

ChaMPPlane discovery of candidate symbiotic binaries in the Bulge

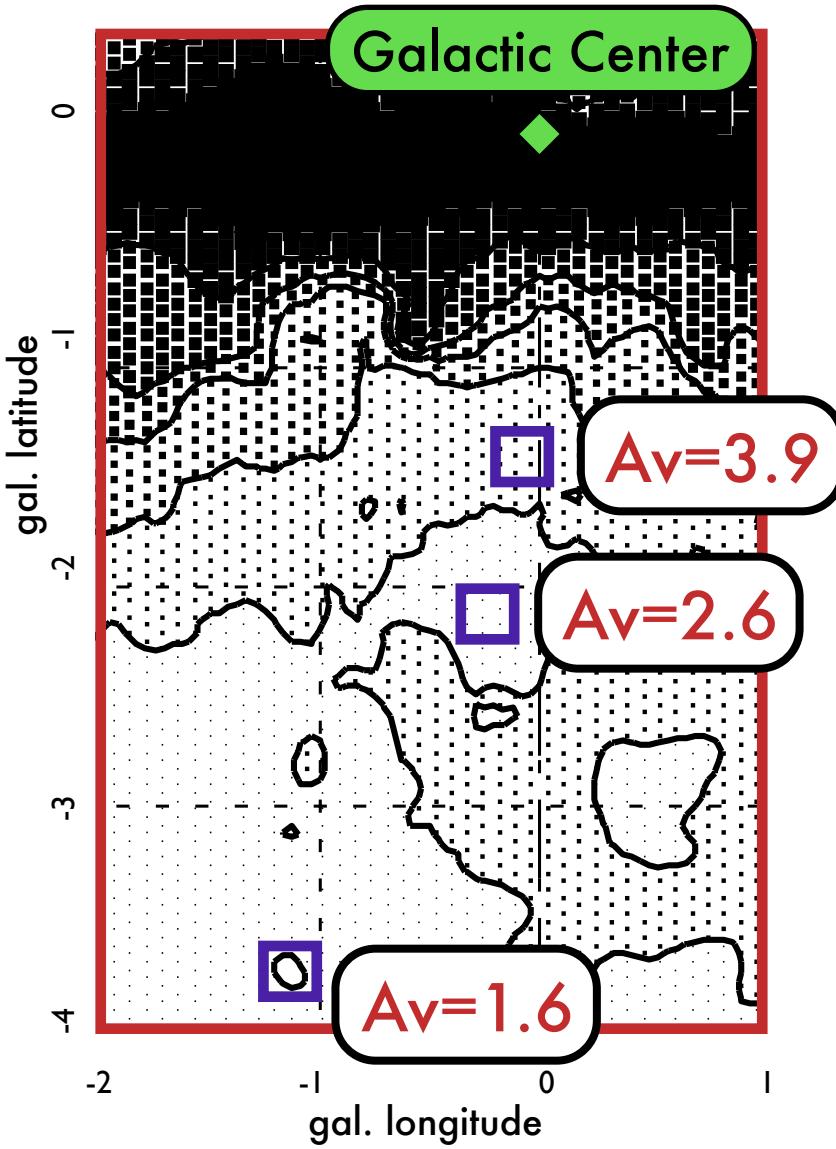
M van den Berg, J Grindlay, S Laycock, J Hong, P Zhao, X Koenig, E Schlegel,
H Cohn, P Lugger, M Rich, A Dupree, G Smith, J Strader

CfA, Univ. Texas, Indiana Univ, UCLA, Lick Obs

new candidate symbiotic binaries (giant + wind-accreting companion)

very hard sources with non-typical “symbiotic” optical spectra
⇒ different class?

