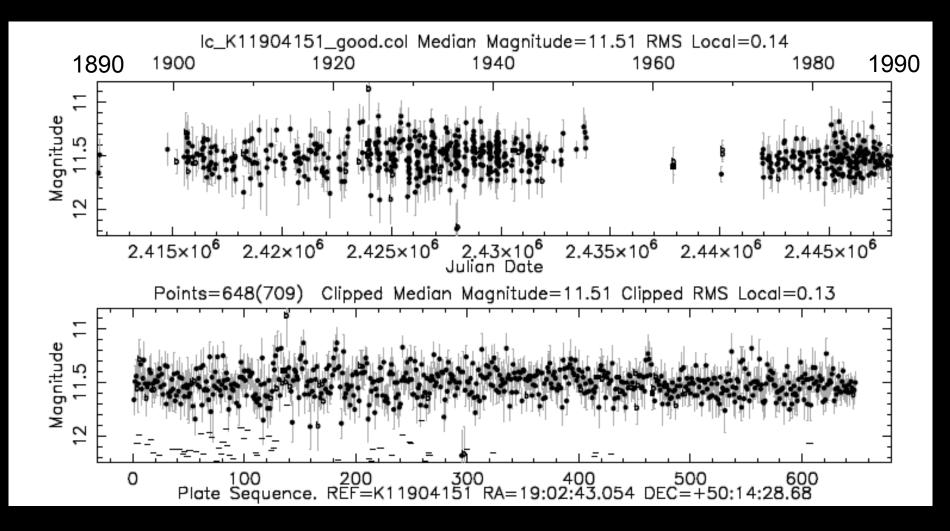


~3000 plates each covering all or part of the Kepler field:

Long-term **DASCH** coverage + short-timescale **Kepler** data.

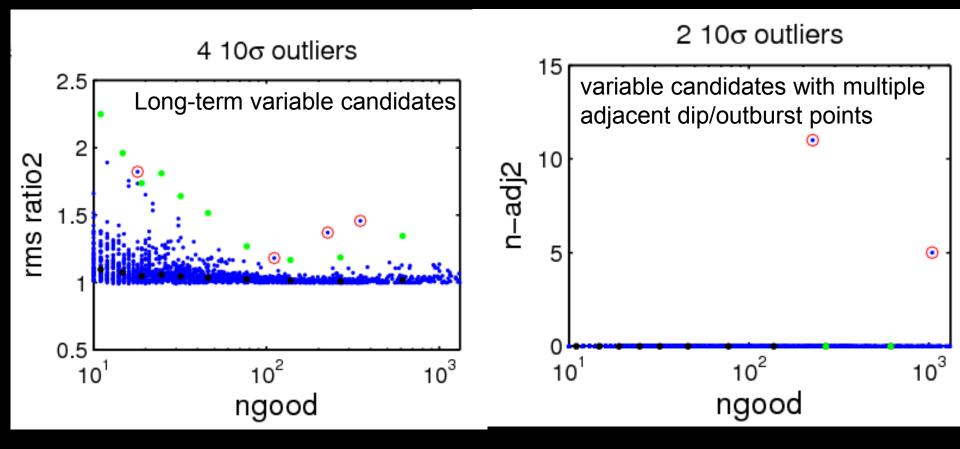


100 yr DASCH lightcurve of Kepler-10



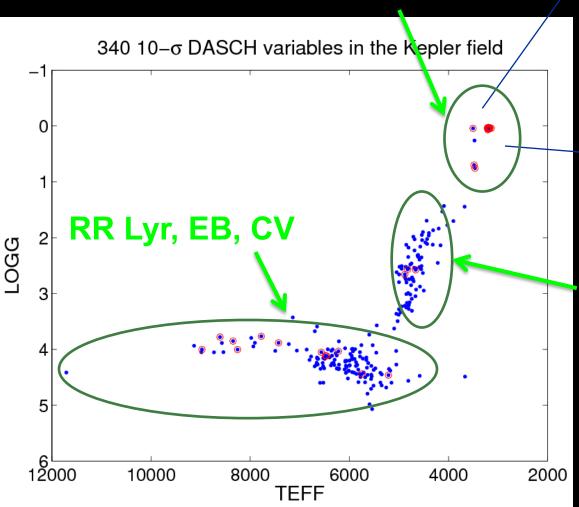
Variable Search:

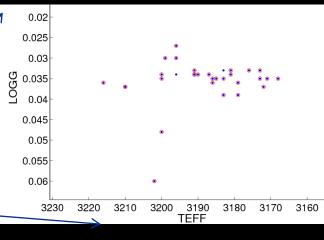
- 1. divided into local 2°x2° bins
- 2. calculate lightcurve statistics for each star
- 3. pick up outliers (variable candidates)
- 4. Visual examination of plate images