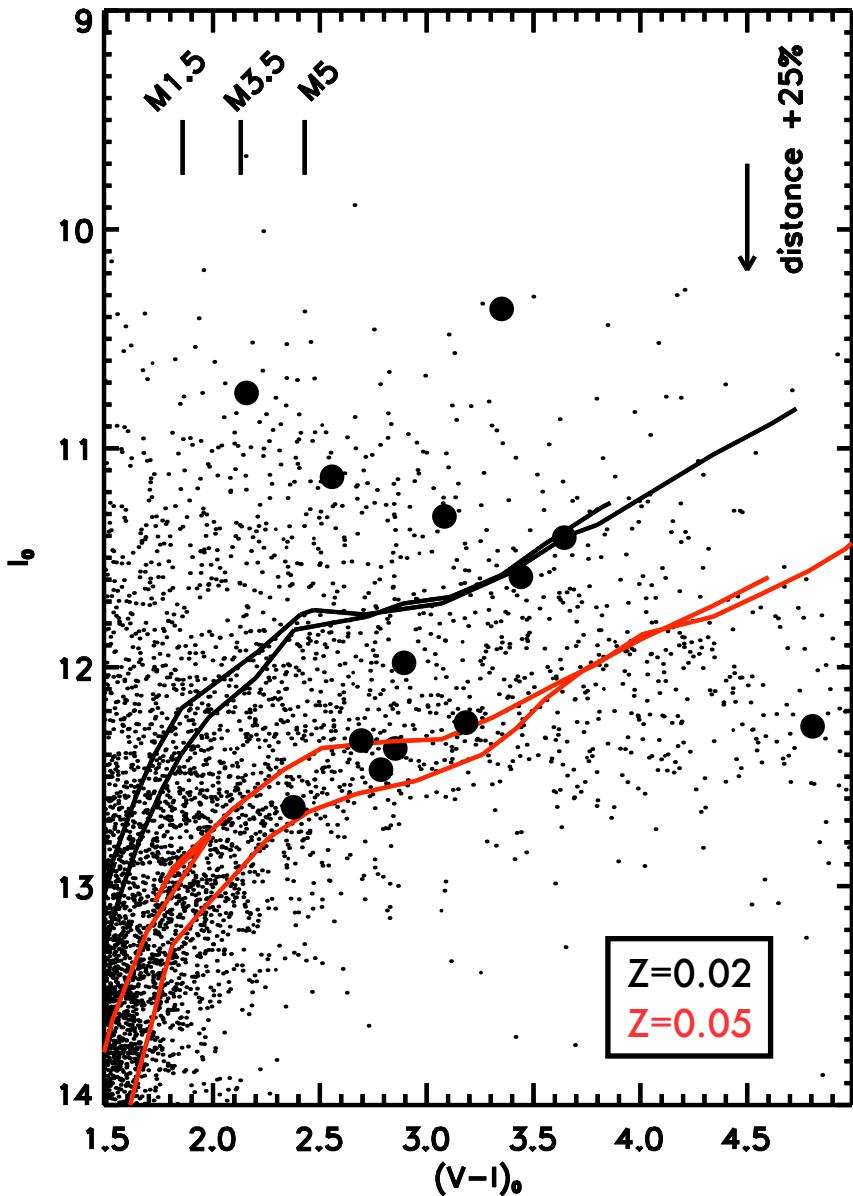
ChaMPPlane Bulge Windows Survey



reddening map (Stanek 1998)

- ❖ GOALS :
Bulge sources
compare with other old populations
- ❖ DATA :
3x100 ks Chandra/ACIS-I
HST/ACS

Optical ids with OGLE-II: M-giants

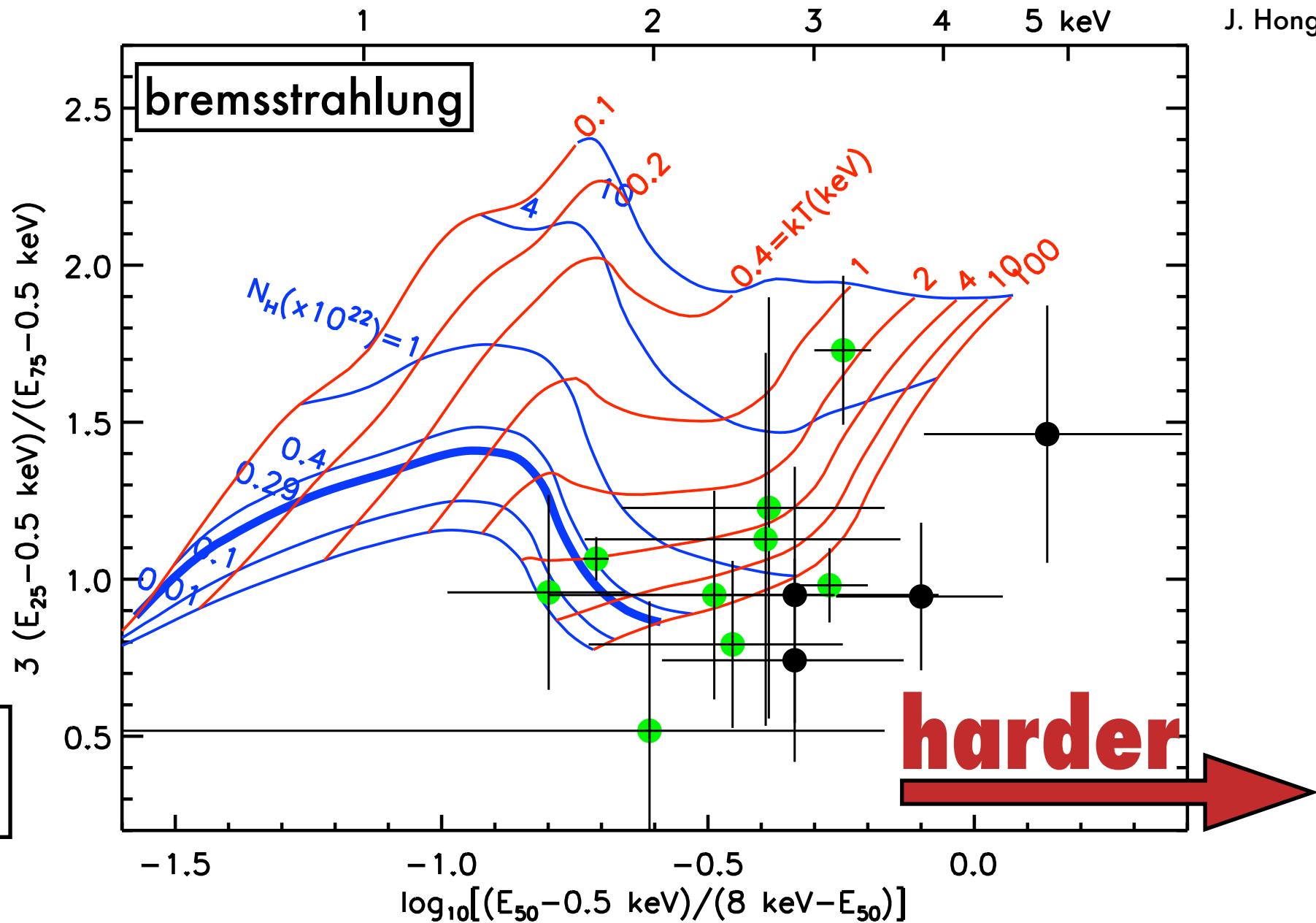


Optical/nIR properties:

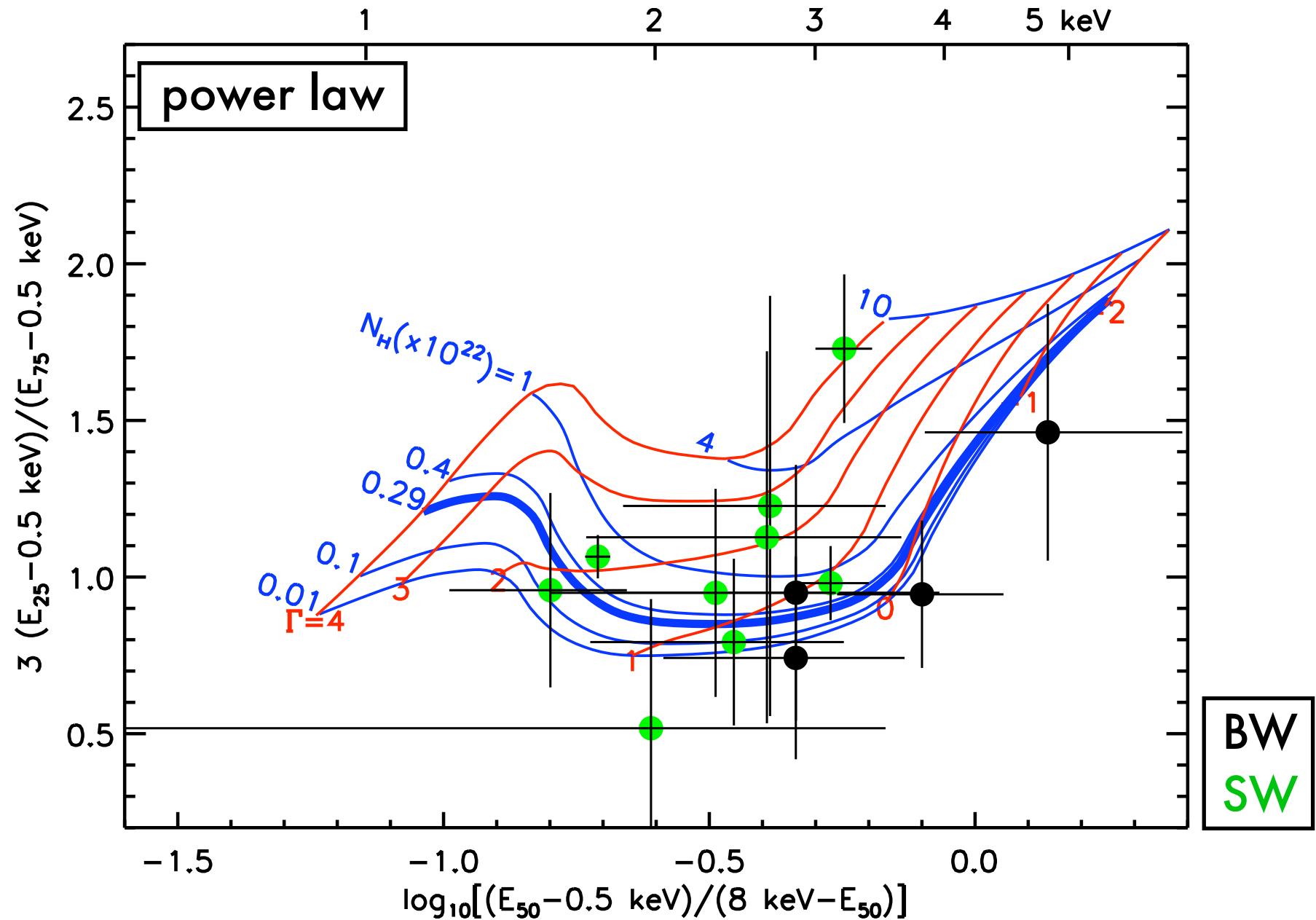
- ❖ (semi-)regular photometric variability
- ❖ mass loss $\sim 10^{-8} - 10^{-7}$ M_{\odot}/yr
- ❖ high outflow velocities (180 - 200 km/s)
- ❖ low-res spectra: like normal M-star,