Different types of Variable Candidates

Red giant variables: Mira, LPV

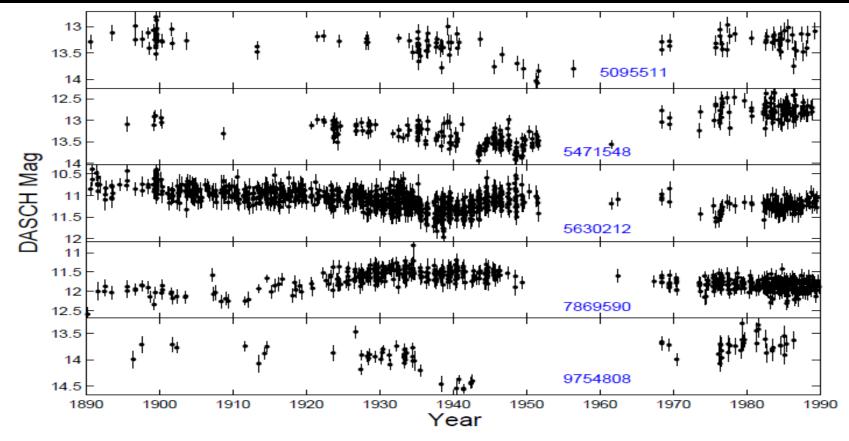




MISC, including some unusal longterm variables

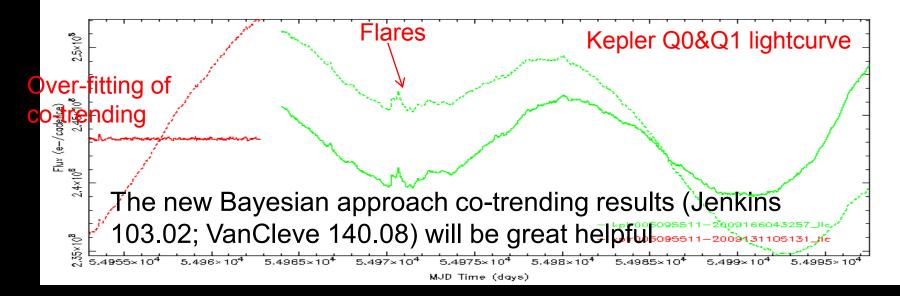
Visual validation still in progress... Only ~30 most interesting candidates have been done so far

K Giants with 10-50 yr variation: An unknown phase of RGB/AGB evolution? similar to what we found in the M44 field (Tang et al. 2010)

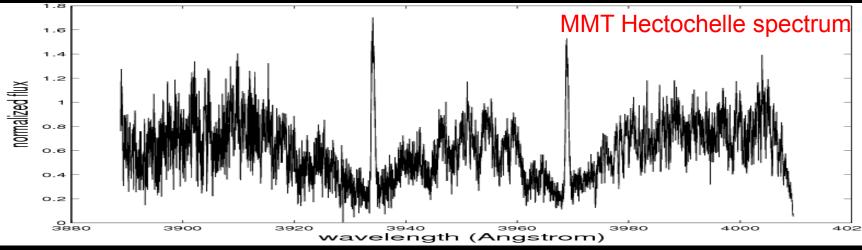


Teff ~4400-4500 K, log g~2-3 cm/s2, R~6-15 R \odot in the KIC 5630212: also a K giant according to its spectra

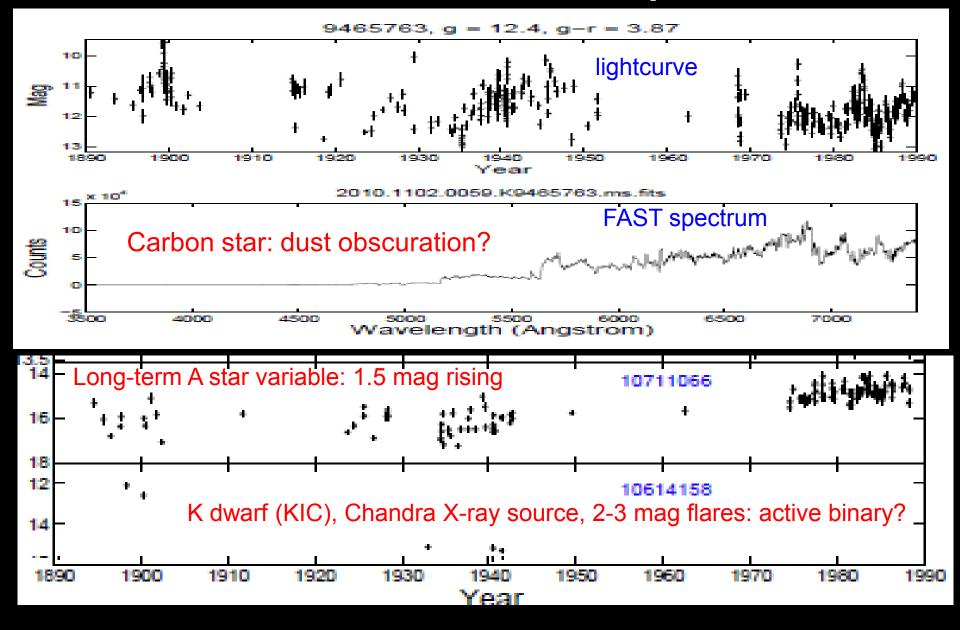
K5095511: a long term variable and active/spotted Giant?



Strong Ca K&H emission lines: indicate strong magnetic activity



A few more examples



Combining DASCH

- 100yr lightcurves for B<15 mag objects in the whole sky
- Discovery of exciting new types of variables
- Explore long-term variability for known variables

with the high-cadence great accuracy Kepler data

- Short-term variability (flares, dimmings, oscillations)
- Magnetic activity, star spots and rotation
- Asteroseismology: derive stellar parameters, pin down their position in the evolutionary track. Compare the stellar parameters derived from atmosphere fitting (MMT Hectochelle/Magellan MIKE spectra).

Enables a unique study of stellar variability on the widest range of timescales. We have applied for the Kepler Cycle 3 GO program for ~20 of our most interesting DASCH variables.