X-ray properties: quantile analysis

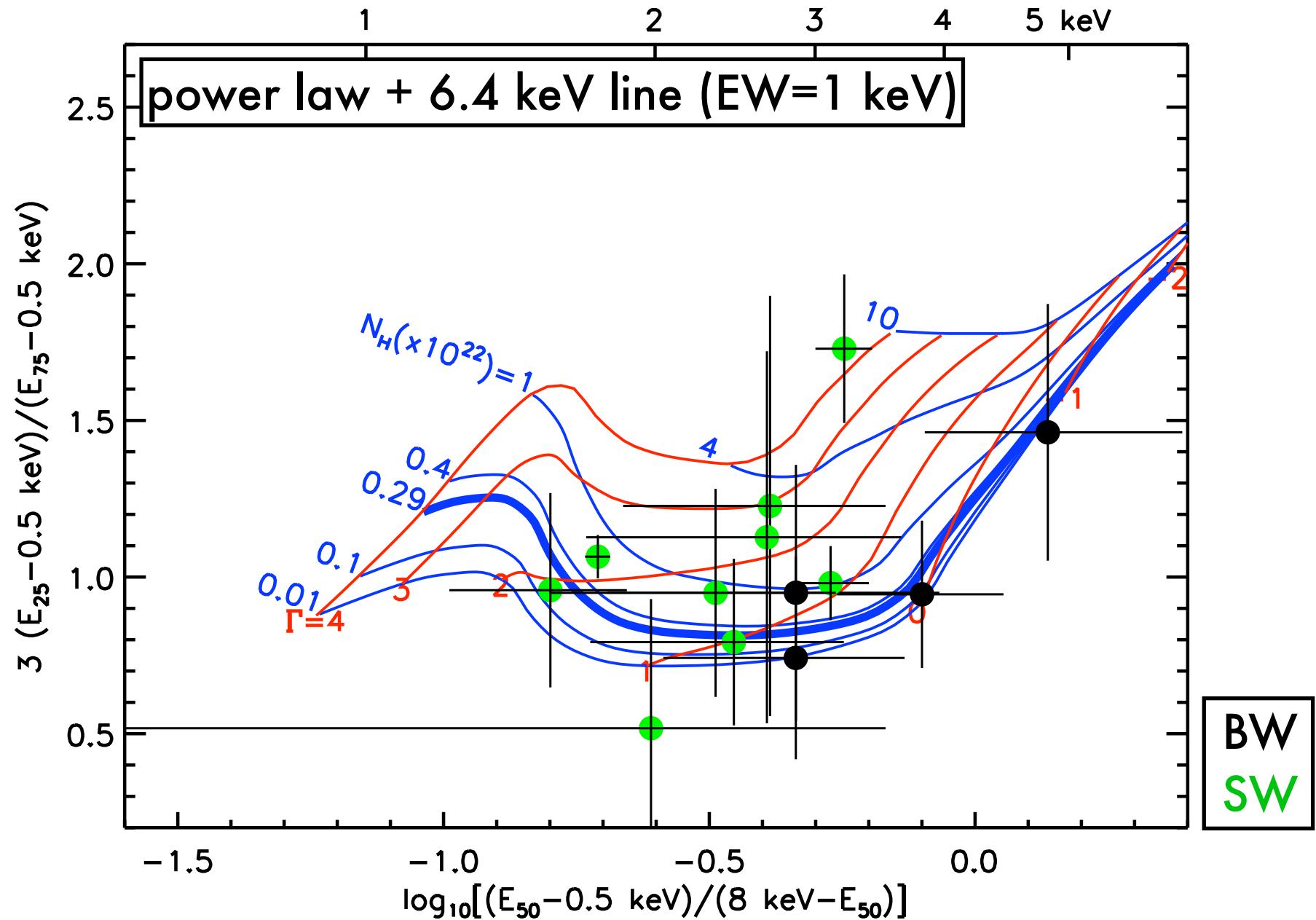
poster 181.04
J. Hong et al.



X-ray properties: quantile analysis



X-ray properties: quantile analysis



X-ray properties

- ❖ hard spectra:
 - 1.5 - 5.5 keV thermal bremsstrahlung, or
 - power law index between 1 and –1
 - Fe K lines ?
- ❖ $L(0.5\text{-}8 \text{ keV}) \approx (2\text{-}90) \times (\text{distance}/8.2\text{kpc})^2 10^{31} \text{ erg/s}$
- ❖ source $N_H \geq$ field N_H

Source classification

- ❖ single ? M-giants rarely detected in X-rays
- ❖ detached active binaries ? no RS CVn with M-giant
- ❖ wind-accreting symbiotic binaries ?
 - typical X-rays: very soft → a few keV
 - $L_x \sim 10^{29}$ - 10^{34} erg/s (white dwarf), $\sim 10^{37}$ erg/s (neutron star GX1+4)
 - wind: intrinsic absorption, Fe K lines (e.g. CH Cyg)

Comparison with symbiotics

✓ Lx range, ~few keV spectra

✓ intrinsic N_H, fast wind

BUT typically:

- not very hard in X-rays
- strong optical nebular emission-line spectra

⇒ different class ?

similar properties for RT Cru (INTEGRAL)
poster 32.07/J. Sokoloski et al.

Summary

- ❖ new candidate symbiotics
- ❖ hard x-rays, optical spectrum: not “typical”
- ❖ think about:
 - what produces hard X-rays ?
 - implications for binary population & evolution